Scalar instantons and the Hartle-Hawking-Maldacena proposal for dS/CFT

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(based on 1406.6148 with Sebastian de Haro)

Instantons in dS/CFT



- Scalar Instantons in Euclidean AdS₄ and dS₄
- 3 Geometric Interpretation
- 4 Discussion

 The HHM proposal: consider scalar fields φ on a fixed dS₄ background i.e.

$$ds^2 = rac{\ell^2}{\eta^2} (-d\eta^2 + dec{x}^2), \ \eta \in (-\infty, 0), \ \phi(\eta, ec{x}) \sim \eta arphi_{ec{k}}(\eta) e^{iec{k}ec{x}}$$

The path integral of configurations born out of the Bunch-Davies vacuum |E⟩ (i.e. having positive frequencies at I⁻) and ending into real configurations φ at late times (i.e. at I⁺) yields the probability for the late time configuration φ

$$\Psi_E[\eta_0,\varphi] \sim e^{iS(\eta_0,\varphi)} \quad , \ \mathcal{P}_{\varphi} = |\Psi_E[\eta_0,\varphi]|^2 \, , \ \eta_0 \to 0$$

• The latter can be used to calculate e.g. late time correlation functions in dS₄

$$\lim_{\eta,\eta'\to 0} \langle E | \phi(\eta,\vec{x}) \phi(\eta',\vec{x}') | E \rangle = \int (\mathcal{D}\varphi) \mathcal{P}_{\varphi} \varphi(\vec{x}) \phi(\vec{x}')$$

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$S_E[\Phi]$ is the *Euclidean* action of all bulk fields, including the metric.

- ∂*M* is a 3-dimensional spacelike hypersurface near the future infinity (i.e. of an asymptotically de Sitter space-time with radius ℓ_{ds}.)
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- The HHM proposal appears to give a boost to dS/CFT, however there is still no completely satisfactory string theory description for gravity with a positive cosmological constant.
- An adventurous set up for a concrete realisation of dS/CFT was proposed by [Anninos et. al (11)] based on Vassiliev's higher-spin (HS) theory which provides the only known consistent classical description of interacting higher-spin gauge fields in a de Sitter background.
- The holographic dual is the Euclidean Sp(*N*) vector model with anti-commuting scalars. It is a free CFT₃ when the higher-spin symmetry is unbroken.
- An analytic continuation is also at work here, as N → −N, where N ~ l²_{dS}/G_N (G_N fixed), maps the Sp(N) anti-commuting vector model to the usual (commuting) O(N) vector model which is believed to be the holographic dual of HS theory on AdS₄.

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- Also, generating functionals for 2-pt functions in EAdS₄ and dS₄ where shown to be related by suitable analytic continuations i.e. *[Harlow et. al. (11)]*
- The HHM proposal has not been tested for cosmological theories with matter fields and generic potentials. This would entail a discussion of exact non-trivial solutions and the analytic continuation of their moduli spaces.
- Also, it is important to know how boundary deformations (such as multitrace deformations) of the CFT dual to EAdS₄ carry over to dS₄, i.e. complexification of the deformation parameters.

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- In AdS/CFT the properly renormalized bulk on-shell action is usually evaluated as a functional of the boundary conditions for a suitably regular solutions of the bulk e.o.m. Depending on the choice of boundary conditions it yields either a generating functional for quantum correlation functions or an effective action for a putative CFT living on the boundary.
- However, given a *particular* solution of the bulk equations of motion, namely one where the bulk fields assume fixed boundary values, the renormalized bulk on-shell action evaluates the *free energy F* of the boundary theory and therefore the partition function as:

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- Conformally coupled scalars in EAdS₄, in the Poincaré patch

$$\phi(z, \vec{x}) \to z \phi_{(0)}(\vec{x}) + z^2 \phi_{(1)}(\vec{x}), \ z \to 0$$

• When $\phi_{(0)}(\vec{x})$ takes a *fixed* form - instanton solutions

$$S_{\text{on-shell}}[\phi_{(0)}] = -\ln \tilde{Z}_0$$

gives the partition function of the *dual* boundary CFT, that has in its spectrum the operator O_1 , $\Delta = 1$ and $\langle O_1 \rangle \sim \phi_{(0)}$.

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Consider the Euclidean action

$$S = \frac{1}{2\kappa^2} \int_{M_{\epsilon}} d^4 x \sqrt{g} \left(-R + 2\Lambda\right)$$
$$+ \int_{M_{\epsilon}} d^4 x \sqrt{g} \left(\frac{1}{2} \left(\partial_{\mu}\phi\right)^2 + \frac{R}{12} \phi^2 + \frac{\lambda}{4!} \phi^4\right)$$
$$- \frac{1}{2\kappa^2} \int_{\partial M_{\epsilon}} d^3 x \sqrt{\gamma} \, 2K \left(1 - \frac{\kappa^2}{6} \phi^2\right),$$

where $\kappa^2 = 8\pi G_N$ and $\Lambda = \mp \frac{3}{\ell^2}$.

• We add the usual counter terms in the conformal boundary ∂M_{ϵ}

$$S_{\rm ct}^{\rm EAdS} = \frac{1}{\kappa^2} \int_{\partial M_{\epsilon}} d^3 x \, \sqrt{\gamma} \left(\frac{4}{\ell} + \ell \, R[\gamma] \right)$$

• For the Euclidean dS₄ case, no counterterms are needed since metric configurations are asymptotically regular.

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$$\mathrm{d} s^2_{\scriptscriptstyle\mathsf{EAdS}} = rac{1}{\sinh^2 rac{ au}{\ell}} \left(\mathrm{d} au^2 + \ell^2_{\scriptscriptstyle\mathsf{AdS}} \, \mathrm{d} \Omega^2_3
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 The relevant solutions of the scalar sector are obtained by solving the Klein-Gordon equation

$$\Box \phi - \frac{2\Lambda}{3} \phi - \frac{\lambda}{6} \phi^3 = 0$$

together with the requirement that the stress-energy tensor vanishes. The latter requirement turns out to give

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• The solutions on EAdS₄ and S^4 are then given by

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• The moduli spaces are non-trivial,

on EAdS₄:
$$-b_0^2 + b_5^2 + b_i^2 = \frac{\lambda_{AdS}}{12} \ell_{AdS}^2$$
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- For example, the moduli space of the solutions is EAdS₄ if $\frac{\lambda_{AdS}}{12} \ell_{AdS}^2 < b_5^2$, which is the condition required for regularity of the bulk solution. At the critical value $\frac{\lambda_{AdS}}{12} \ell_{AdS}^2 = b_5^2$ the effective potential of the dual field theory becomes unbounded from below, which was interpreted as an instability against marginal deformations [*de Haro (06)*].

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 Not all of the parameters a_I, b_I (I = 0,...,5) are moduli. Two of them are boundary conditions relating the leading and subleading modes of the scalar:

$$\phi_{S_{-}^{4}}(-\epsilon\,\ell,\Omega) = \frac{\varepsilon}{a_{5}+a_{i}\Omega_{i}} + \frac{\epsilon\,\varepsilon\,a_{0}}{(a_{5}+a_{i}\Omega_{i})^{2}} + \mathcal{O}(\epsilon^{2}) = \Phi_{(0)}^{a_{5},a_{i}} + \epsilon\,\Phi_{(1)}^{a_{5},a_{i}} + \dots$$
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where $\alpha = b_5$ for EAdS₄ and $\alpha = a_0$ for S_{-}^4 . Thus, b_5 , a_0 parametrize marginal triple trace deformations that change the boundary conditions from Dirichlet to mixed.

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• The solutions are exact: they have zero stress-energy tensor hence the EAdS/dS background stays unmodified. Thus we can compute the on-shell effective action including its finite part. For simplicity we set the spherical modes $a_i = b_i = 0$ and get:

$$\begin{split} S_{\text{EAdS}}^{\text{on-shell}} &= \frac{4\pi^2 \ell_{\text{AdS}}^2}{\kappa^2} - \frac{\lambda_{\text{AdS}} \pi^2 \ell_{\text{AdS}}^4}{12} \frac{2b_0 + b_5}{3b_0^3 (b_0 + b_5)^2} + \pi^2 \ell_{\text{AdS}}^2 \frac{b_5}{b_0^3} + \mathcal{O}(\epsilon) \\ &= \frac{4\pi^2 \ell_{\text{AdS}}^2}{\kappa^2} + \frac{2\pi^2 \ell_{\text{AdS}}^2}{3b_0^3} \frac{b_0^2 + b_0 b_5 + b_5^2}{b_0 + b_5} + \mathcal{O}(\epsilon), \end{split}$$

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Instantons in dS/CFT

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- The matching under analytic continuation of the gravity part is already nontrivial i.e. the EAdS₄ calculation needs to be regularized and renormalized, whereas the S⁴₋ calculation of the HH wave function is completely finite.
- In order to match the matter contributions, however, in the second term we need to analytically continue the couplings as well. This analytic continuation from EAdS₄ to S⁴₋ is an invertible map γ:

$$\gamma(\ell_{AdS}) = i \, \ell_{dS} \,, \gamma(b_0) = i \, a_5 \,, \gamma(b_5) = -i \, a_0 \,, \gamma(b_i) = i a_i \;.$$

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• The moduli have to be analytically continued since they are dimensionful quantities, to be measured in units of the radius. Defining dimensionless moduli $y_l = a_l/\ell_{dS}$, $z_l = b_i/\ell_{AdS}$, l = 0, ..., 5, we find

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• These are all real quantities on both sides. The moduli space is O(1,5) invariant:

$$\eta^{IJ} y_I y_J = \frac{\lambda_{dS}}{12} , \ I, J = 0, \dots, 5 ,$$

with η^{IJ} the O(1,5) Minkowski metric. The analytic continuation is then simply an SO(1,5) map of the moduli space onto itself:

$$Z_{I} = \epsilon_{I}^{J} y_{J}; \quad \epsilon = \delta \begin{pmatrix} 0 & 1 \\ -1 & 0 \\ & \mathbf{1}_{4 \times 4} \end{pmatrix}, \quad \delta^{2} = 1.$$

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 We note here that in order to enforce the conformal boundary conditions to the solutions of the bulk equations of motion, one needs to add a further "triple trace deformation" term to the action

$$S_{
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This of course agrees, after the analytic continuation with the term one gets in the dS₄ case, $-\frac{2\pi^2 \ell_{dS}^2 a_0}{3a_5^3}$. These terms simply add up to the solutions above.

 The 3-sphere partition function Z_{S³} of a three-dimensional CFT is a measure of its degrees of freedom. It has been argued that for unitary CFTs the corresponding free energy is given (in suitably chosen units) by

$$F=-\log\left|Z_{\mathcal{S}^{3}}\right|,$$

which is positive and satisfies an *F*-theorem, namely it decreases along RG flows from the UV to the IR [*Jafferis (10)*].

- Holographically, the partition function is usually calculated using the bulk gravitational action on EAdS₄ with matter fields set to zero. This is the first term of the results above. It is proportional to the dimensionless ratio ℓ_{AdS}^2/κ^2 .
- One may then wonder what the physical interpretation is of the second term in which corresponds to the contribution of the bulk scalar fields. Notice that for $\lambda_{AdS} < 0$ this term is also positive.

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 The 3-sphere partition function Z_{S³} of a three-dimensional CFT is a measure of its degrees of freedom. It has been argued that for unitary CFTs the corresponding free energy is given (in suitably chosen units) by

$$F=-\log\left|Z_{\mathcal{S}^{3}}\right|,$$

which is positive and satisfies an *F*-theorem, namely it decreases along RG flows from the UV to the IR [*Jafferis (10)*].

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• Consider the conformally related metrics

$$g_{\mu
u}=\Omega^{-2}\,h_{\mu
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• One can then show that

$$-I_h := \frac{1}{2\kappa^2} \int \mathrm{d}^4 x \sqrt{h} \left(R[h] - 2\Lambda_h \right)$$

$$= \int d^4x \sqrt{g} \left(\frac{1}{2} \left(\partial_\mu \phi \right)^2 + \frac{R[g]}{12} \phi^2 + \frac{\lambda}{4!} \phi^4 \right) - \frac{1}{2} \int d^4x \ \partial_\mu \left(\sqrt{g} g^{\mu\nu} \phi \partial_\nu \phi \right)$$
$$=: I_{(g,\phi)}$$

where we have defined

$$\phi := \sqrt{rac{\kappa^2}{6}} \, \Omega \,, \qquad \lambda := -rac{2\kappa^2}{3} \, \Lambda_h \,.$$

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• One can also show that on the instanton solutions

$$\frac{1}{2}\int_{\partial M_{\epsilon}}\mathrm{d}^{3}x\sqrt{g}g^{00}\phi\partial_{0}\phi=-\frac{1}{2\kappa^{2}}\int_{\partial M_{\epsilon}}\mathrm{d}^{3}x\sqrt{\gamma}2K\frac{\kappa^{2}}{6}\phi^{2}$$

• Then, the on-shell value of the instanton action $I_{(g,\phi)}$ gives

$$I_{(g,\phi)}^{\text{on-shell}} = -\frac{\lambda}{4!} \int \mathrm{d}^4 x \, \sqrt{g} \; \phi_{\text{inst}}^4 = \frac{\Lambda_h}{\kappa^2} \int \mathrm{d}^4 x \; \sqrt{h} \equiv -I_h^{\text{on-shell}} \, ,$$

which is *minus* the on-shell value for the Einstein-Hilbert action I_h , since $R[h] = 4\Lambda_h$.

- The crucial observation now is that the instanton solutions ϕ_{inst} on either S_{-}^4 and EAdS₄, with the moduli set to specific values, correspond exactly to the conformal factor relating the two metrics
- Hence, the on-shell action of instantons on EAdS₄ corresponds to the volume of (half) S⁴ and conversely, the on-shell action of instantons on the half S⁴ corresponds to the volume of EAdS₄,

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- The instanton deformation parameter *b* (which corresponds to *a*₀ in the previous section) regulates the volume of EAdS₄.
- Let us see how this arises. On the half S^4 with curvature radius ℓ and metric

$$\mathrm{d}s^2 = \frac{4}{\left(1 + \frac{\rho^2}{\ell^2}\right)^2} \left(\mathrm{d}\rho^2 + \rho^2 \mathrm{d}\Omega_3^2\right)$$

the instanton solutions are given by

$$\phi_{\text{inst}}(\rho) = \pm \sqrt{\frac{12}{\lambda_{S^4}}} \frac{1}{b} \frac{1 + \frac{\rho^2}{\ell^2}}{1 - \frac{\rho^2}{b^2}}$$

where $\lambda_{S^4} > 0$.

 It is important to note the presence of the instanton modulus b which is in principle unrelated to ℓ. In particular, since the range of the radial coordinate is ρ ∈ [0, ℓ), if we consider b > ℓ the solution is everywhere regular.

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 Next we notice that half S⁴ and EAdS₄ are conformally related metrics. In particular

$$ds^{2} = \frac{4}{\left(1 + \frac{\rho^{2}}{\ell^{2}}\right)^{2}} \left(d\rho^{2} + \rho^{2}d\Omega_{3}^{2}\right) = \\ = \left(\frac{1 - \frac{\rho^{2}}{b^{2}}}{1 + \frac{\rho^{2}}{\ell^{2}}}\right)^{2} \frac{4}{\left(1 - \frac{\rho^{2}}{b^{2}}\right)^{2}} \left(d\rho^{2} + \rho^{2}d\Omega_{3}^{2}\right)$$

where on the right EAdS₄ has a generally different radius *b* which is set equal to the instanton modulus. Hence, the calculation of the on-shell action for instantons on half S^4 boils down to the calculation of the *regularized* volume of EAdS₄.

A. C. Petkou (AUTH)

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 Notice that it is the presence of the instanton modulus b > ℓ that gives rise to a particular regularization of the volume of the EAdS₄ space with radius ℓ.

• Explicitly, we obtain $(\alpha = \frac{\ell}{b} < 1)$

$${}^{\text{on-shell}}_{(g,\phi)}\left(S^4_{-}
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• Next we consider the instanton solutions on $EAdS_4$ with radius ℓ . These have the form

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with $\lambda_{AdS} < 0$ and *b* the modulus which we take here again to be $b > \ell$. The on-shell action would give part of the volume of S^4 with radius *b*. We find

$$J_h^{\text{on-shell}}\left(\mathsf{EAdS}_4\right) = rac{8\pi^2 b^2}{\kappa^2} rac{lpha^4(lpha^2+3)}{(lpha^2+1)^3} = -rac{16\pi^2}{\lambda_{ ext{AdS}}} rac{lpha^4(lpha^2+3)}{(lpha^2+1)^3} \ ,$$

where here $\lambda_{AdS} < 0$.

- The EAdS₄ and dS₄ results are mapped to each other by the transformation b → i b. This is natural, if we recall that b is the radius of both the S⁴ and EAdS₄ of the associated EH actions.
- Thus, we have shown that the on-shell instanton action on EAdS₄ can be viewed as a partition function on a 3-sphere of radius *b*, arising from a bulk gravitational action with Newton's constant κ , if we identify the coupling $\lambda \sim \kappa^2/b^2$.
- The instantons in this part are related to the previously discussed ones by

$$b^2 = \ell_{\rm dS}^2 \left(1 - \frac{a_5}{a_0}\right),$$

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- It would also be interesting to work out the relationship of this result with holographic stochastic quantisation.

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 We also found an interesting geometric realization of the same computation, in which the S^4 instantons (for particular values of the moduli) are seen to compute the regularized volume of $EAdS_4$, and the $EAdS_4$ instantons are seen to compute the volume of the four-sphere. The regulator of the EAdS₄ volume is a_0 , with the divergence appearing precisely for the critical value $a_{n}^{2} = \frac{\ell_{S^{4}}}{12} \ell_{ds}^{2}$. This might imply that a sector of the Sp(N) model with this marginal deformation is dual to a pure gravitational theory with no scalars, and hence signal a duality between Sp(N)models with different values of the deformation parameter.

- Our instanton solutions are intimately related to the SO(4) and SO(3, 1) invariant solutions of 4-dimensional HS theory found by Sezgin and Sundell, which are also related to a consistent truncation of $\mathcal{N} = 8$ gauge supergravity down to a single scalar of the SO(8) group. In that sense, our results should also provide the partition functions of both the above theories, at the scalar instanton vacua.
- Finally, our instanton partition functions describe nucleation of spherical and hyperbolic bubbles. Fluctuations around these solutions should give the correlation functions on these situations.

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