

KONISHI ANOMALY Approach

To Gravitational F-Terms

based on 0304227

with

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Introduction

- Consider $U(N)$ $N=1$ SYM with a single adjoint chiral Multiplet Φ and an arbitrary super potential

$$V(\Phi) = \sum_{k=1}^{n+1} \frac{g_k}{k} \text{Tr}(\Phi^k) \quad g_2 > 0$$

- F-terms

- Feynman diagram contribution to F-terms obtained by integrating out the chiral multiplet reduce to planar graphs
- The diagrams reduce to the planar graphs of an associated matrix model with

$$V(M) = \frac{\hat{N}}{g_m} \sum_k \frac{g_k}{k} \text{Tr}(M^k) = \frac{\hat{N}}{g_m} V(M)$$

F-terms are given by

$$\int d^2\theta \left[N_i \frac{\partial F_i}{\partial S_i} + \frac{1}{2} \omega^i \omega_{\alpha\beta} \frac{\partial^2 F_i}{\partial S_i \partial S_j} \right]$$

$S_i = \omega^i \omega_{\alpha\beta}$ chiral superfield whose lowest component is the gaugino bilinear in the i 'th broken phase

$$U(N) \rightarrow \prod_{i=1}^n U(N_i)$$

$\omega_{\alpha\beta}^i$ $U(1)$ chiral gauge superfield of $U(N_i)$

• The proof of the above statements

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• Diagrammatic

• Anomaly Equations

• F-Terms in gravitational background

• The Feynman diagrams contributing to gravitational F-Terms reduce to genus 1 diagrams of the corresponding matrix model

F-Term $\int d^2\theta G^2 F_1(S)$

S: gaugino bilinear

$$G^2 = G^{\alpha\beta\gamma} G_{\alpha\beta\gamma}$$

$$G_{\alpha\beta\gamma} = \psi_{\alpha\beta\gamma} + \theta^\delta R_{\alpha\beta\gamma\delta} + \dots$$

N=1 Weyl Multiplet

• Diagrammatically: shown by Ooguri & Vafa

• We show genus 1 (γ_{N^2}) corrected resolvent equation for the matrix model agrees with the gravitational corrected anomaly equation in the gauge theory.

Outline of Talk

- Ingredients in the proof
 - Chiral Ring
 - Generalized Konishi anomaly
- Genus One Corrections
 - Connected Two point functions of gauge invariant chiral operators
- Comparison with Matrix Model
- Conclusions

Chiral Ring

- Consists of all operators annihilated by $\bar{D}_{\dot{\alpha}}$
- Relations in chiral ring are defined modulo $\bar{D}_{\dot{\alpha}}$ exact terms
- Restricting to the chiral ring is an efficient way of focussing on F-terms as $\bar{D}_{\dot{\alpha}}$ exact terms in the ring can be written as

$$\int d^4x d^2\theta d^2\bar{\theta} S(\theta, \bar{\theta})$$

- For $\mathcal{N}=1$ in absence of gravity chiral ring relations are

$$[W_{\alpha}, \bar{\Phi}] = 0 \quad \{W_{\alpha}, W_{\beta}\} = 0$$

- $$D^{\dot{\alpha}}(D_{\alpha\dot{\alpha}} W_{\beta}) = [D^{\dot{\alpha}}, D_{\alpha\dot{\alpha}}] W_{\beta} + \cancel{D_{\alpha\dot{\alpha}} D^{\dot{\alpha}} W_{\beta}}$$
$$= \{W_{\alpha}, W_{\beta}\} \quad (\text{Branchi Identity})$$

$$D_{\alpha\dot{\alpha}} = \frac{\partial}{\partial x^{\alpha\dot{\alpha}}} + [D_{\dot{\alpha}}(e^{-V} D_{\alpha} e^V), \cdot]$$

$$D_{\dot{\alpha}} e^{-V} D_{\alpha} e^V \sim \sigma_{\alpha\dot{\alpha}}^{\mu} A_{\mu} + \bar{\Theta}_{\dot{\alpha}} \lambda_{\alpha}$$

$$[D^{\dot{\alpha}}, D_{\alpha\dot{\alpha}}] = Ad(W_{\alpha})$$

$$\begin{array}{l} \overline{\hspace{1cm}} \\ W_{\alpha} = \lambda_{\alpha} + \Theta^{\epsilon} F_{\epsilon\alpha} \\ \overline{\hspace{1cm}} \\ + \dots \end{array}$$

- Using the ring relations all gauge invariant operators can be arranged as

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$$\text{Tr}(W^2 \varphi^k) \rightarrow R = -\frac{1}{32\pi^2} \text{Tr}\left(\frac{W^2}{z-\varphi}\right)$$

$$\text{Tr}(W_\alpha \varphi^k) \rightarrow \omega_\alpha = \frac{1}{4\pi} \text{Tr}\left(\frac{W_\alpha}{z-\varphi}\right)$$

$$\text{Tr}(\varphi^k) \rightarrow T = \text{Tr}\left(\frac{1}{z-\varphi}\right)$$

- In presence of gravity

$$[W_\alpha, \Phi] = 0 ; [W_\alpha, W_\beta] = 2G_{\alpha\beta\gamma} W^\gamma$$

again: $D^{\dot{\alpha}}(D_{\alpha\dot{\alpha}} W_\beta) = [D^{\dot{\alpha}}, D_{\alpha\dot{\alpha}}] W_\beta$

but $D_{\alpha\dot{\alpha}} W_\beta = \left(\frac{\partial}{\partial x^{\alpha\dot{\alpha}}} + \sigma_{\alpha\dot{\alpha}}^\mu \omega_{\mu ab} \sigma^{ab} + \sigma_{\alpha\dot{\alpha}}^\mu A_\mu \right) W_\beta$
(lowest components)

$$\sigma_{\alpha\dot{\alpha}}^\mu \Omega_{\mu ab} (\sigma^{ab})_{\beta\gamma} \sim \sigma_{\alpha\dot{\alpha}}^\mu \omega_{\mu ab} \sigma^{ab} + \bar{\Theta}_{\dot{\alpha}} \psi_{\alpha\beta\gamma}$$

gravitino
field strength

- Again can organize fields into R, ω_α, T

- Other relations in the Ring

$$G_{\alpha\beta\gamma} \omega^\gamma = 0$$

$$G^2 W_\alpha = 0$$

$$G^4 = 0$$

- Gravitational Corrections truncate at order G^2

Generalized Konishi Anomaly

• To obtain equations which completely determine the operators R, T, W_α we need Ward Identities. Provided by the Konishi Anomaly

• Consider the variation

$$\delta\phi = \epsilon\phi \quad \delta\bar{\phi} = -\epsilon\delta\phi$$

Conserved current

$$0 = \langle \bar{D}^2 (\bar{\phi} e^V \phi) \rangle = \frac{1}{32\pi^2} \text{Tr}_{Ad} W^\alpha W_\alpha + \text{Tr} \bar{\phi} \frac{\partial V}{\partial \phi}$$

Holomorphic contribution

Anomaly

classical variation

Consider \mathcal{Q}^2 with $\bar{\mathcal{Q}}^2$ component

$$\partial_\mu \text{Tr} (\bar{\psi} \bar{\sigma}^\mu \psi - \psi \sigma^\mu \bar{\psi}) = \frac{1}{32\pi^2} \text{Tr}_{Ad} (F \wedge F) + \frac{N^2}{32\pi^2} \cdot \frac{1}{24} R \wedge R$$

superfield

In presence of gravity

$$+ \frac{1}{32\pi^2} \text{Tr}_{Ad} W^\alpha W_\alpha + \frac{N^2}{32\pi^2} \cdot \frac{1}{3} (C^{\alpha\beta\gamma} \epsilon_{\alpha\beta\gamma})$$

for a generalized variation $\delta\phi_{ij} = \epsilon f_{ij}(\phi, W)$

one gets $0 = A_{ij;kl} \frac{\delta f_{ij}}{\delta \phi_{kl}} + \text{Tr}(fV')$

$$A_{ij;kl} = \frac{1}{32\pi^2} \left\{ (W^2)_{kj} \delta_{il} + (W^2)_{il} \delta_{kj} - 2(W^2)_{ij} (W^2)_{kl} \right. \\ \left. + \frac{1}{3} \delta_{kl} \delta_{ij} C^2 \right\}$$

Using The generalized Anomaly & the Equations in the Chiral Ring We get

$$\langle R^2(z) \rangle - \langle \text{Tr}(V'(\Phi) R) \rangle = 0$$

$$2 \langle R(z) W_\alpha(z) \rangle - \langle \text{Tr}(V'(\Phi) W_\alpha(z)) \rangle = 0$$

$$2 \langle R(z) T(z) \rangle - \langle \text{Tr} V'(\Phi) T(z) \rangle + \langle W^\alpha(z) W_\alpha(z) \rangle - \frac{G^2}{3} \langle T(z) T(z) \rangle = 0$$

Note: In flat space correlation fns of chiral gauge Invariant Operators factorize.

Here a priori eg. $\langle R^2 \rangle \neq \langle R \rangle \langle R \rangle$

strategy: To solve above equations

$$\rightarrow \text{write } \langle R^2 \rangle = \langle R \rangle^2 + \langle RR \rangle_c$$

$$\langle RR \rangle_c \propto G^2$$

$$\rightarrow \text{Expand } \langle R \rangle = \langle R_0 \rangle_0 + G^2 \langle R_1 \rangle$$

\rightarrow Substitute in above Eqns & match order by order

$$\text{eg } \langle R_0 \rangle^2 - V'(z) \langle R(z) \rangle - \frac{f(z)}{4} + G^2 [(2R_0 - V'(z)) \langle R_1 \rangle] + \langle RR \rangle_c = 0$$

$$\text{where } \langle \text{Tr} V'(\Phi) - V'(z) R \rangle = \frac{1}{4} f(z)$$

$f(z)$: polynomial of degree $n-1$

$\rightarrow f(z)$ is fixed by demanding $\frac{1}{2\pi i} \int_{C_i} R(z) dz = S_i$
 $i = 1, \dots, n$

Conditions For other Operators

$$\frac{1}{2\pi i} \int_{C_i} w_k(z) dz = W_{ki} \quad ; \quad \frac{1}{2\pi i} \int_{C_i} T(z) dz = N_i$$

Genus One Corrections

How to obtain equations for the connected correlators:

KONISHI ANOMALY

Consider for e.g. $\langle R(z) T(w) \rangle_c$

Variation $f_{ij} = R_{ij}(z) T(w)$

Jacobian

$$\frac{\delta f_{ij}}{\delta \phi_{kl}} = \frac{\delta R_{ij}(z) T(w)}{\delta \phi_{kl}} + R_{ij}(z) T_{kl}(w) T_{em}(w)$$

1st Term

with classical piece $\langle R^2 - \text{Tr}(V'(\Phi) R(z)) \cdot T(w) \rangle$
 $= \langle R^2 - \text{Tr} V'(\Phi) R(z) \rangle \langle T(w) \rangle$
 $+ 2 \langle R(z) \cdot \rangle \langle R(z) T(w) \rangle_c + \langle RRT \rangle_c$
 $- \langle \text{Tr} V'(\Phi) R(z) T(w) \rangle_c$

2nd Term

$$= \frac{1}{3} G^2 \langle \text{Tr}(R(z) T(w) T(w)) \rangle = -\frac{G^2}{3} \partial_w \left(\frac{R(z) - R(w)}{z-w} \right)$$

Net Eq

$O(G^4)$ ⑨

$$\left[2R(z) - I(z) \right] \langle R(z) T(w) \rangle_c + \langle R(z) T(z) T(w) \rangle_c - \frac{1}{3} G^2 \partial_w \left(\frac{R(z) - R(w)}{z-w} \right) = 0$$

where: $I(z) A(z) = \frac{1}{2\pi i} \int_{C_z} dy \frac{v'(y) A(y)}{y-z}$

NOTE: $\langle R(z) T(w) \rangle_c \neq 0 \propto G^2$

→ The Equations for the various Connected 2pt fns are consistent

e.g $M(z) \langle R(z) T(w) \rangle_c = \frac{1}{3} G^2 \partial_w \left(\frac{R(z) - R(w)}{z-w} \right)$

$M(z) = 2R(z) - I(z)$

also get

$$M(w) \langle R(z) T(w) \rangle_c = \frac{1}{3} G^2 \partial_z \left(\frac{R(z) - R(w)}{z-w} \right)$$

Non trivial condition

$$\frac{G^2}{3} M(w) \partial_w \left(\frac{R(z) - R(w)}{z-w} \right) = \frac{G^2}{3} M(z) \partial_z \left(\frac{R(z) - R(w)}{z-w} \right)$$

is satisfied.

→ Solutions to the 2-pt fns. fixed by physical Requirement

$$\frac{1}{2\pi i} \int_{C_1} dz R(z) = S_i ; \quad \frac{1}{2\pi i} \int_{C_1} dw T(w) = N_i ; \quad \frac{1}{2\pi i} \int_{C_1} dz w_a(z) = w_a$$

→ Contour Integrals of connected 2pt fns around

z & w branch cuts vanish

Soln is unique

let $\langle R(z) T(w) \rangle_c = G^2 H(z, w)$ $H(z, w)$ is symmetric

• Solutions for connected 2pt fns

$$\begin{aligned} \langle R(z) R(w) \rangle_c &= \langle \omega_\alpha(z) T(w) \rangle_c = \langle \omega_\alpha(z) R(w) \rangle_c \\ &= \langle \omega_{\beta\alpha}(z) \omega_{\beta\alpha}(w) \rangle_c = 0 \end{aligned}$$

$$\langle R(z) T(w) \rangle_c = G^2 H(z, w)$$

$$\langle \omega^\alpha(z) \omega_\alpha(w) \rangle_c = 10 G^2 H(z, w)$$

$$\langle T(z) T(w) \rangle_c = G^2 N_i \frac{\partial}{\partial s_i} H(z, w)$$

• Solutions to 1-pt fn at $O(G^2)$

$\langle R \rangle$; $\langle \omega_\alpha \rangle$ receive no corrections

Corrections to $T(z)$ consists of 2 Terms

$$T_{(0)}^{(1)}(z) = -\frac{1}{6} N_i \frac{\partial}{\partial s_i} T^{(0)}(z) \quad (O(N^2) \text{ piece})$$

$$T_1^{(1)}(z) = \frac{-12}{(2R^0 - V')} [H(z, z) + C(z)]$$

$C(z)$: polynomial of degree $n-2$ uniquely determined by requirement $\int_{C_i} T_1^{(1)}(z) dz = 0$

• $O(N^2)$ genus (0) term can be absorbed in $T^{(0)}$ by field redefinition $s_i \rightarrow s_i + \frac{G^2}{6} N_i$

Comparison with the Matrix Model

Consider a hermitian matrix model

$$S = \frac{\hat{N}}{g_m} \sum_k \frac{g_k}{k} \text{Tr}(M^k)$$

The resolvent $\Omega = \frac{g_m}{\hat{N}} \text{Tr} \left(\frac{1}{z-M} \right)$ satisfies

$$\langle \Omega^2 \rangle - \mathbb{I}(z) \langle \Omega(z) \rangle = 0$$

$$\rightarrow \langle \Omega \rangle^2 + \langle \Omega(z) \Omega(z) \rangle_c - \mathbb{I}(z) \langle \Omega(z) \rangle = 0$$

$\sim \frac{1}{N^2}$

Identical to gauge theory Resolvent $R(z)$

To compute $\frac{1}{N^2}$ corrections we need $\langle \Omega(z) \Omega(z) \rangle_c$

→ Similar procedure

Consider the variation

$$\delta M_{ij} = \Omega_{ij}(z) \Omega(w)$$

obtain

$$\left(2 \Omega(z) - \mathbb{I}(z) \right) \langle \Omega(z) \Omega(w) \rangle_c + \langle \cancel{\Omega(z)} \Omega(z) \Omega(w) \rangle_c \sim \frac{1}{N^2}$$

$$- \left(\frac{g_m}{N} \right)^2 \partial_w \left(\frac{\Omega(z) - \Omega(w)}{z-w} \right) = 0$$

Identical to eq for $3 \langle R(z) T(w) \rangle$

Demanding $\int_{C_i} \langle \Omega(z) \Omega(w) \rangle dz = 0 = \int_{C_i} \langle \Omega(z) \Omega(w) \rangle dw$

Soln is unique .8

$$\langle \Omega(z) \Omega(w) \rangle_c = \left(\frac{g_m}{N} \right)^2 3 H(z, w)$$

Similarly can show that

$$\Omega^{(1)}(z) = \frac{1}{4} T^{(1)}(z)$$

Conclusions :

- We have used **Konishi Anomaly**
Modified Chiral Ring to show gravitational
 F-terms are captured by genus one
 diagram of the matrix model.

crucial ingredient : **lack of factorization**

- Including the graviphoton field strength $F_{\alpha\beta}$
 in the ring.

$$\{W_\alpha, W_\beta\} = F_{\alpha\beta} + G_{\alpha\beta\gamma} W^\gamma$$

One can show that the gauge theory resolvable R
 captures the all genus expansion of the
 matrix model

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