

# **The Trailing String in Confining Holographic Theories**

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**Crete Center for Theoretical Physics, 22-10-2013**

**Work with E. Kiritsis, L. Mazzanti, to appear soon**

# Introduction: AdS/CFT

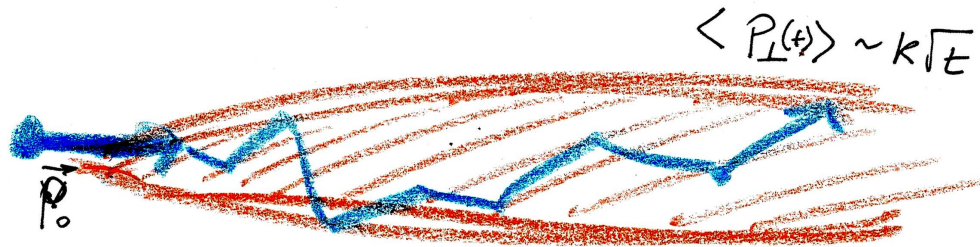
Since 15 years, the AdS/CFT correspondence or gauge/gravity duality, offered us another way to think about QFTs

- The QFT degrees of freedom and dynamics can be recast in the language of a theory of gravity in a higher dimensional, curved space-time.
- The higher dimensional (bulk) theory becomes simple e.g. classical GR in the regime where the number of degrees of the QFT becomes large, and the couplings become strong.
- The deconfined, high-temperature phase of gauge theories is mapped to a black hole solution.

Strongly coupled quantum field theories may be described using the tools of classical General Relativity

# Introduction: Heavy quarks in QGP

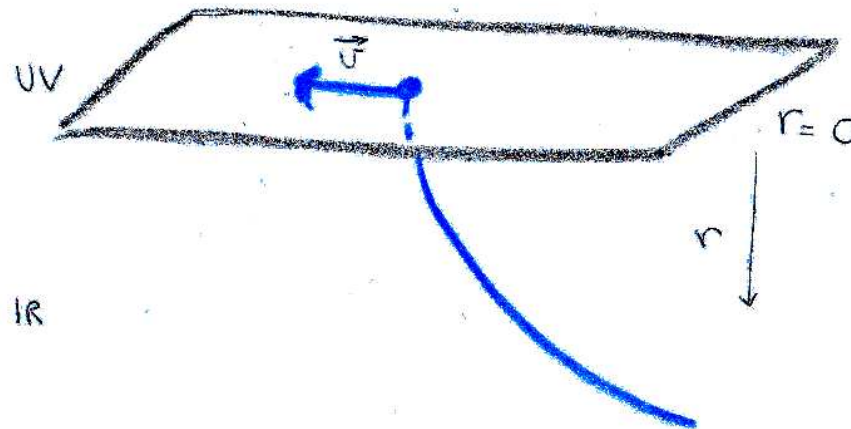
- in a heavy-ion collision experiment, a collective state (quark-gluon plasma) is formed that undergoes fast thermalization and can be described by hydrodynamics.
- A **heavy quark** can be created out of equilibrium in the QGP produced by a heavy ion collision. It then undergoes a diffusion process governed by the interactions with the medium.



The brownian-like dynamics causes viscous **energy loss** and a **spread in transverse momentum**

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Dual dual to a **trailing string** with the quark as its endpoint.

# Motivation

I will discuss the **trailing string solution in  $T = 0$  vacuum geometries**, in particular those dual to a **confining theory**.

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- From a practical point of view, in order to correctly obtain the dynamics of the probe in the deconfined medium, one needs a **subtraction procedure** to make basic quantities (Boundary retarded correlators) well defined. **The natural way to operate this subtraction is through the vacuum correlator**. Whence the need of the vacuum trailing string solution.

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- The vacuum trailing string in confining theories, and its fluctuations, have an interesting and non-trivial structure
  - presence of a **confining horizon**
  - long time **dissipation effects**

# Outline

- Confinement in AdS/CFT
- The trailing string picture of a probe quark, and connection to Brownian motion
- Review of trailing string in a black hole solution
- Static trailing string in a confining background
- Dragged confining trailing string
- Conclusion



# AdS/CFT and Confinement

Confinement in AdS/CFT is decided via the dual of the Wilson loop test: In confining gauge theories, the Wilson Loop operator

$$W(\gamma) = \mathcal{P} \exp i \oint_{\gamma} A$$

exhibits an **Area Law**:  $W \sim \exp \sigma_c \text{Area}$

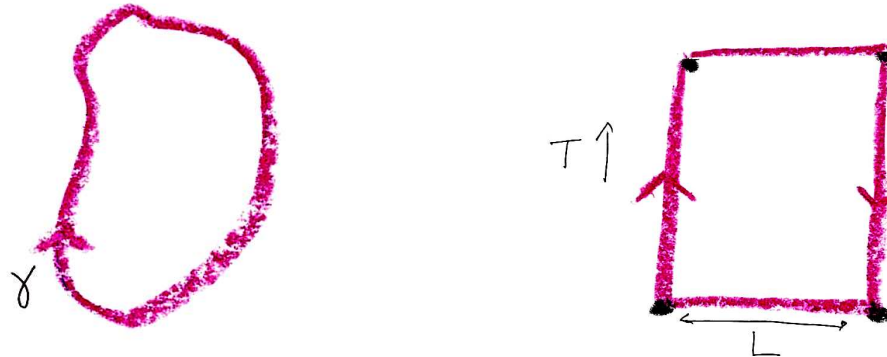


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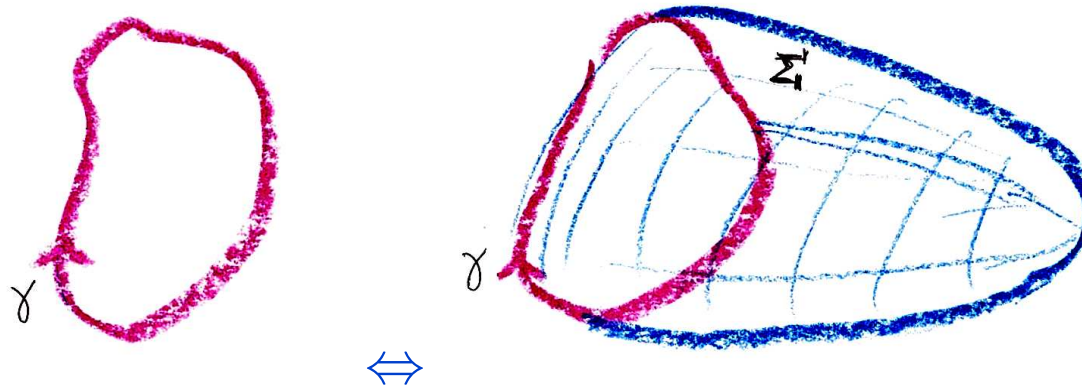


Area law implies a **linear potential** between two static quarks,

$$S_{\gamma} = TV(L) \sim \sigma_c TL \Rightarrow V(L) = \sigma_c L$$

# AdS/CFT and Confinement

The holographic dual of the Wilson Loop is the action of a string attaching to the contour on the boundary, and closing into the interior.



$$S_\gamma = \frac{1}{2\pi\ell_s^2} \text{Area}_\Sigma$$

# AdS/CFT and Confinement

Confinement if:

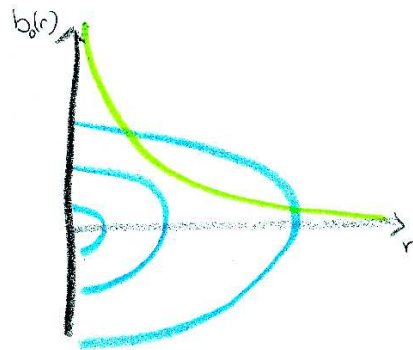
- The higher-dimensional space ends regularly (or at a hard wall) at some coordinate  $r_m$
- The space is non-compact in the IR but the metric functions have a non-zero minimum at some point  $r_m$

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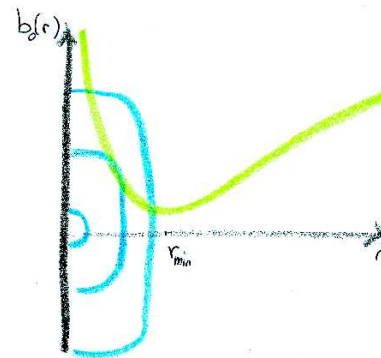
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$$ds^2 = b(r) [dr^2 + dx_\mu^2]$$

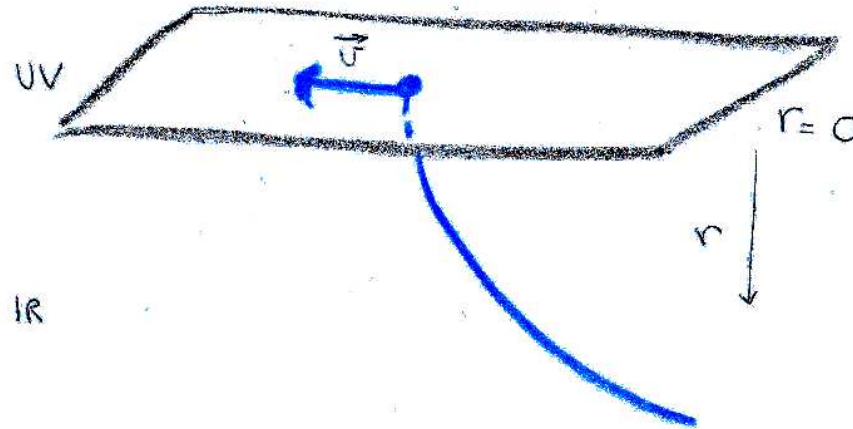


non-confining,  $\sigma_c = 0$



confining,  $\sigma_c = b^2(r_m)$

# The Trailing String



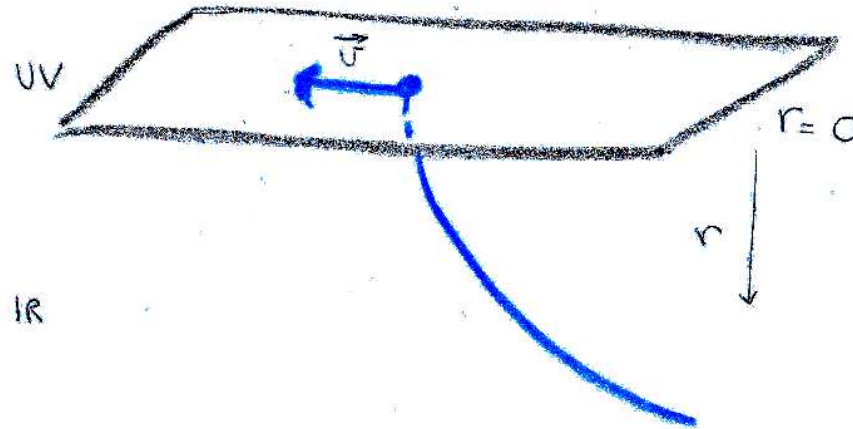
Probe quark on the boundary a 5D asymptotically  $AdS$  spacetime



Classical string attached at the boundary and extending in the interior.

(Gubser '06)

# The Trailing String



The string profile is found by extremizing the surface spanned by the string

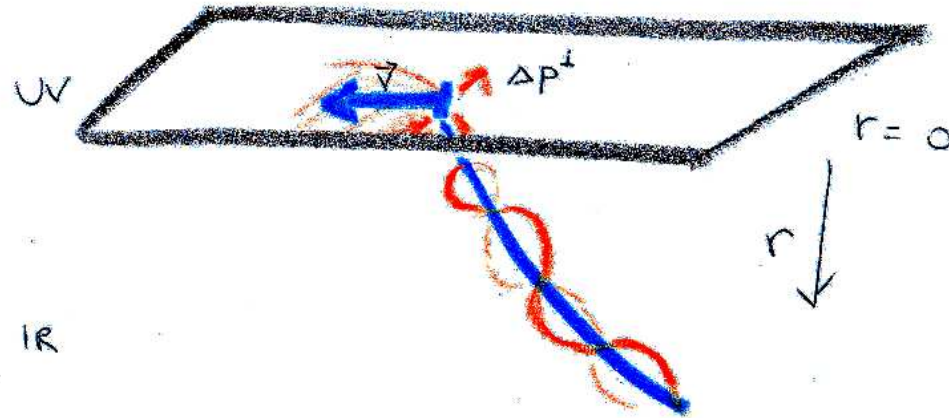
$$S = \frac{1}{2\pi\ell_s^2} \int dt dr \sqrt{-\det g_{ind}},$$

with respect to the embedding coordinates:  $\vec{X}(t, r) = \vec{v}t + \vec{\xi}(r)$ .

The string exerts a drag force which causes the quark to lose energy:

dual description of in-medium energy loss

# The Trailing String Fluctuations



Add small fluctuations along the string:

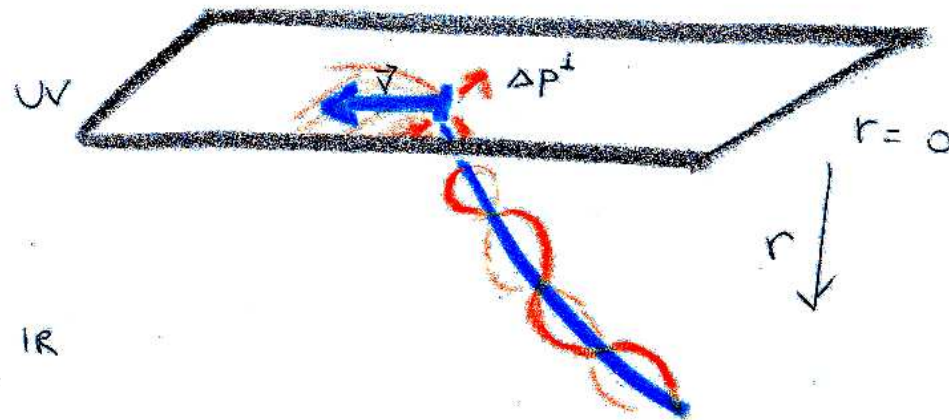
$$X(t, r) = \xi(r) + \delta X(t, r)$$

they induce a Brownian-like dynamics for the boundary quark, governed by a Langevin equation, and leading to a spread in momentum. (Gubser '05, De Boer *et al* 06, Herzog *et al* 06, Son and Teaney 09) .

**Dual description of transverse momentum broadening.**



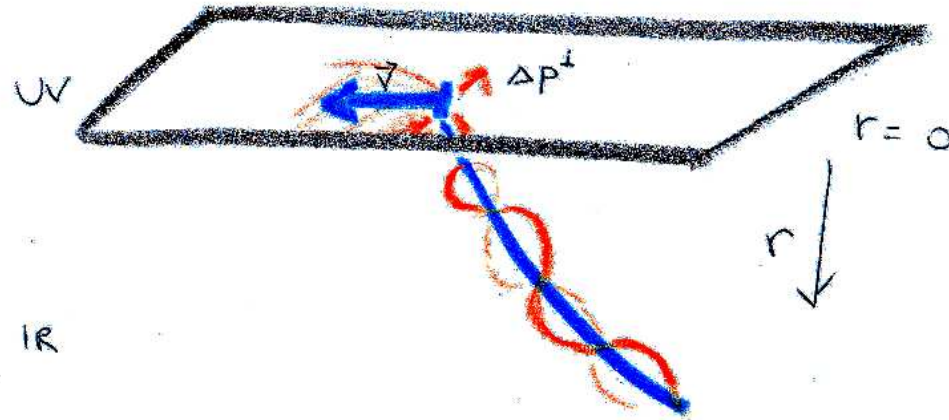
# The Trailing String Fluctuations



$$M\delta\ddot{X}(t) + \int dt' G_R(t-t')\delta X(t') = \zeta(t), \quad \langle \zeta(t)\zeta(t') \rangle = G_s(t-t')$$

Generalized Langevin equation.

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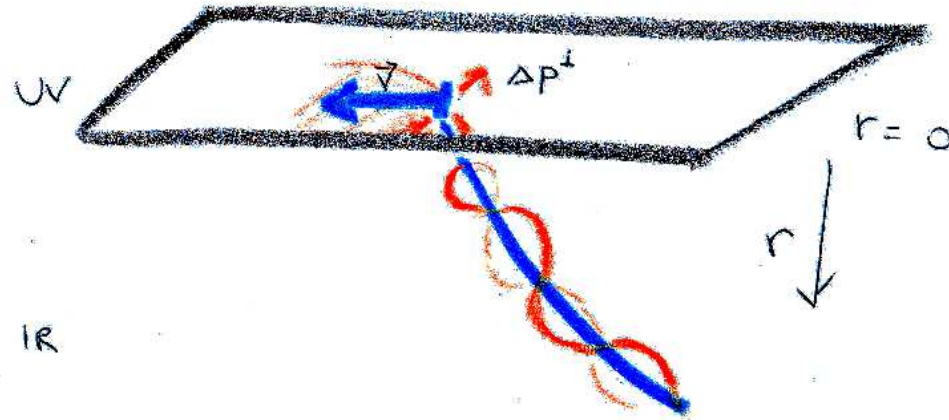
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**Generalized Langevin equation.** Two force terms:

- a **classical** force with retardation effects;
- a **stochastic** force with a Gaussian distribution.

Both terms arise from the same underlying physics.

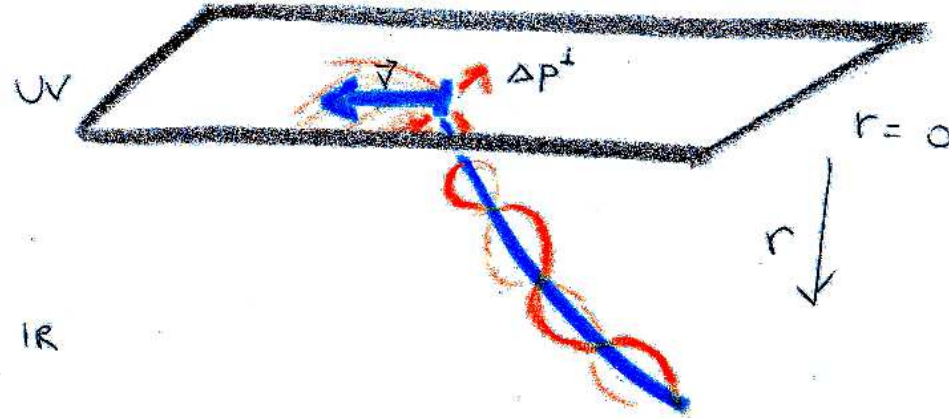
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- $G_R(t)$  is the **retarded** boundary correlator associated to the fluctuations  $\delta X(t, r)$  around the classical trailing string.
- $G_s(t)$  is the associated the **symmetric** correlator, obtained from  $G_R(t)$  via a Fluctuation-Dissipation relation, characteristic of the ensemble.

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long-time limit:

$$M\delta\ddot{X}(t) + \eta\dot{X}(t) = \zeta(t), \quad \langle \zeta(t)\zeta(t') \rangle = \kappa\delta(t-t')$$

$$\eta = \lim_{\omega \rightarrow 0} \frac{\text{Im } G_R(\omega)}{\omega}, \quad \kappa = \lim_{\omega \rightarrow 0} G_s(\omega)$$

# Trailing string in 5D black hole

Consider a generic asymptotically *AdS* 5D black hole:

$$ds^2 = b^2(r) \left[ \frac{dr^2}{f(r)} - f(r)dt^2 + dx^i dx_i \right]$$

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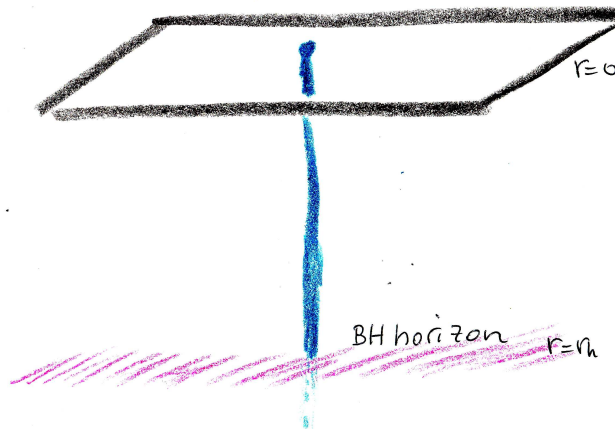
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- Dual to a **non-conformal** gauge theory in thermal equilibrium **at** a temperature  $T_h$ , in a **deconfined phase**.



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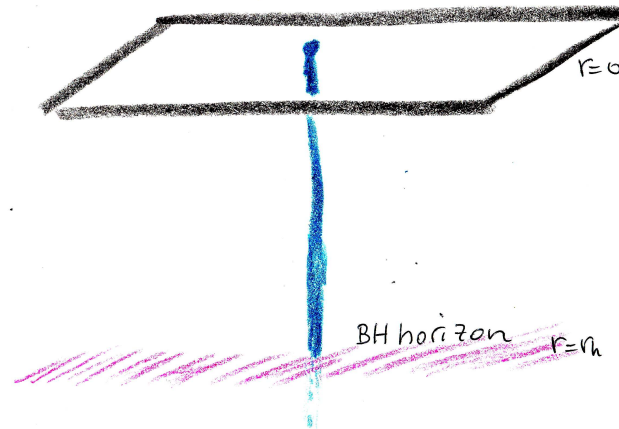
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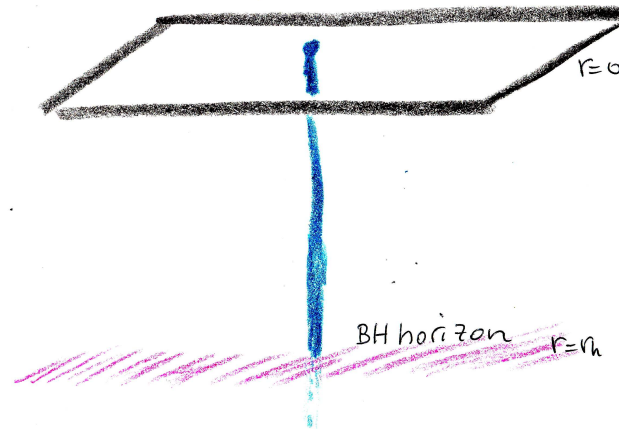
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The induced 2D worldsheet metric is a **2D black hole** with horizon  $r_h$  and temperature  $T_h$ .

$$ds_{ind}^2 = b^2(r) \left[ -f(r)dt^2 + f^{-1}(r)dr^2 \right],$$

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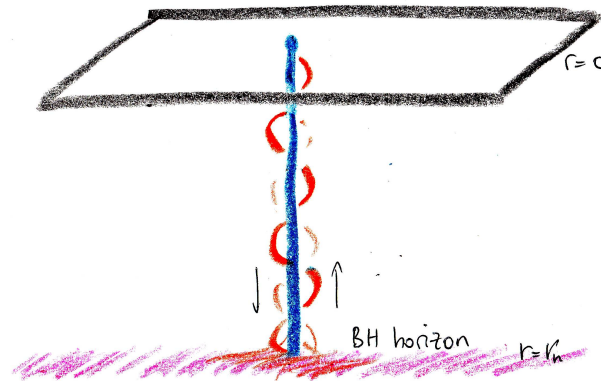
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$$ds_{ind}^2 = b^2(r) [-f(r)dt^2 + f^{-1}(r)dr^2], \quad f(r) \simeq 4\pi T_h(r_h - r), \quad r \simeq r_h$$

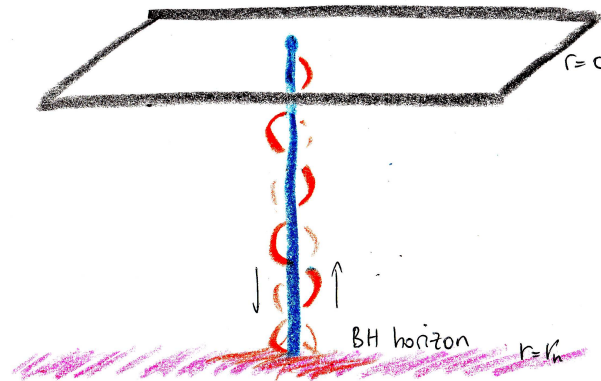
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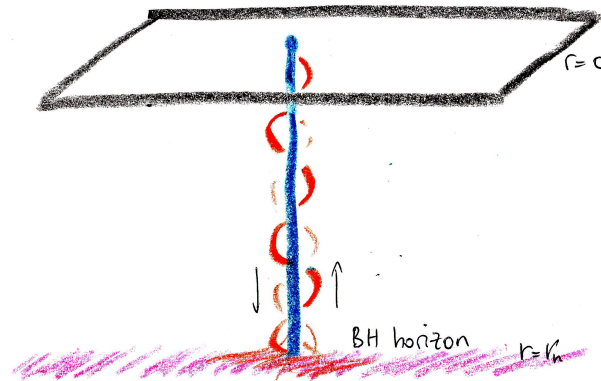


The fluctuation equation close to the horizon is

$$\delta X'' - \frac{1}{(r_h - r)} \delta X' + \frac{\hat{\omega}^2}{(r_h - r)^2} \delta X = 0, \quad \hat{\omega} \equiv \frac{\omega}{4\pi T_h}$$

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The solutions have infalling/outgoing behavior near  $r_h$ ,

$$\delta X(\omega, r) \simeq (r_h - r)^{\pm i\hat{\omega}}$$

# Correlators

The retarded correlator is found by the Policastro-Son-Starinets prescription

$$G_R(\omega) = [\mathcal{G}(r) \delta X'_R(\omega, r)]_{r \rightarrow 0}, \quad \delta X_R(\omega, r) \rightarrow \begin{cases} 1 & r \rightarrow 0 \\ (r - r_h)^{-i\hat{\omega}} & r \rightarrow r_h \end{cases}$$

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At long times, the dynamics is encoded by the **Langevin coefficient**

$$\kappa = \lim_{\omega \rightarrow 0} G_s(\omega) = 2T_h \eta$$

# Green's functions: High frequency limit

The large  $\omega$  limit obtained via WKB approximation

Gursoy, Mazzanti, Kiritsis, FN 1006.3261:

$$\text{Im } G_R(\omega) \simeq \omega^3 h \left( \frac{\sqrt{2}}{\gamma\omega} \right) \quad b(r) \sim \frac{\ell}{r} h(r)$$

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This diverges too fast for  $G_R$  to be physical:

- Dispersion relations that allow to write

$$G_R = \int \frac{\text{Im } G_R(\omega')}{\omega' - \omega - i\epsilon}$$

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- Remark: the leading behavior is **temperature-independent**.

# Dressed spectral density

UV-safe spectral densities can be defined: **Subtract the correlator obtained from the vacuum background.** [Mazzanti, Kiritsis, FN, 1111.1008:](#)

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- This prescription can be obtained with a change of variable in the quark path integral.

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Two qualitatively very different cases:

- the vacuum could be confining (as in QCD)
- or non-confining (as in  $\mathcal{N} = 4$  SYM).

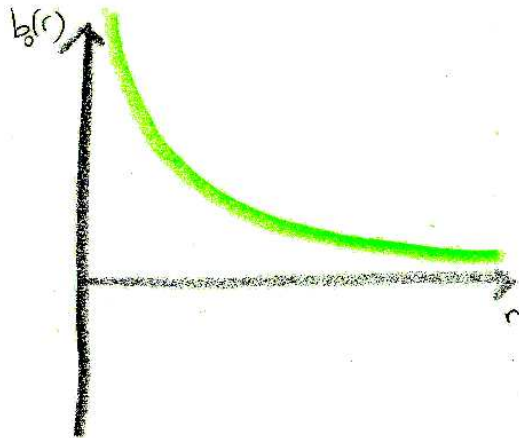
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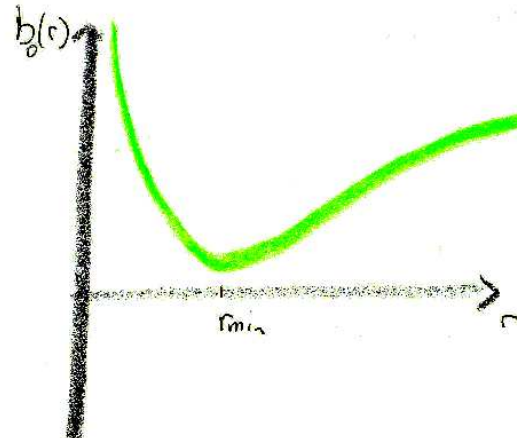
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Confinement is essentially equivalent to the presence of a minimum of the **bulk scale factor**  $b(r)$  (cfr. J .Sonnenschein's talk)



non-confining

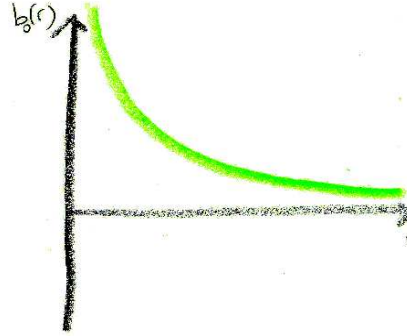


confining  $\sigma_c = b^2(r_m)$

# Non-confining case

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the string profile satisfies:

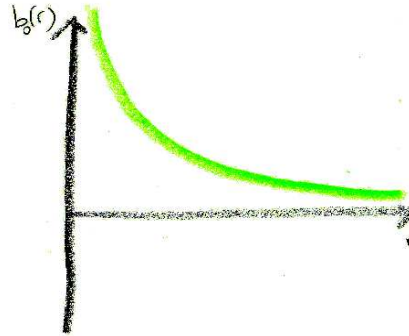
$$\xi'(r) = \frac{C}{\sqrt{b^4(r) - C^2}}$$



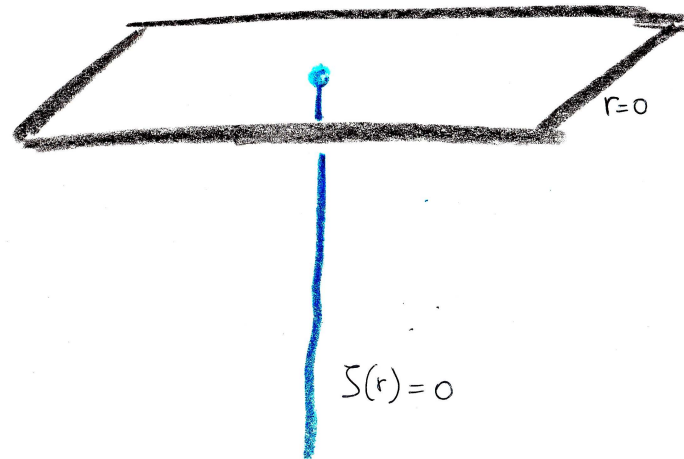
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As  $b \rightarrow 0$ , regularity requires  $C = 0$ : the embedding is trivial,  $\xi = 0$

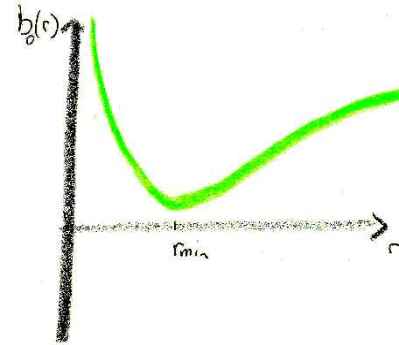


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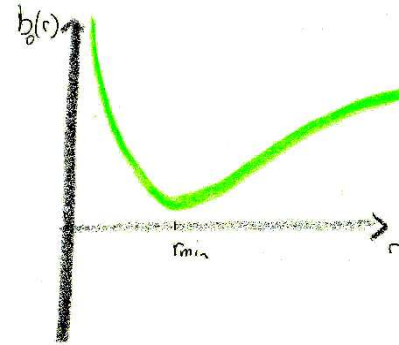
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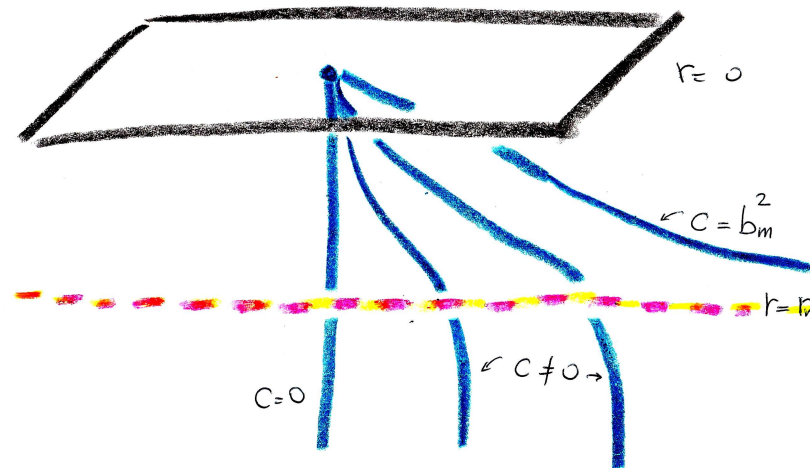
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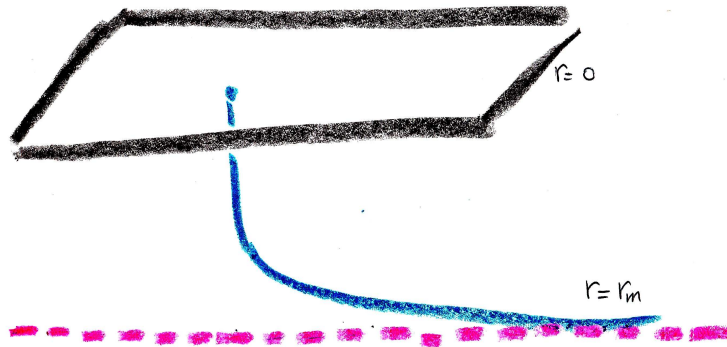
Now the minimum of  $b(r)$  is non-zero: the constant  $C$  is not fixed. One-parameter family of solutions with  $0 \leq C \leq b^2(r_m)$





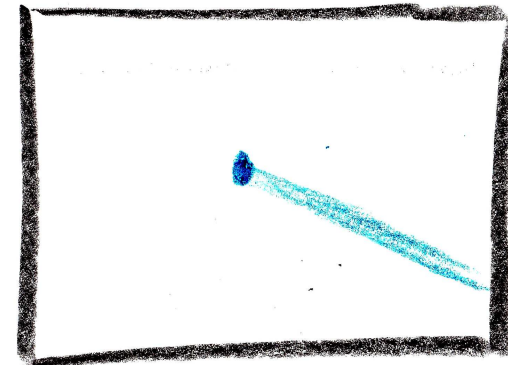
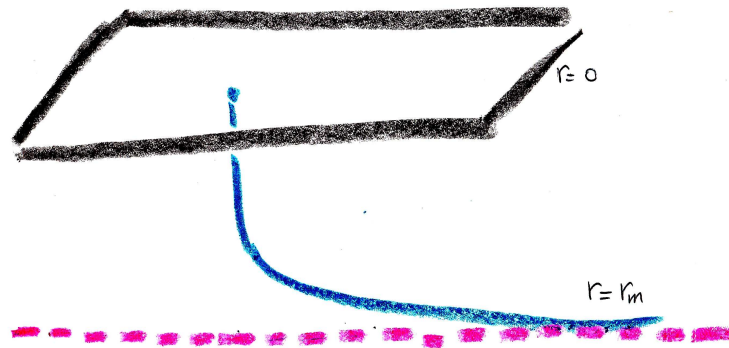
# Confining Trailing String

The extremal  $C = b_m^2$  string is the one with lowest action. It does not extend beyond the *confining horizon*  $r = r_m$ , and it extends to infinity along one of the spatial directions.



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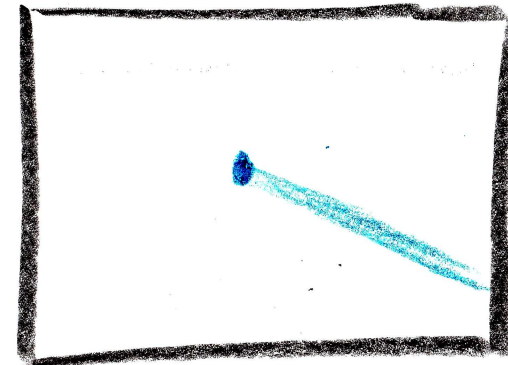
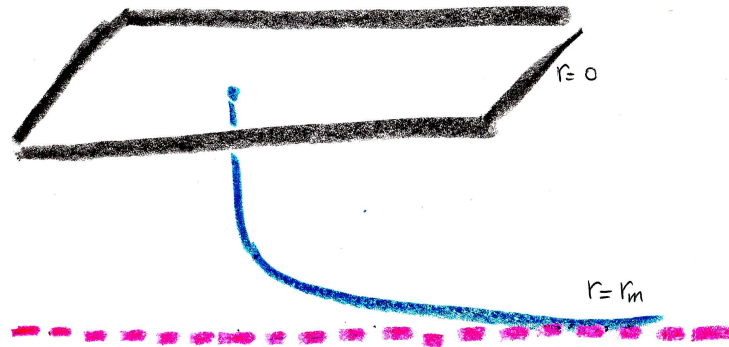
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Asymptotically it looks like a straight string with fixed tension  $b_m^2$  i.e. **the confining string tension of the dual theory**: it is the QCD flux-tube.

# Confining Trailing String

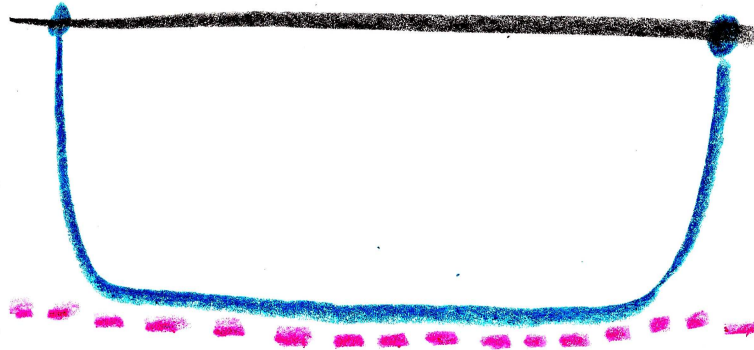
The extremal  $C = b_m^2$  string is the one with lowest action. It does not extend beyond the *confining horizon*  $r = r_m$ , and it extends to infinity along one of the spatial directions.



The string breaks rotational invariance. It can be recovered summing over different directions.

# Physical picture: the Shadow Quark

This has a simple physical interpretation:



look at the trailing string as **half** of the confining string connecting two quarks, one of which is observed, the other (**shadow quark**) infinitely far.

# Confining string geometry

Worldsheet induced metric:

$$ds^2 = b^2(r) \left[ -dt^2 + \frac{b^4}{R^2} dr^2 \right], \quad R = \sqrt{b^4 - b_m^4}$$

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- close to the confining horizon  $r_m$  it reduces to

$$ds^2 \sim b_m^2 \left[ -dt^2 + (4\pi T_m)^2 \frac{dr^2}{(r_m - r)^2} \right] \quad T_m = (4\pi)^{-1} \sqrt{b_m''/b_m}$$

# Confining horizon geometry

Metric close to the confining horizon  $r_m$ :

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$$z \sim -4\pi T_m \log(r_m - r) \rightarrow +\infty.$$

The full metric is conformally flat, and interpolates between  $AdS_2$  in the UV and flat Minkowski in the IR.

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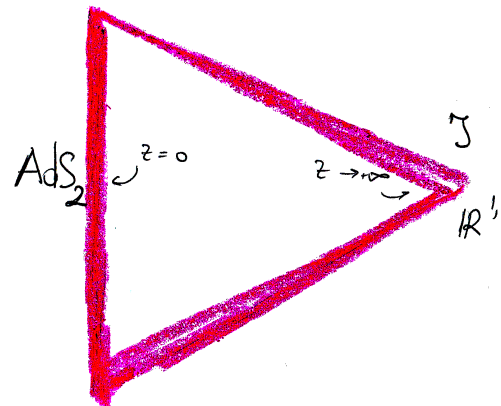
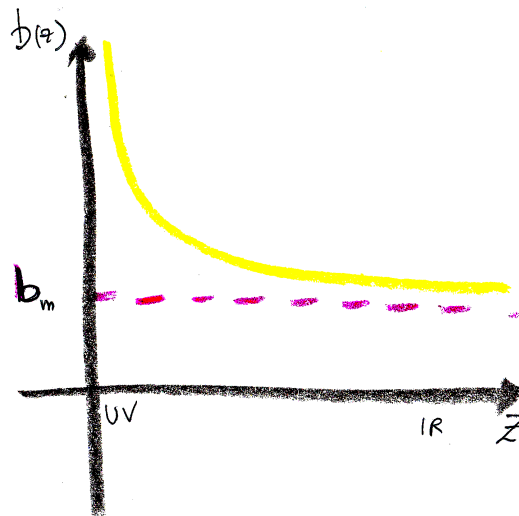
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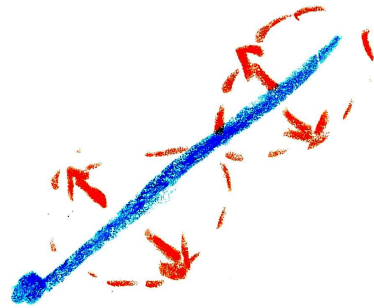
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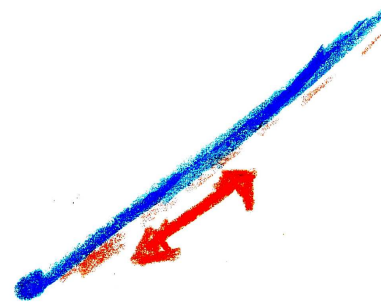


# Transverse vs. longitudinal fluctuations

We can distinguish between fluctuations longitudinal and transverse to the boundary direction of the string:



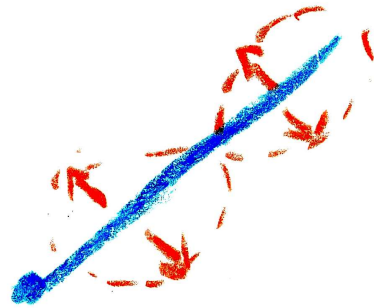
Transverse



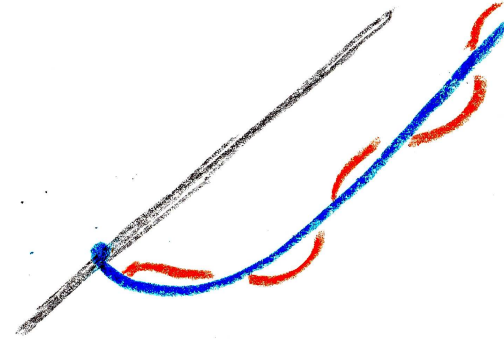
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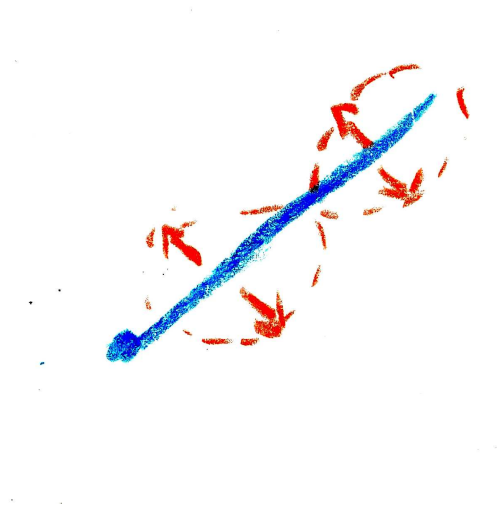
Transverse



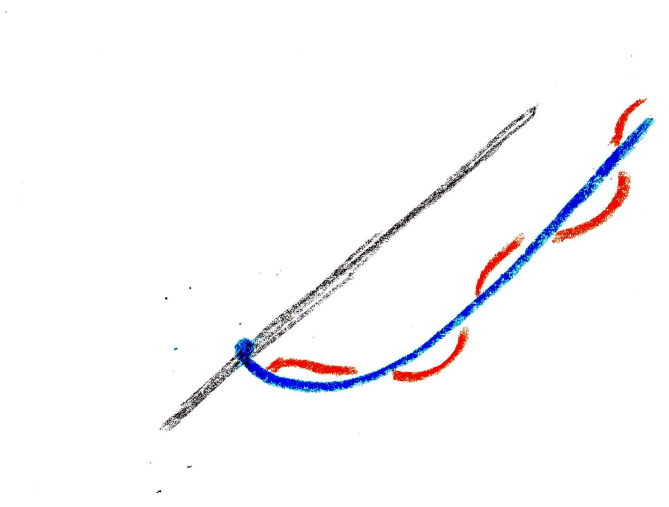
Longitudinal

# Transverse vs. longitudinal fluctuations

We can distinguish between fluctuations longitudinal and transverse to the boundary direction of the string:



Transverse



Longitudinal

We require **no information** come from the quark at infinity.

$\Rightarrow$  we impose **infalling** boundary conditions as  $r \rightarrow r_m$ .

# Transverse vs. longitudinal fluctuations

The  $\perp$  and  $\parallel$  fluctuations behave differently:

$$\partial_r \left[ R \partial_r \left( \delta X^\perp \right) \right] + \frac{\omega^2 b^4}{R} \delta X^\perp = 0,$$

$$\partial_r \left[ \frac{R^3}{b^4} \partial_r \left( \delta X^\parallel \right) \right] + \omega^2 R \delta X^\parallel = 0$$

$$R(r) = \sqrt{b^4(r) - b_m^2}$$

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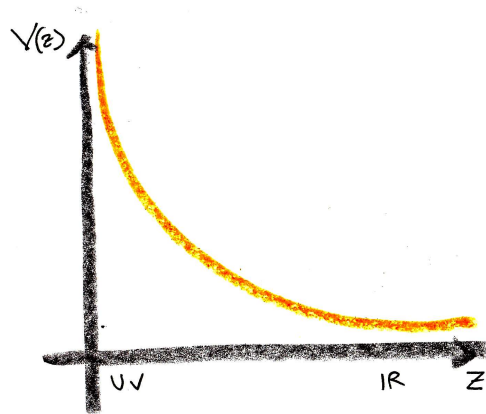
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$$R(r) = \sqrt{b^4(r) - b_m^2} \sim b_m^2 4\pi T_m (r_m - r), \quad r \simeq r_m.$$

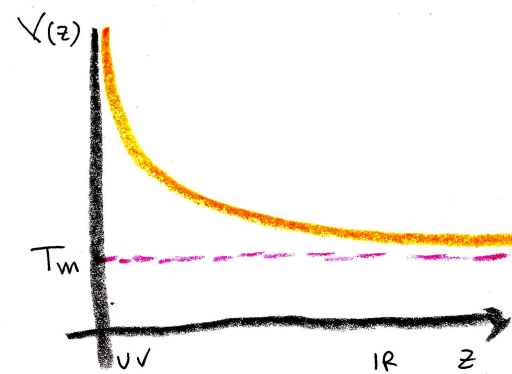
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One can transform the equation into the form of a Schrödinger problem:



Transverse



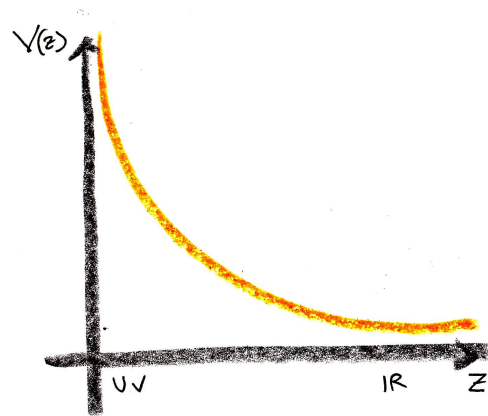
Longitudinal



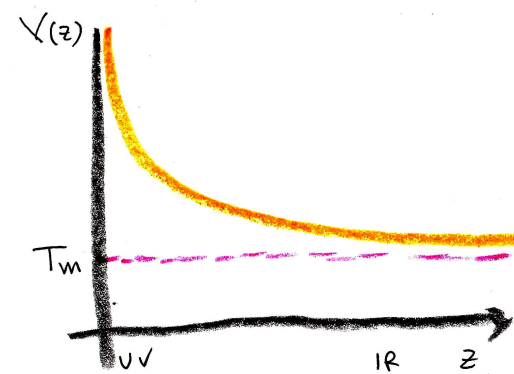
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Transverse



Longitudinal

The transverse modes have a continuous spectrum starting at  $\omega = 0$ , the longitudinal modes are **gaped** and start at  $\omega = 4\pi T_m$

# An Effective Temperature?

The equation for **transverse fluctuations** close to  $r_m$  is:

$$\delta X'' - \frac{1}{(r_m - r)} \delta X' + \frac{\hat{\omega}^2}{(r_m - r)^2} \delta X = 0, \quad \hat{\omega} \equiv \frac{\omega}{4\pi T_m}$$

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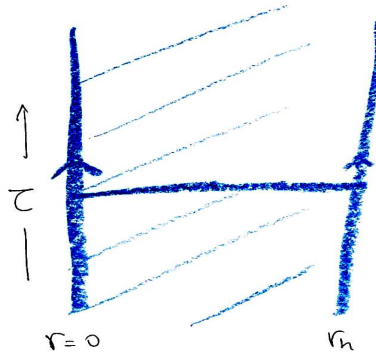
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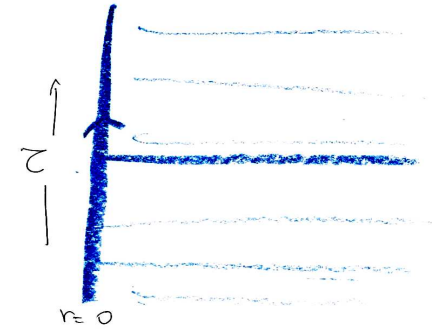
Are we seeing the emergence of an *effective temperature* set by the *confinement scale* ?

# An Effective Temperature?

BH: finite length

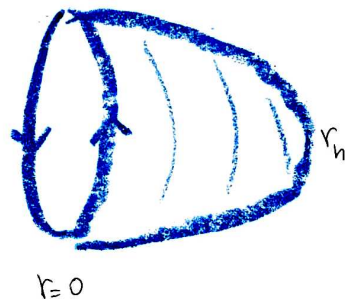
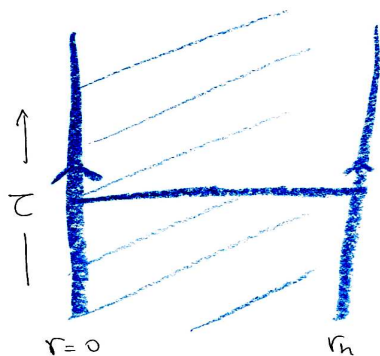


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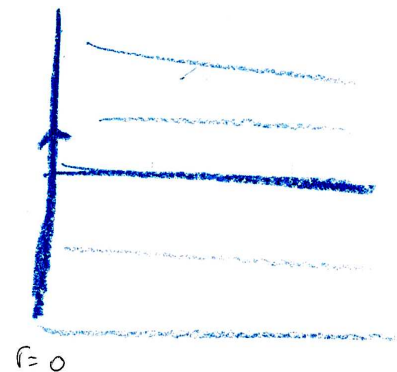
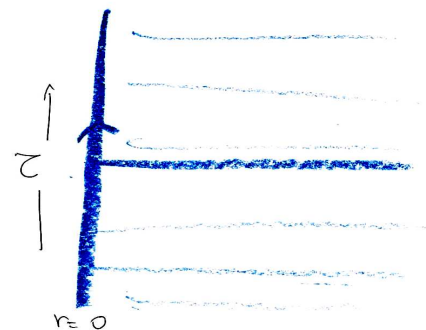
BH: finite length



Periodic euclidean time

$$T = T_h$$

Confining :  $\infty$  length



Non-compact Euclidean time

$$T=0$$

# Dissipation at zero temperature

The **transverse** boundary correlator at small frequency behaves as:

$$G_R^T(\omega) \simeq i b_m^2 \omega + O(\omega^2), \quad b_m^2 = \sigma_c \quad \textit{the confining string tension}$$

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The confining vacuum is dissipative for the fluctuations of a single **quark** and the dissipation time scale is again set by the confinement scale.



# Dissipation at zero temperature

Because of the gap in the longitudinal mode, the imaginary part of the **longitudinal** boundary correlator vanishes identically at small frequency:

$$G_R^L(\omega) = 0 \quad \omega < 4\pi T_m$$

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$[ImG_R^L/\omega]_{\omega \rightarrow 0} = 0$  and there is no viscous friction for longitudinal modes.

# Averaging over angles

The string solution breaks spontaneously spatial rotations  $SO(3) \rightarrow SO(2)$ . For a given solution, the correlator will be **anisotropic**:

$$G^{ij} = G^L(\omega)n^in^j + G^T(\omega)(\delta^{ij} - n^in^j) \quad \vec{n} = \vec{n}(\theta, \varphi)$$

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To quadratic order in the fluctuations:

$$Z = \int \frac{d\Omega}{4\pi} \exp i \int d\omega \delta X^i(\omega) G_{ij}(\omega, \Omega) \delta X^j(-\omega)$$

$$\simeq \exp i \int \frac{d\Omega}{4\pi} \int d\omega \delta X^i(\omega) G_{ij}(\omega, \Omega) \delta X^j(-\omega)$$

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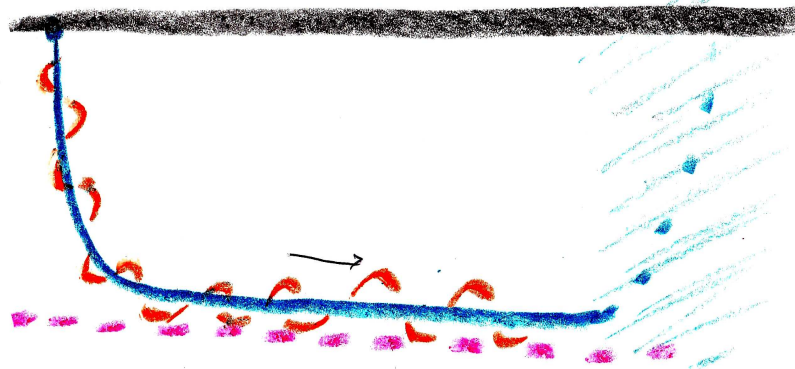
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# Physical picture: the Shadow Quark

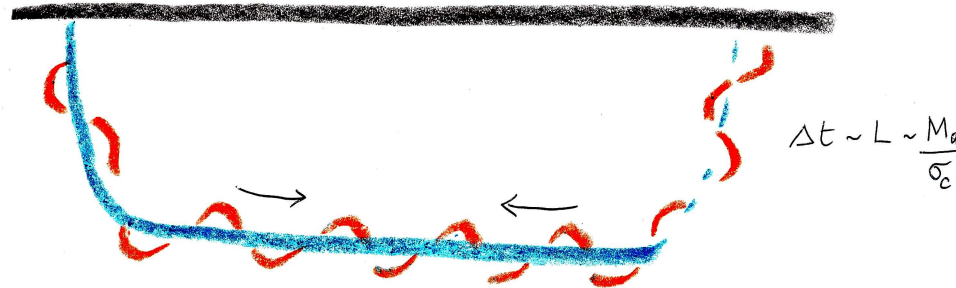
The friction coefficient arises because of infalling boundary condition.



All calculation done on a single (observed) quark should be done by assuming that **no information** is available or comes from the shadow quark. E.g. the infalling wave condition at  $z \rightarrow \infty$

# Physical picture: the Shadow Quark

The friction coefficient arises because of infalling boundary condition.



In practice, for any finite string, sooner or later information will come back and the system will become non-dissipative: a finite length of the string destroys the small- $\omega$  linear term in  $ImG_R$ . A finite quark mass  $M$  will introduce an IR cutoff to the string length.



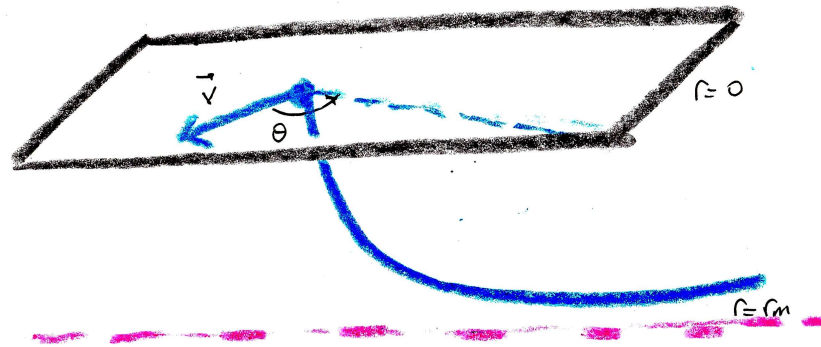
# Physical picture: the Shadow Quark

The friction coefficient arises because of infalling boundary condition.

# Trailing String at finite velocity

Suppose the boundary quark has a constant velocity  $\vec{v}$ . The more general solution is now:

$$X(r, t) = \vec{v}t + \vec{\xi}(r), \quad \xi'(r) = \frac{\vec{c}}{\sqrt{b^4 - C^2}}, \quad |\vec{c}|^2 + \frac{(\vec{v} \cdot \vec{c})^2}{1 - v^2} = C^2$$



Again, the action does not depend on the string direction, and the preferred solution is the one with  $C = b_m^2$ .

# Drag Force

From the classical solution we can compute the net force exerted by the string, as the worldsheet momentum. This corresponds to the vev of the force operator dual to  $\delta X$ :

$$\langle \vec{\mathcal{F}} \rangle = \frac{\sigma_c}{(1 - v^2 \sin^2 \theta)^{3/2}} \begin{pmatrix} \cos \theta \\ (1 - v^2) \sin \theta \cos \varphi \\ (1 - v^2) \sin \theta \sin \varphi \end{pmatrix}$$

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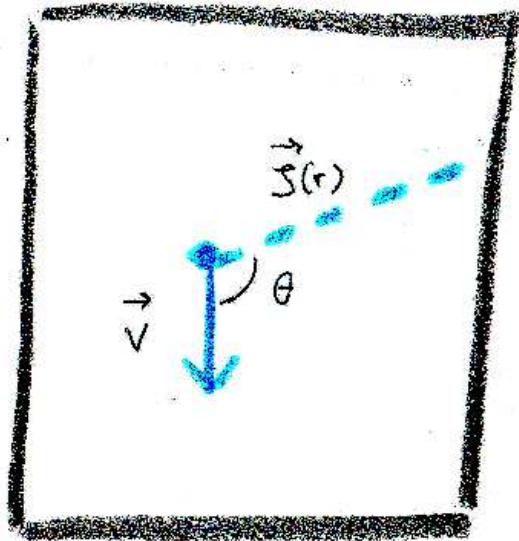
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For  $v \ll 1$  it is directed along the string, and it represents the constant force applied by a string with tension  $\sigma_c$  on its endpoint.

# Fluctuations

Now we have two ways of defining transverse and longitudinal modes:

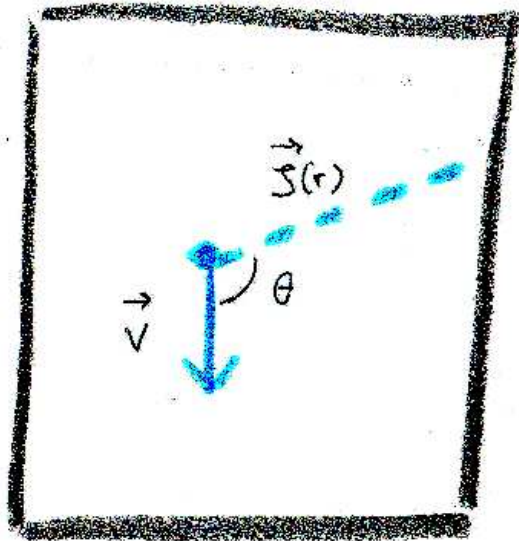
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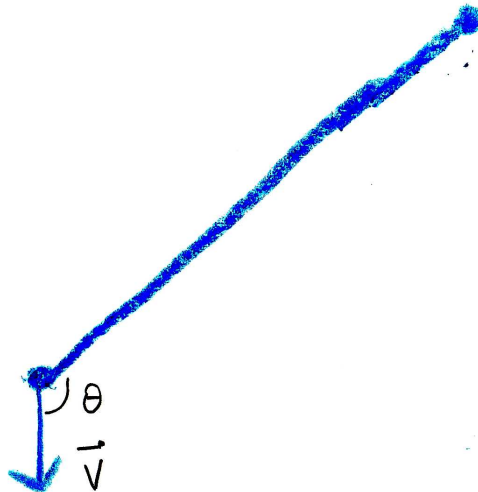


Only the first distinction has a meaning on the boundary.

# Integrating over angles

The way to perform the angular average is not unique.

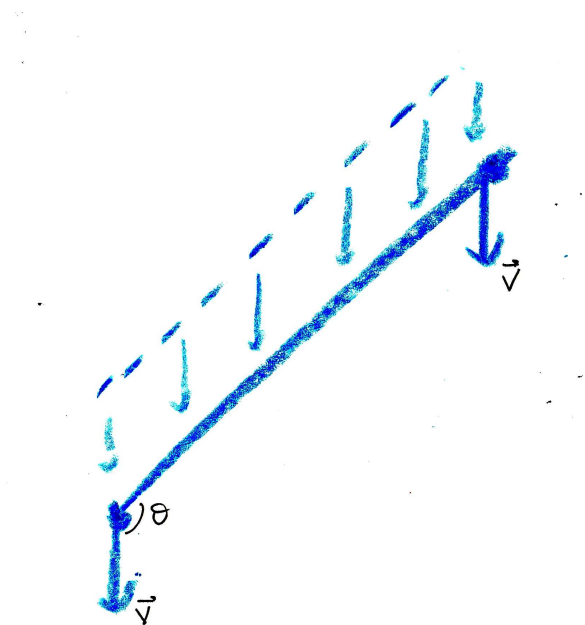
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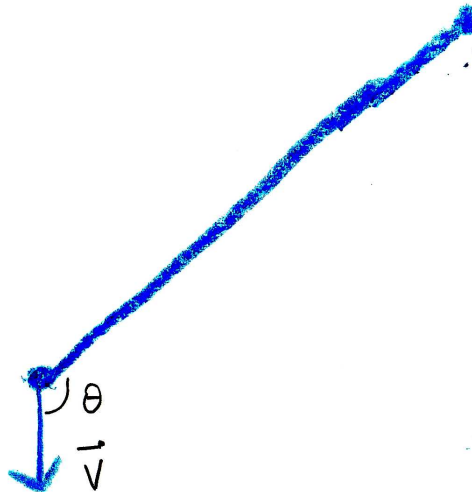




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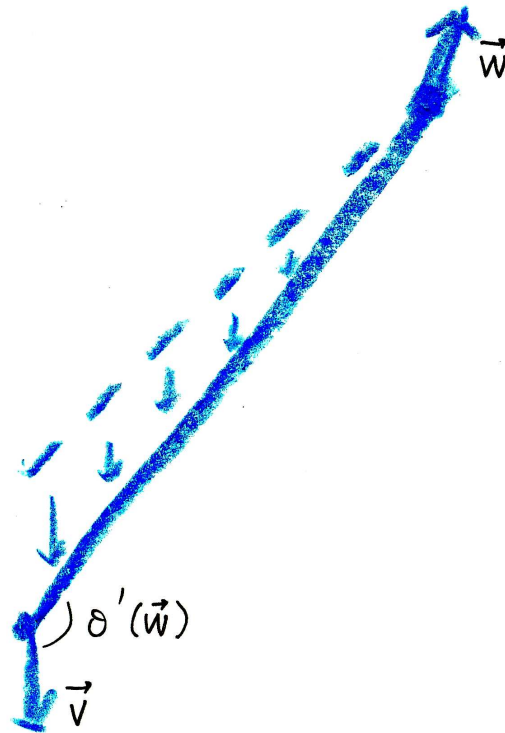
- We can assume the shadow quark has an arbitrary constant velocity  $\vec{w}$  and average over that. As  $t \rightarrow \infty$ , the angle  $\theta$  will be constant and function of  $\vec{w}$  and  $\vec{v}$ :



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$$\cos\theta = \frac{\vec{v} \cdot \vec{w} - v w}{|\vec{v}| |\vec{w} - \vec{v}|}$$

# Averaged correlators

Both averaging procedures can be carried out **analytically**. The result is always in the form:

$$\langle G^{\parallel} \rangle = A(v)G^L(\omega) + B(v)G^T(\omega), \quad \langle G^{\perp} \rangle = C(v)G^T(\omega) + D(v)G^L(\omega)$$

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- $G^{T,L}(\omega)$ : (essentially) the same correlators we found in the static case;
- $A, B, C, D$ : simple functions of  $v$  which **depend on the kind of average**.
- In the static limit we obtain the expected isotropic result:

$$\langle G^{\parallel, \perp} \rangle \rightarrow \frac{1}{3}G^L + \frac{2}{3}G^T \quad v \rightarrow 0$$

- The large- $\omega$  behavior is **universal**.

# Consistency condition

How to decide which average is the correct one?

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Under the same general assumptions that give the Generalized Langevin equation as the effective description of the quark fluctuations, a **low-frequency Ward identity** must be obeyed:

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- the simple angular average over the string direction DOES NOT satisfy this relation;
- the average over the velocity of the shadow quark  $\vec{w}$  DOES.
- This is a consistency check that the shadow-quark picture (complete ignorance about the string boundary at infinity) is the correct one if we want to describe a consistent ensemble



# Conclusion

- We computed the trailing string solution and corresponding fluctuation in confining holographic backgrounds.
- The solution exhibits new features (e.g. spontaneous breaking of isotropy), and interesting dynamics (viscous drag and low frequency modes at zero temperature)
- Integration over moduli gives correlators which can be used to define physical Langevin correlators for the diffusion problem at finite temperature.
- The consistent interpretation of the vacuum correlator is in terms of a quark pair, one of which is very far and unobserved.
- Next: finite mass quarks, comparison with experimental results.