

07.09.2014

Holographic Superconductors ...In Helical Backgrounds & Homes' Relation

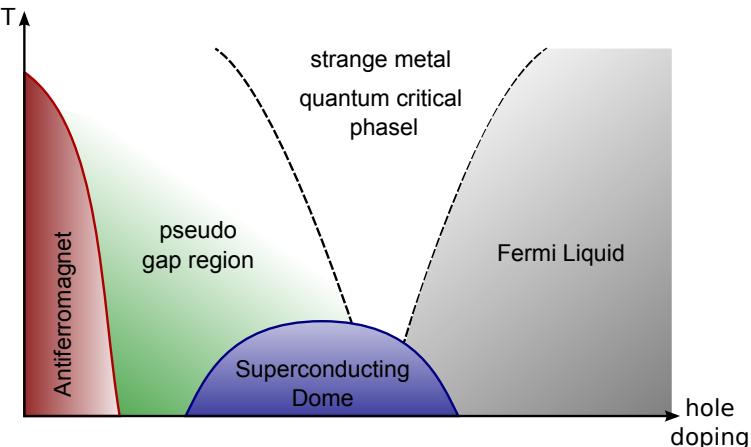
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René Meyer & Koenraad Schalm

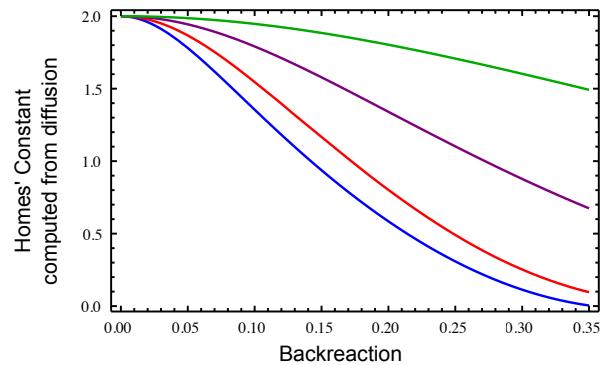
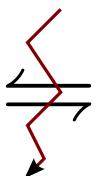
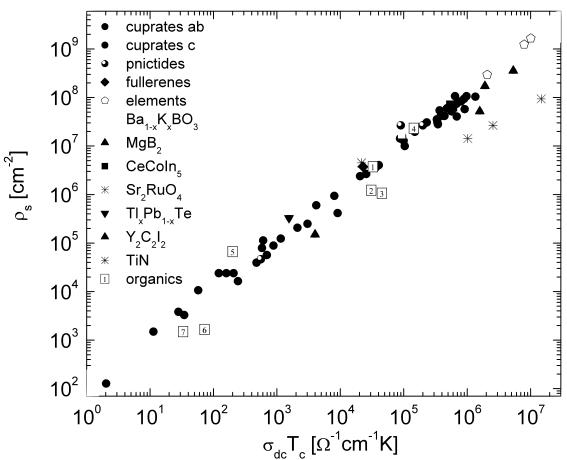
Quantum Field Theory, String Theory & Condensed Matter Physics
Κολυμβαρι

Motivation

- ▶ Holography is a powerful tool to describe strongly correlated systems
- ▶ Experimental puzzling results: Homes' Relation



$$\rho_s = C \sigma_{DC} T_c$$



[Erdmenger,Kerner,SK '12]

[Dordevic,Basov,Homes '12]

Motivation

- ▶ Homes relation cannot hold in "pure" holographic s-wave superconductors, where transport is related to diffusion
- ▶ Momentum is intrinsically conserved in Holography
 - Impossible to discern an ideal metal from a superconductor
- ▶ How to incorporate momentum relaxation?

Q-Lattice

[Donos & Gauntlett '13]

Massive Gravity

[Vegh '13]

Helical Lattice

[Donos & Hartnoll '12]

Inhomogenous Background

[Hartnoll & Hofman '12]

[Horowitz, Santos & Tong '12]

- ▶ Holographic Setup
- ▶ Thermodynamics of the Superfluid
Insulator Transition
- ▶ Optical Conductivity at Finite Temperature
- ▶ Homes' Relations of high T_c Superconductors
- ▶ Zero Temperature Solutions

Part I:

Holographic Setup

Holographic Setup

► Action

$$S = \int \left[R e^{\wedge 3} + \star 12 - \frac{1}{4} F \wedge \star F - \frac{1}{4} W \wedge \star W - m^2 B \wedge \star B \right.$$
$$\left. - \int d^{4+1}x \left[|\partial \chi - iqA\chi|^2 + V(|\chi|) \right] \right]$$
$$- \frac{\kappa}{2} \int B \wedge F \wedge W$$

Gauge fields

$$F = dA$$

$$W = dB$$

Gravity/Geometry

$$0 = de + \Omega \wedge e$$

$$R = d\Omega + \Omega \wedge \Omega$$

$$ds^2 = "e * e"$$

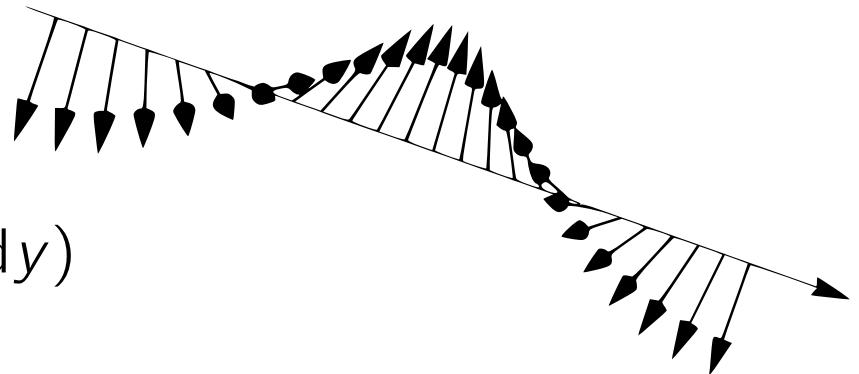
Holographic Setup

- Bianchi VII Ansatz for the metric

$$ds^2 = -U(r) dt^2 + \frac{dr^2}{U(r)} + e^{2v_1(r)} \omega_1^2 + e^{2v_2(r)} \omega_2^2 + e^{2v_3(r)} \omega_3^2$$

$$\omega_1 = dx$$

$$\omega_2 + \omega_3 = e^{ip \cdot x} (dx + i dy)$$



- Asymptotically AdS spacetime $r \rightarrow \infty$

$$U(r) = r^2 \quad v_i(r) = \log(r) \quad \text{for } i = 1, 2, 3$$

Holographic Setup

- ▶ Ansatz for the gauge fields

→ encodes finite density ←

$$A = a(r) dt$$

$$a(\infty) = \mu$$

$$B = w(r)\omega_2$$

$$w(\infty) = \lambda$$

→ generates helix
with "lattice strength" λ ←

- ▶ Homogenous charged scalar field

$$\chi = \chi(r)$$

Equation of Motion

$$\left. \begin{aligned}
 0 &= a'' + a' (v'_1 + v'_2 + v'_3) - \frac{2aq^2\rho^2}{U} + \kappa p e^{-\nu_1-\nu_2-\nu_3} w w', \\
 0 &= w'' + w' \left(\frac{U'}{U} + v'_1 - v'_2 + v'_3 \right) + \frac{w}{U} \left(\kappa p e^{-\nu_1+\nu_2-\nu_3} a' - m^2 - p^2 e^{-2(\nu_1-\nu_2+\nu_3)} \right), \\
 0 &= 2\rho^2 \left(m_\rho^2 - \frac{a^2 q^2}{U} \right) + a'^2 + w^2 \left(m^2 e^{-2\nu_2} + p^2 e^{-2(\nu_1+\nu_3)} \right) + 4p^2 e^{-2\nu_1} \sinh^2 (\nu_2 - \nu_3) \\
 &\quad + 2U' (v'_1 + v'_2 + v'_3) - U (2\rho'^2 + e^{-2\nu_2} w'^2 - 4v'_1 v'_2 - 4v'_1 v'_3 - 4v'_2 v'_3) - 24, \\
 0 &= 2\rho^2 \left(\frac{m_\rho^2}{U} - \frac{a^2 q^2}{U^2} \right) - \frac{a'^2}{U} + \frac{w^2}{U} \left(m^2 e^{-2\nu_2} - p^2 e^{-2(\nu_1+\nu_3)} \right) \\
 &\quad + \frac{p^2}{U} \left(-2e^{-2\nu_1} + 3e^{-2(\nu_1+\nu_2-\nu_3)} - e^{-2(\nu_1-\nu_2+\nu_3)} \right) + 2\rho'^2 + \frac{2U''}{U} \\
 &\quad + 4 \left(\frac{U'}{U} (v'_1 + v'_2) + v'_1^2 + v'_2 v'_1 + v'_2^2 \right) - \frac{24}{U} + e^{-2\nu_2} w'^2 + 4(v''_1 + v''_2), \\
 0 &= \frac{2w^2}{U} \left(p^2 e^{-2(\nu_1+\nu_3)} - m^2 e^{-2\nu_2} \right) + \frac{4p^2}{U} \left(e^{-2(\nu_1-\nu_2+\nu_3)} - e^{-2(\nu_1+\nu_2-\nu_3)} \right) \\
 &\quad + \frac{4U'}{U} (v'_3 - v'_2) - 2e^{-2\nu_2} w'^2 + 4(-v'_2^2 - v'_1 v'_2 + v'_3^2 + v'_1 v'_3) + 4(v''_3 - v''_2), \\
 0 &= \frac{p^2}{U} \left(e^{-2\nu_1} - e^{-2(\nu_1+\nu_2-\nu_3)} \right) + \frac{U'}{U} (v'_3 - v'_1) - v'_1^2 + v'_3^2 - v'_1 v'_2 + v'_2 v'_3 - v''_1 + v''_3, \\
 0 &= \rho'' + \rho' \left(\frac{U'}{U} + v'_1 + v'_2 + v'_3 \right) + \rho \left(\frac{a^2 q^2}{U^2} - \frac{m_\rho^2}{U} \right)
 \end{aligned} \right\}$$

Only ODEs
as
promised
by
Bianchi VII

Numerical Solution

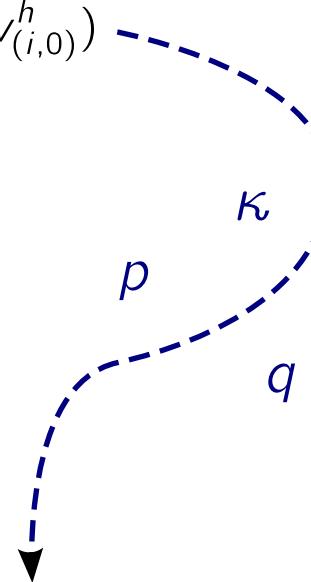
► Expansion at black hole horizon

$$\left. \begin{array}{l} a = a_1^h (r - r_h) + a_2^h (r - r_h)^2 + \dots, \\ w = w_0^h + w_1^h (r - r_h) + \dots, \\ U = U_1^h (r - r_h) + U_2^h (r - r_h)^2 + \dots, \\ v_i = v_{(i,0)}^h + v_{(i,1)}^h (r - r_h) + \dots, \\ \rho = \rho_0^h + \rho_1^h (r - r_h) + \dots, \end{array} \right\} \quad T = \frac{U_1^h}{4\pi} \quad (a_1^h, w_0^h, \rho_0^h, U_1^h, v_{(i,0)}^h)$$

► Expansion at AdS boundary

$$\left. \begin{array}{l} a = \mu + \frac{\nu}{r^2} + \dots, \\ w = \lambda + \frac{\beta - p^2 \lambda \log(r)/2}{r^2} + \dots, \\ \rho = \frac{\rho_b}{r^4} - \frac{q^2 \mu^2 \rho_b}{12r^6} + \dots, \\ U = r^2 - \frac{\epsilon/3 + p^2 \lambda^2 \log(r)/6}{r^2} + \dots, \\ v_i = \log(r) + \frac{g_i - (-1)^i p^2 \lambda^2 \log(r)/24}{r^4} + \dots, \end{array} \right\} \quad (\epsilon, g_1, g_2, \mu, \nu, \lambda, \beta, \rho_b)$$

Shooting Method



Part I:

Superfluid Phase Transition

Thermodynamics

- Free energy density evaluated at the boundary

$$\frac{\Omega}{V} = \epsilon + \mu \underbrace{\left(2\nu - \frac{1}{2} \kappa \lambda^2 p \right)}_{= n} - \underbrace{e^{v_1(r_H) + v_2(r_H) + v_3(r_H)} U'(r_H)}_{= Ts}$$

$$= 4g_1 + \frac{\alpha^4}{16} - \frac{\epsilon}{3} - \beta\lambda - \frac{\lambda^2 p^2}{8}$$

shift symmetry $\rightarrow r \rightarrow r + \alpha/2$

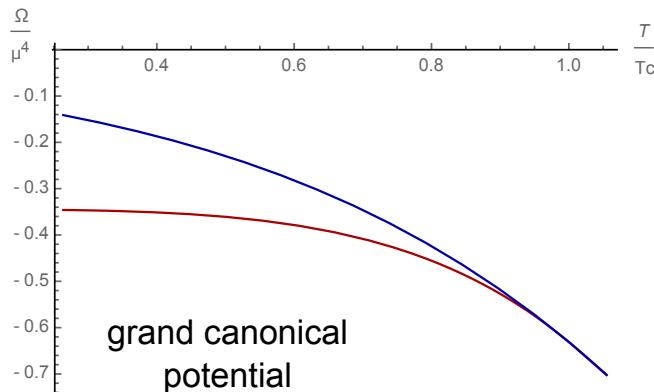
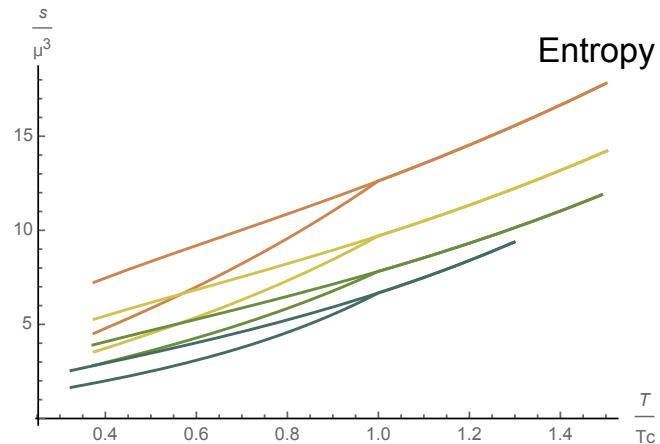
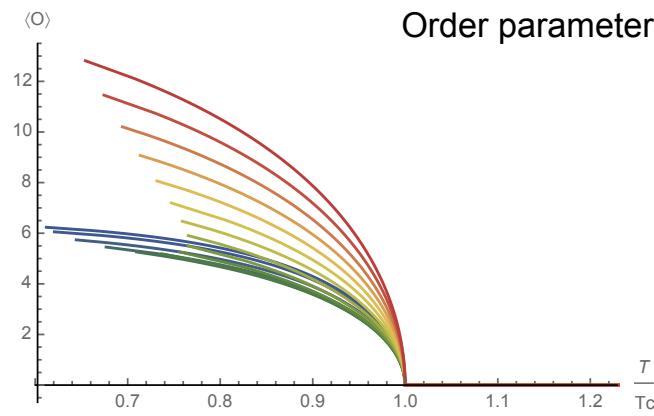
conformal anomaly $4 \langle T_a^a \rangle$

- Physical parameters we can freely tune

$$(\kappa, q, T/\mu, p/\mu, \lambda/\mu)$$

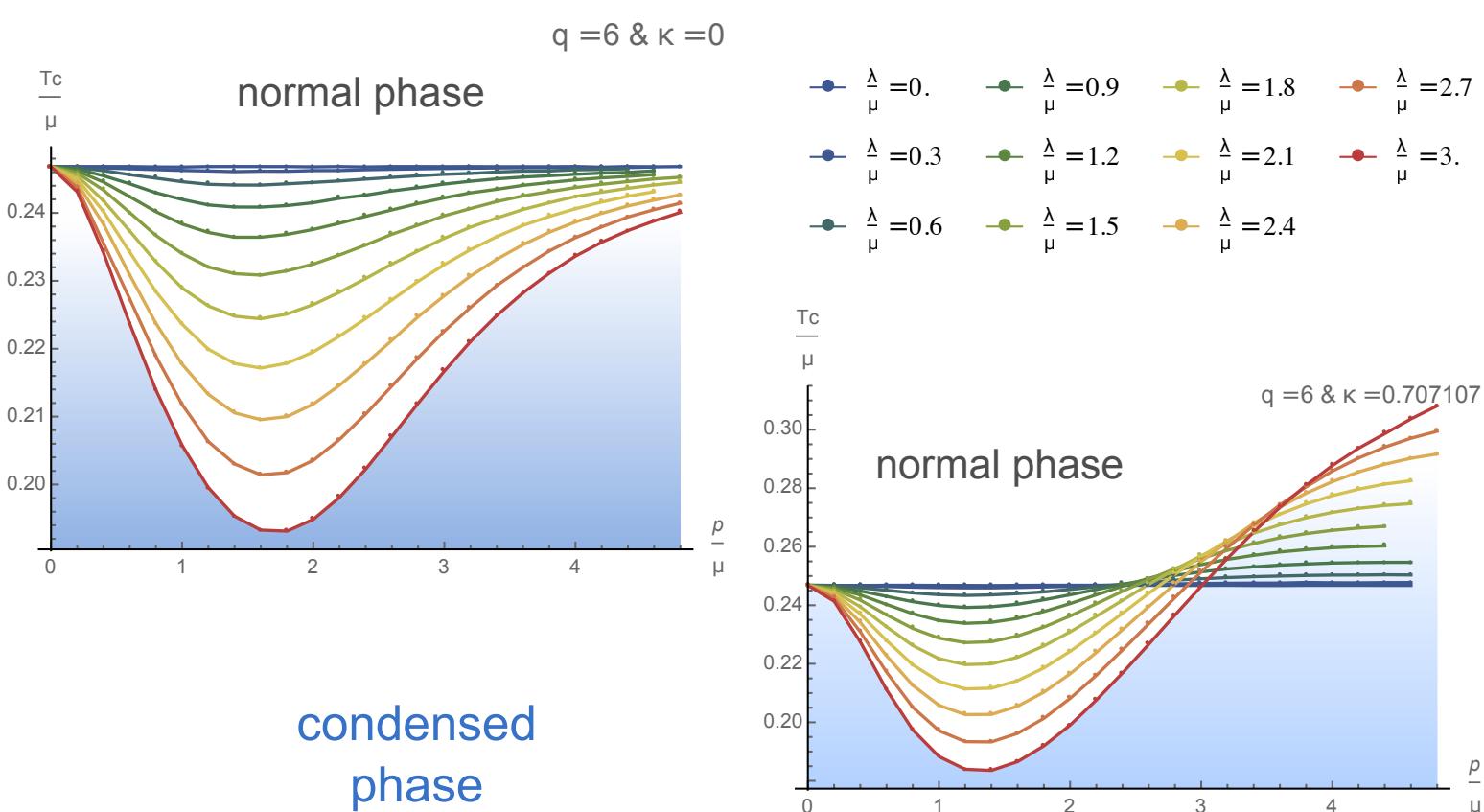
Phase Transition

► Mean field insulator-superfluid transition

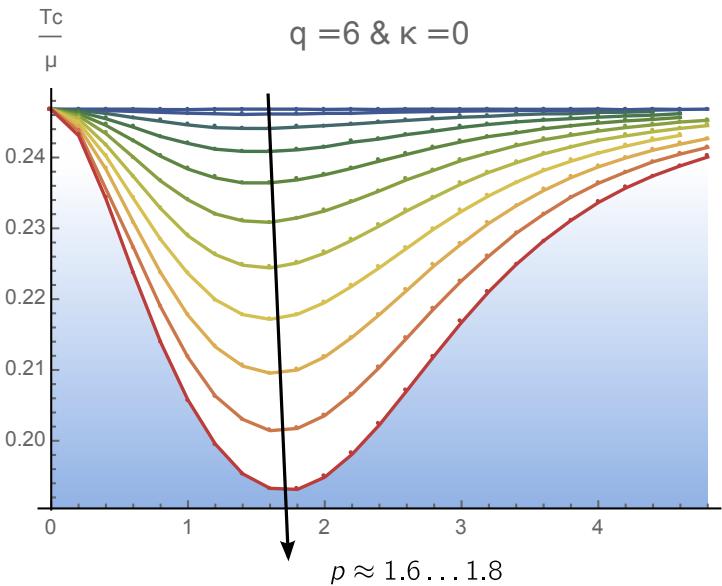


Large N mean field
second order
phase transition

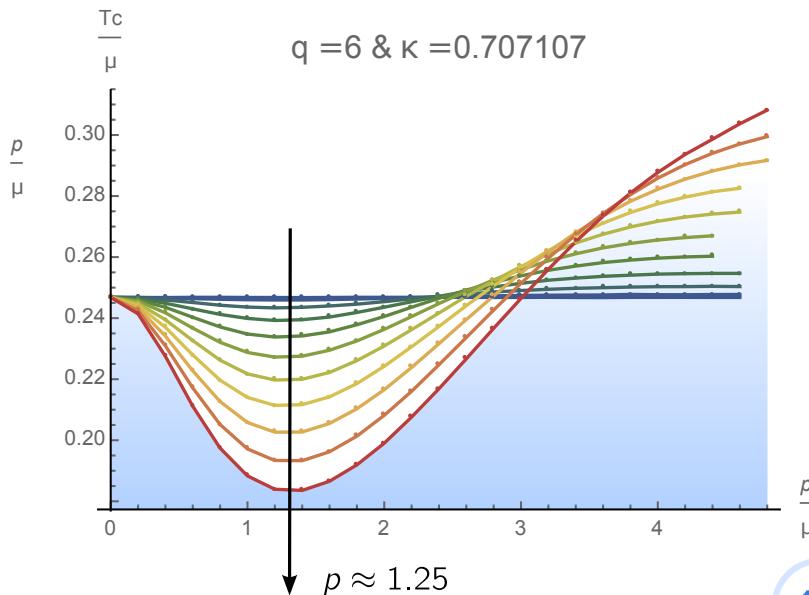
Phase Diagram



Minimal Tc @ Fixed p

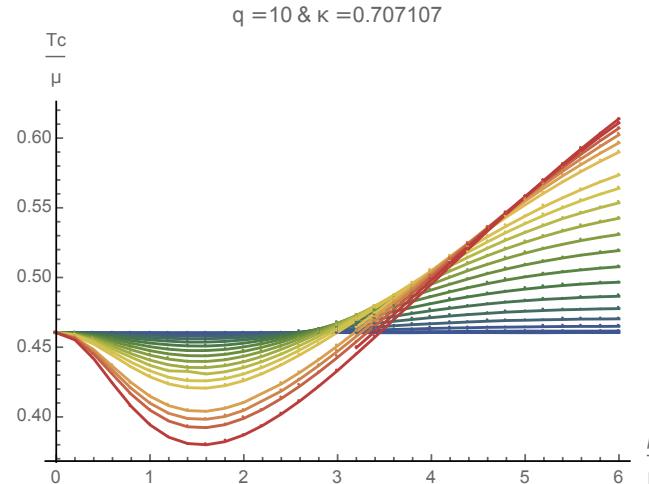
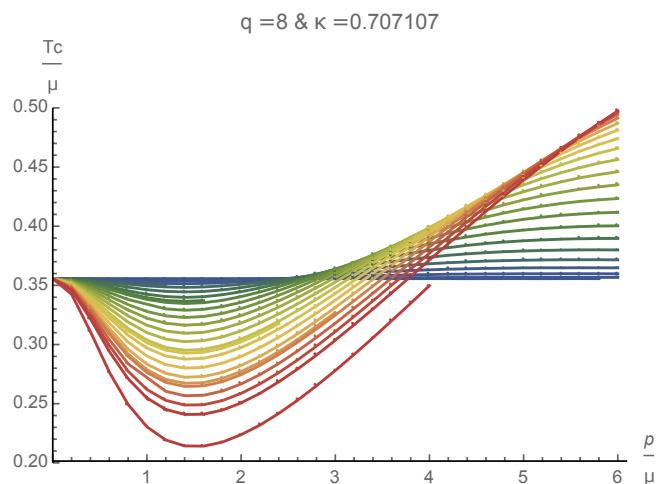
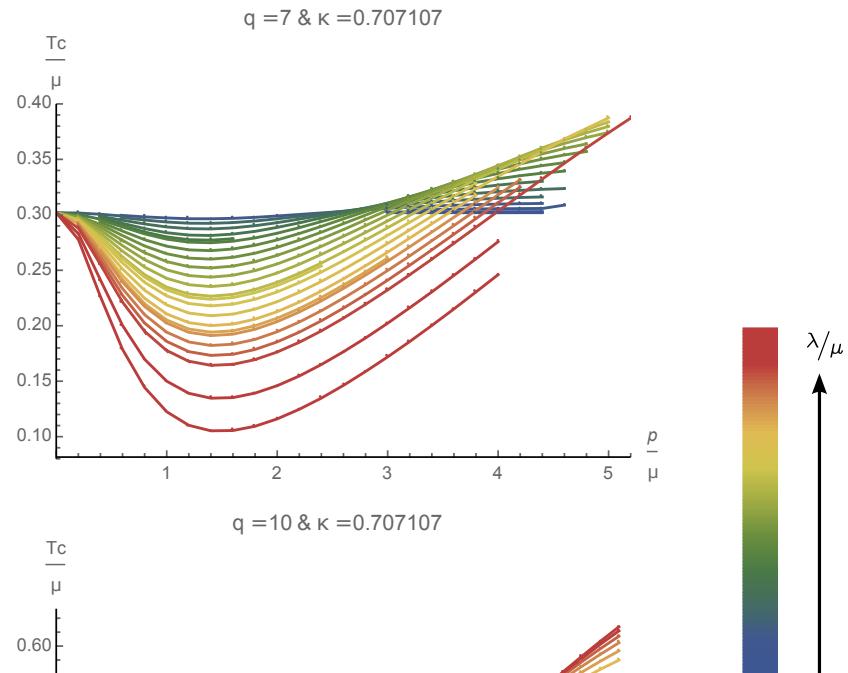
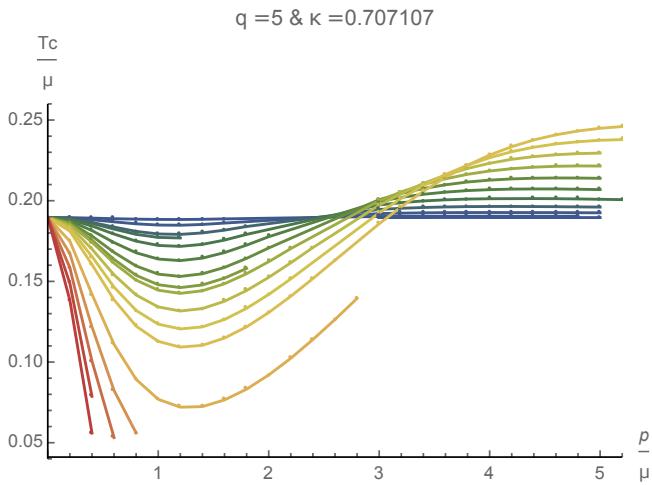


● $\frac{\lambda}{\mu} = 0.$	● $\frac{\lambda}{\mu} = 0.9$	● $\frac{\lambda}{\mu} = 1.8$	● $\frac{\lambda}{\mu} = 2.7$
● $\frac{\lambda}{\mu} = 0.3$	● $\frac{\lambda}{\mu} = 1.2$	● $\frac{\lambda}{\mu} = 2.1$	● $\frac{\lambda}{\mu} = 3.$
● $\frac{\lambda}{\mu} = 0.6$	● $\frac{\lambda}{\mu} = 1.5$	● $\frac{\lambda}{\mu} = 2.4$	

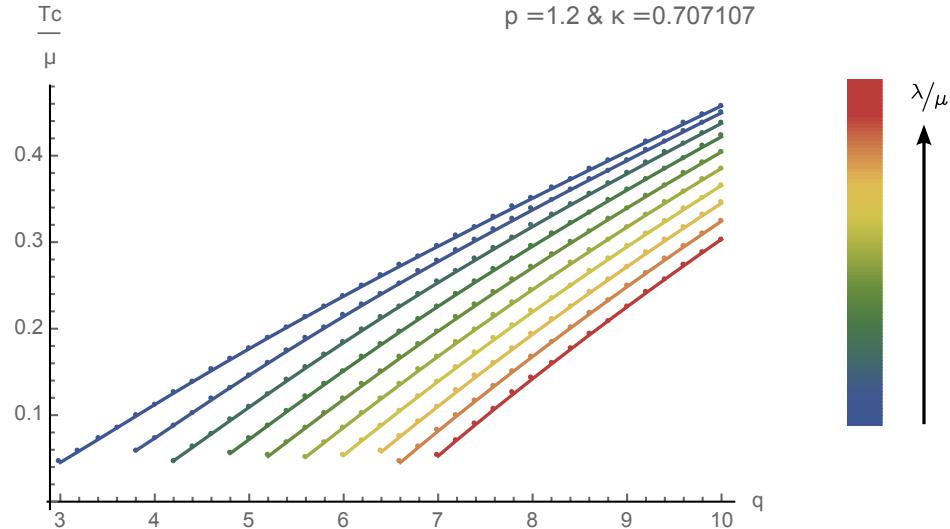


- ▶ Minimal Tc occurs at a fixed p for all λ
- ▶ p-value increases with decreasing q

Phase Diagram



Variation of Tc with q



- ▶ T_c is reduced when increasing backreaction controlled by q
 - matches known behavior of holographic s-wave superconductor in pure AdS-RN background

Part II:

Optical Conductivity

Fluctuation Equations of Motion

	h_{tt}	h_{rr}	h_{11}	h_{22}	h_{33}	h_{tr}	h_{t1}	h_{r1}	h_{t2}	h_{t3}	h_{r2}	h_{r3}	h_{12}	h_{13}	h_{23}	A_t^f	A_r^f	A_1^f	A_2^f	A_3^f	B_t^f	B_r^f	B_1^f	B_2^f	B_3^f	ρ^f	$(\rho^f)^*$
h_{tt}	•	•	•	•	•	•										•	•					•		•	•		•
h_{rr}		•	•	•	•	•										•	•					•		•	•		•
h_{11}			•	•	•	•										•	•					•		•	•		•
h_{22}				•	•	•										•	•					•		•	•		•
h_{33}					•	•										•	•					•		•	•		•
h_{tr}						•																•			•		•
h_{t1}							•																				
h_{t2}								•																			
h_{t3}									•																		
h_{r1}										•																	
h_{r2}											•																
h_{r3}												•															
h_{12}													•														
h_{13}														•													
h_{23}															•												
A_t^f																•	•										
A_r^f																	•										
A_1^f																		•									
A_2^f																			•								
A_3^f																				•							
B_t^f																				•							
B_r^f																					•						
B_1^f																					•						
B_2^f																						•					
B_3^f																							•				
ρ^f																								•			
$(\rho^f)^*$																									•		
$\delta A = \mathcal{A}(t, r) \omega_1$																											
$\delta B = \mathcal{B}(t, r) \omega_3$																											
$\delta (ds^2) = h_{t1}(t, r) dt \otimes \omega_1$																											
$+ h_{23}(t, r) \omega_2 \otimes \omega_3 + h_{r1}(t, r) dr \otimes \omega_1$																											

Transport Properties

- Horizon expansion of fluctuation equations

$$\mathcal{A} = (r - r_h)^{\pm i\omega/(4\pi T)} (A_0^h + A_1^h (r - r_h) + \dots)$$

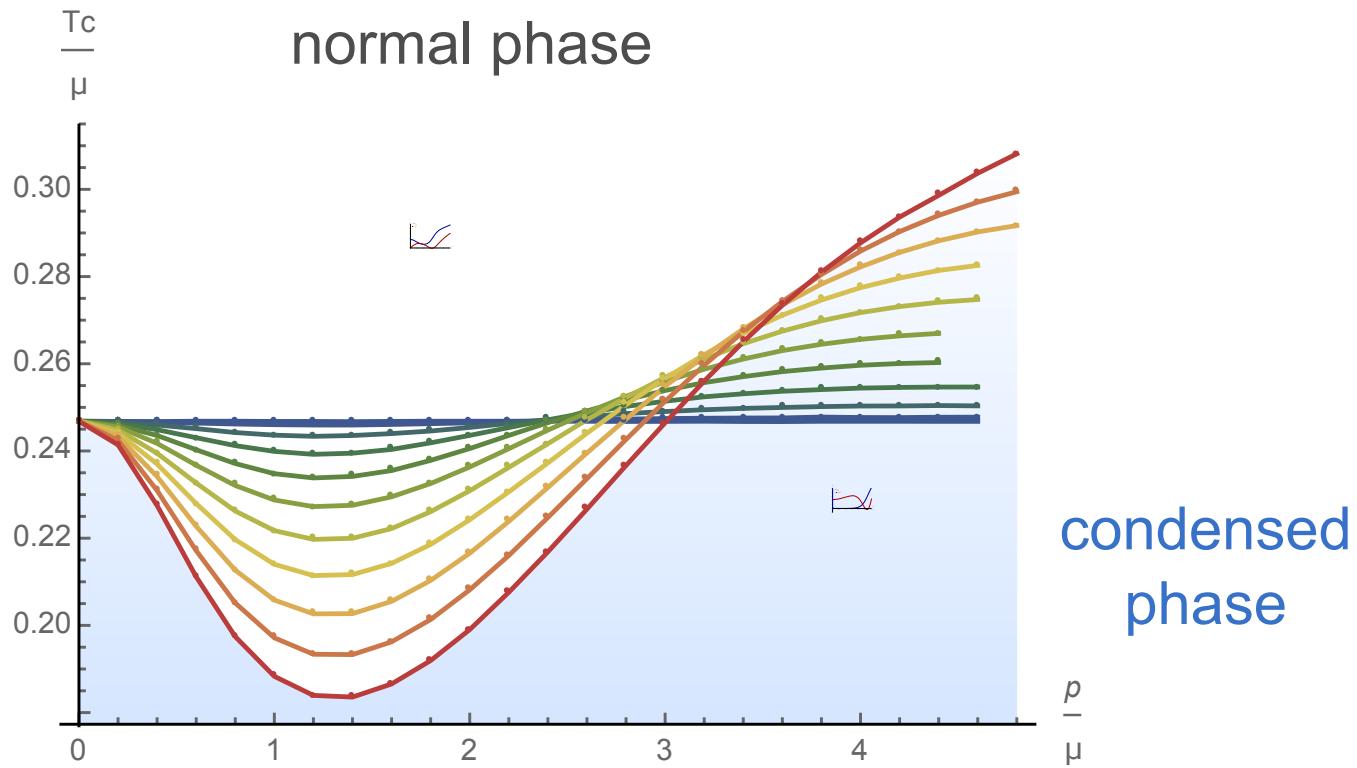
$$\mathcal{A} = A_0^b + \frac{A_2^b + A_0^b \omega^2 \log(r)/2}{r^2} + \dots$$

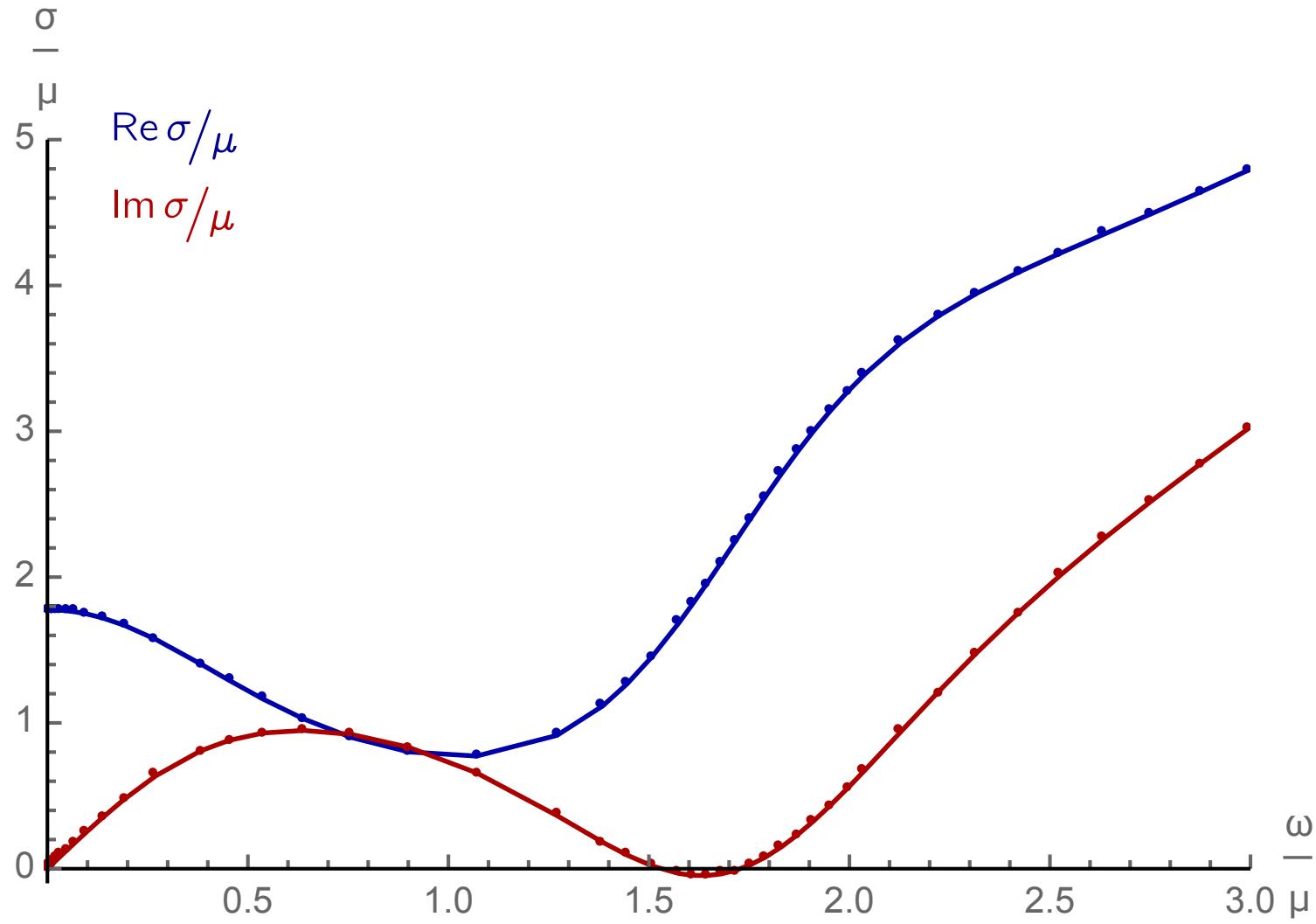
- Photon propagator at zero momentum

$$\sigma_x(\omega) = \lim_{\mathbf{k} \rightarrow 0} \frac{G_{xx}^R(\omega, \mathbf{k})}{i\omega} \quad G_{xx}^R(\omega, 0) = 2 \left(\frac{A_2^b(\omega)}{A_0^b(\omega)} - \frac{\omega^2}{4} \right)$$

Optical Conductivity

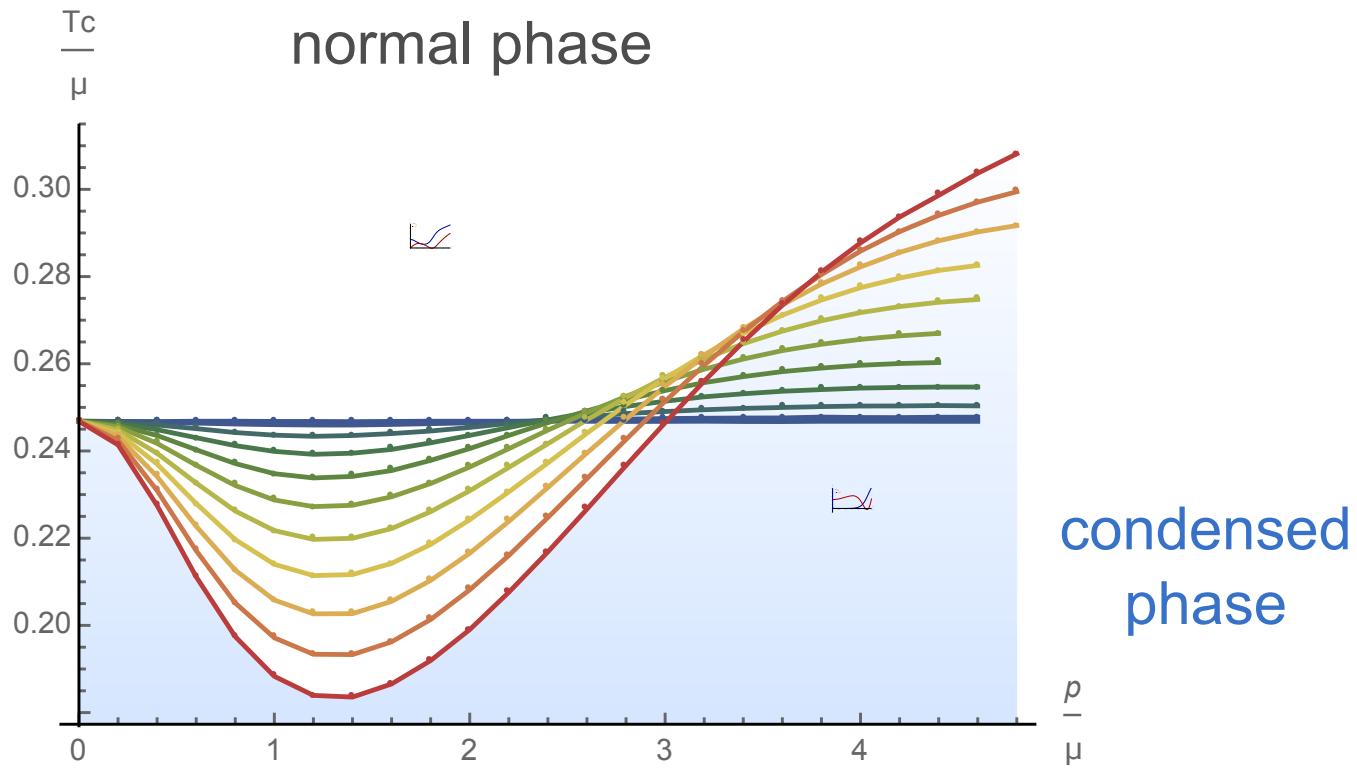
- Optical conductivity in normal and condensed phase

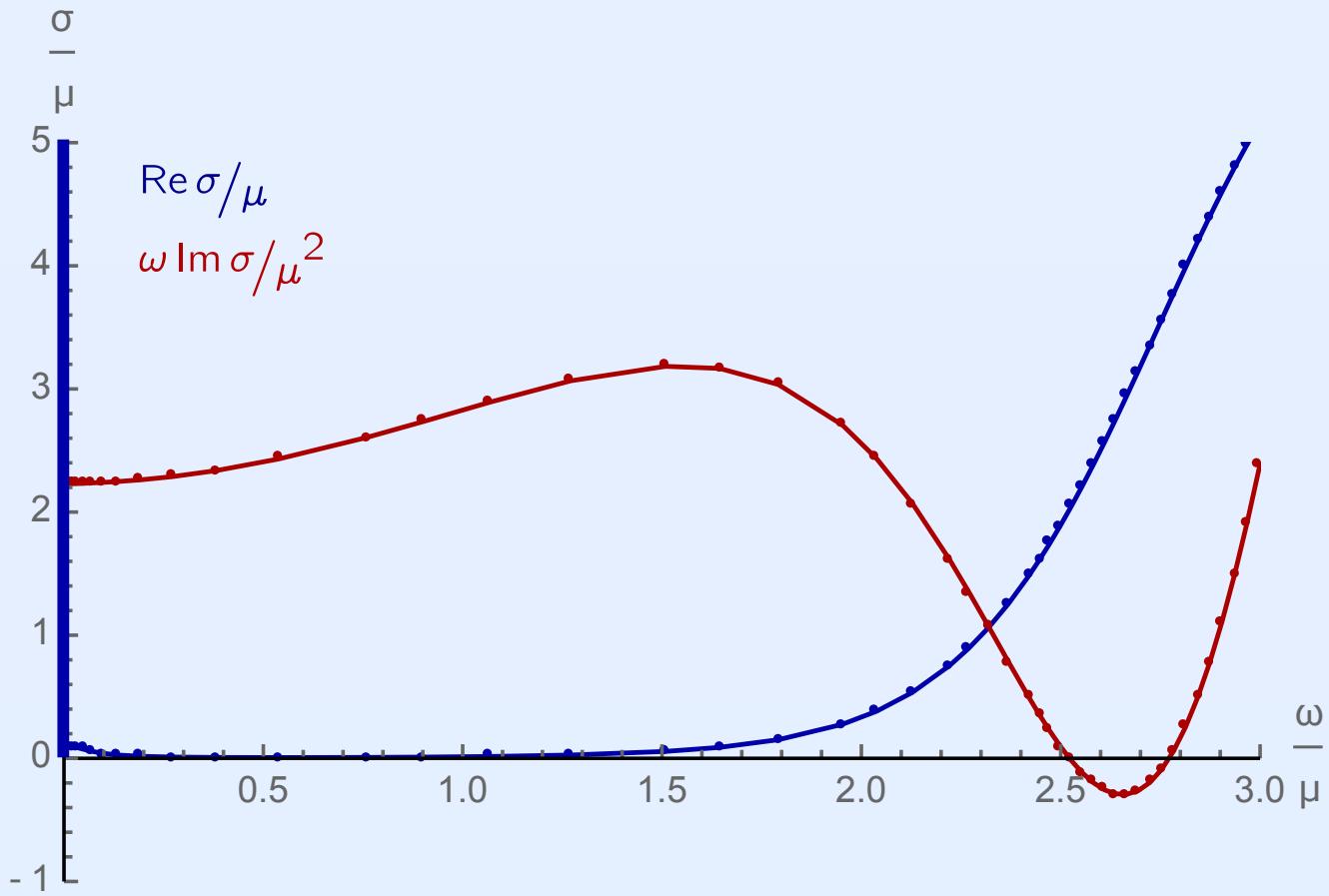




Optical Conductivity

- Optical conductivity in normal and condensed phase

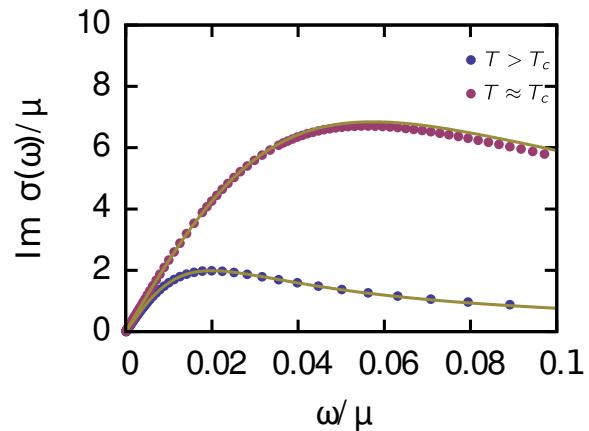
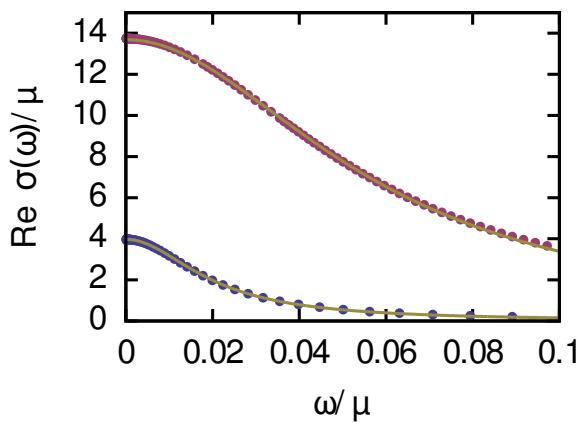




Drude Model

► Drude model

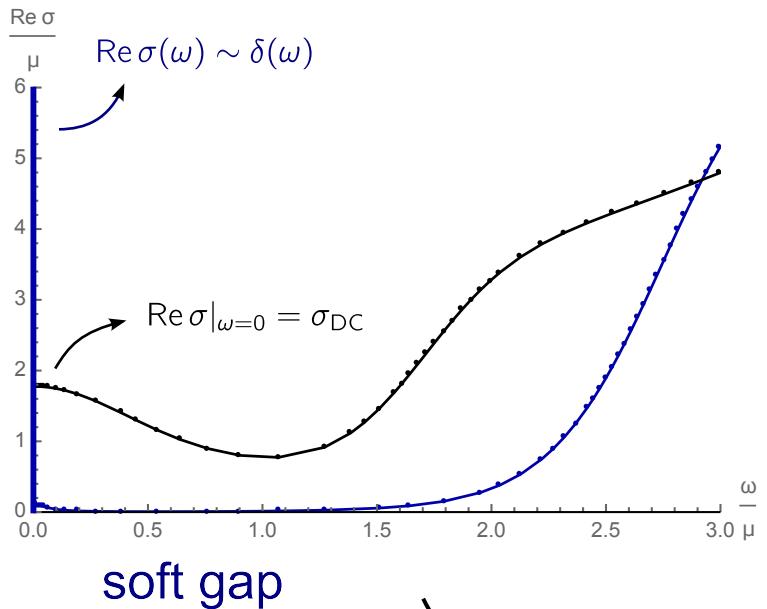
$$\sigma(\omega) = \frac{\sigma_{DC}}{1 - i\omega\tau} = \frac{n_n e^2 \tau}{m^*} \frac{1}{1 - i\omega\tau} = \frac{\omega_P^2}{4\pi} \frac{1}{1/\tau - i\omega}$$



→ $\text{Re } \sigma(\omega) = \frac{\sigma_{DC}}{1 + \omega^2 \tau^2}$

→ $\text{Im } \sigma(\omega) = \frac{\omega \tau}{1 + \omega^2 \tau^2} \sigma_{DC}$

Optical Conductivity Superfluid/Insulator

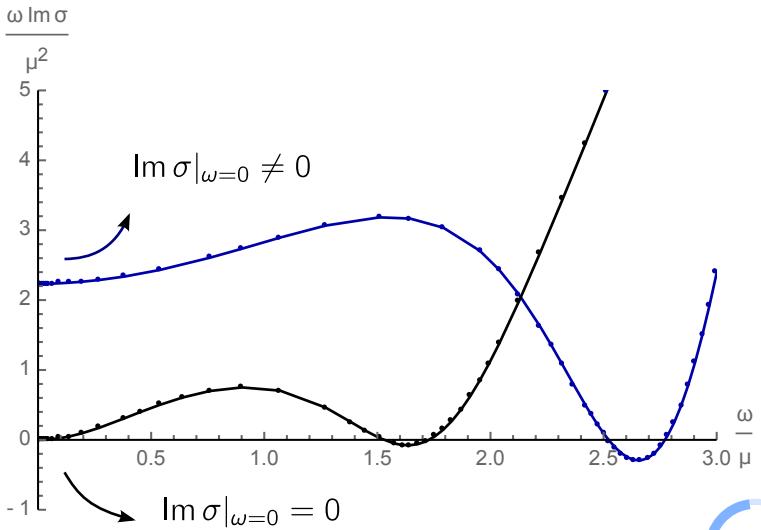


$$\text{Re } \sigma(\omega) \sim \delta(\omega) \Leftrightarrow \text{Im } \sigma(\omega) \sim \frac{i}{2} \frac{1}{\omega}$$

Kramers-Kronig relations

$$\text{Re } f(\omega) = \frac{2}{\pi} \mathcal{P} \int_0^{\infty} d\omega' \frac{\omega' \text{Im } f(\omega')}{\omega'^2 - \omega^2}$$

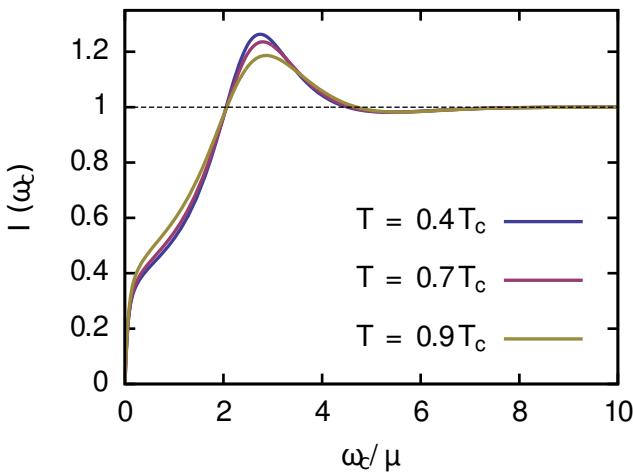
$$\text{Im } f(\omega) = -\frac{2\omega}{\pi} \mathcal{P} \int_0^{\infty} d\omega' \frac{\text{Re } f(\omega')}{\omega'^2 - \omega^2}$$



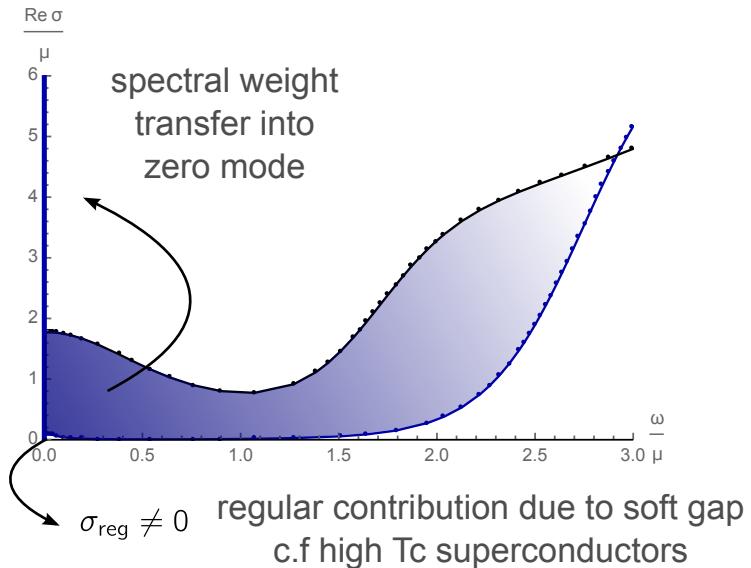
F-Sum Rule

- Plasma frequency
~ Superfluid strength

$$\begin{aligned}\omega_{Ps}^2 &= 8 \int_0^\infty d\omega \operatorname{Re} \sigma_s(\omega) \\ &= 4\pi \frac{n_s e^2}{m} = \lambda_L^{-2} \equiv \rho_s\end{aligned}$$



- Ferrell-Glover-Tinkham sum rule



$$\rho_s = \int_{0^+}^{\infty} d\omega [\operatorname{Re} \sigma_n(\omega) - \operatorname{Re} \sigma_s(\omega)]$$

assuming no residual regular or high frequency contribution

Two Fluid Model

► Two fluid model

$$\text{Re } \sigma(\omega) = \frac{e}{m^*} \left(\chi_n(T) \frac{\tau}{1 + \omega^2 \tau^2} + \frac{\pi}{2} \chi_s(T) \delta(\omega) \right)$$

regular metallic part superfluid part

► χ is controlling the strength

normal phase:

$$\chi_n(T > T_c) = n_n$$

$$\chi_s(T > T_c) = 0$$

$$\rho_s(T \rightarrow 0) \approx n_s(T \rightarrow 0)$$

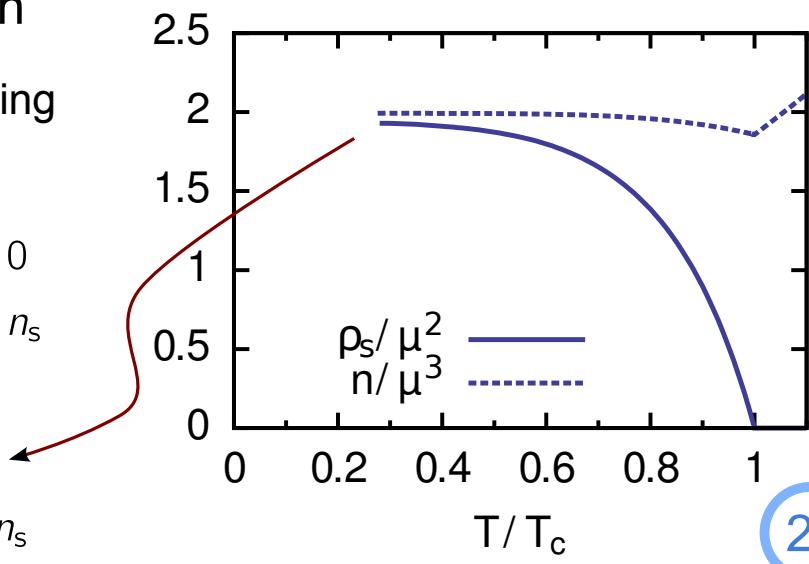
superconducting phase:

$$\chi_n(T = 0) = 0$$

$$\chi_s(T = 0) = n_s$$

$$\chi_n(T \rightarrow 0) \approx 0$$

$$\chi_s(T \rightarrow 0) \approx n_s$$



Part IV:

Homes' & Uemura's Relation

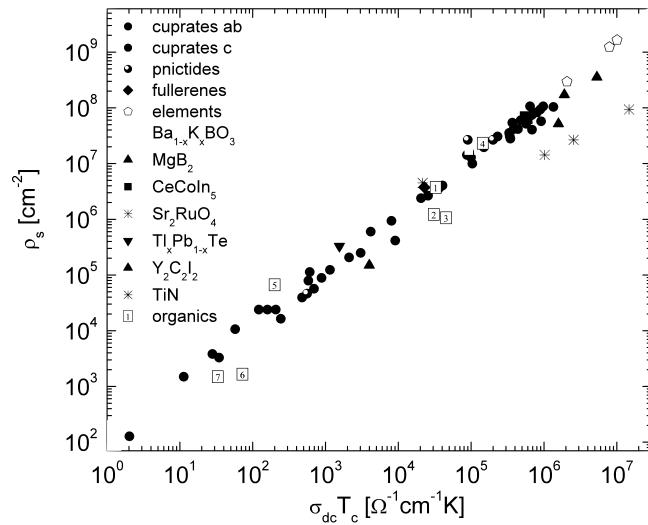
Homes Relation

► Homes' relation

$$\rho_s = C \sigma_{DC}(T_c) T_c$$

Insert Drude model:

$$\rho_s = C \frac{e}{m^*} n_n(T_c) \tau(T_c) T_c$$



[Dordevic, Basov, Homes '12]

► Perfect fluids / "Planckian Dissipators"

$$\tau_\hbar(T) \sim \frac{\hbar}{k_B T} \quad \quad \frac{\eta}{T_S} = \frac{1}{4\pi} \frac{\hbar}{k_B T} = \frac{\tau_\hbar}{4\pi}$$

holographic superconductors
are
Planckian dissipators

Homes Relation

► Holographic Version of Homes' Relation

$$\rho_s = C \frac{e}{m^*} n_n(T_c) \tau(T_c) T_c$$

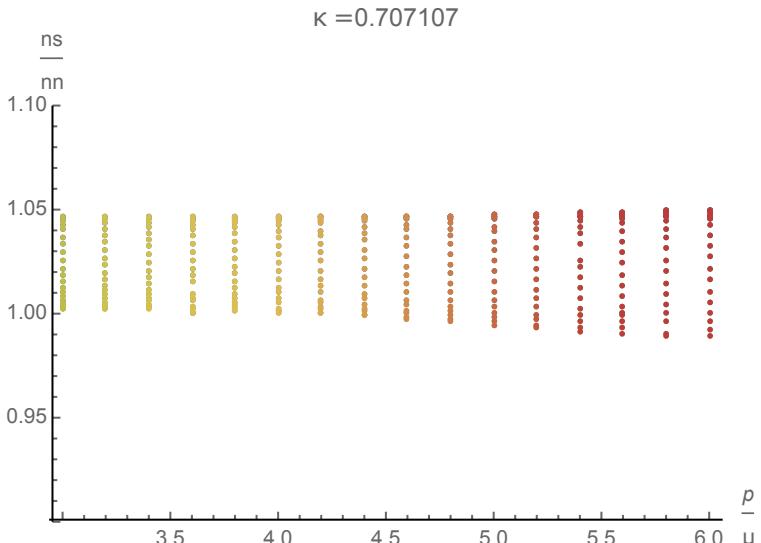
all charge carrier condense
in zero mode

$$\rho_s(T \rightarrow 0) \approx n_s(T \rightarrow 0)$$

quantum critical system

$$\tau(T_c) T_c \sim 1$$

$$\Rightarrow n_s(T \approx 0) = C' \frac{e}{m^*} n_n(T_c) \quad \Rightarrow \quad \frac{n_s(T \approx 0)}{n_n(T_c)} = \text{const.}$$



Part V:

Zero Temperature Solutions

Zero Temperature Solution

- ▶ Scaling solution in the deep IR (no horizon!)

$$a_0 = \frac{9\kappa|p|p e^{-2\nu_{10}}}{5(6\kappa^2 + q^2\rho_0^2 - 4)}$$

$$\rho_1 = -\left(\frac{|p|}{2}\right)^4 \frac{\kappa^2 q^2 \rho_0 e^{-4\nu_{10}}}{(6\kappa^2 + q^2\rho_0^2 - 4)^2}$$

$$e^{\nu_{30}} = \frac{2}{|p|} e^{\nu_{10} + \nu_{20}}$$

$$w_0 = \sqrt{3} \left(\frac{|p|}{2}\right)^{-2} e^{2\nu_{10} + \nu_{20}}$$

$$w_1 = \left(\frac{|p|}{2}\right)^2 \frac{\sqrt{3}(q^2\rho_0^2 - 4)}{2(6\kappa^2 + q^2\rho_0^2 - 4)} e^{\nu_{20} - 2\nu_{10}}$$

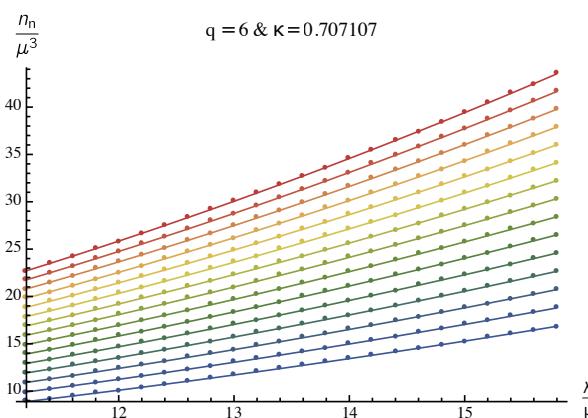
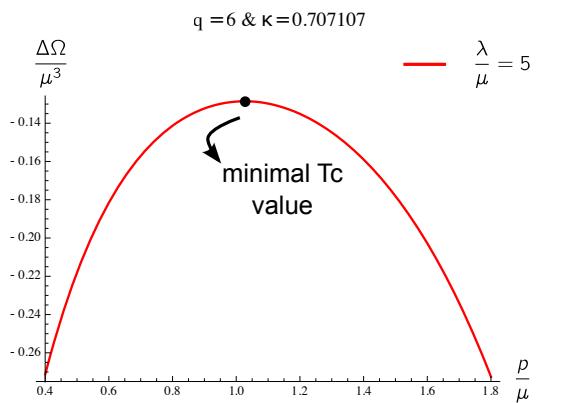
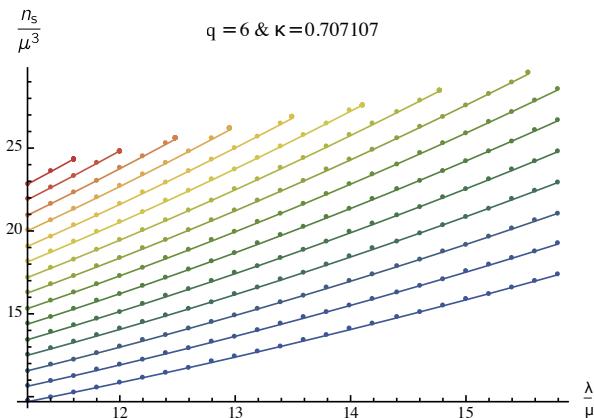
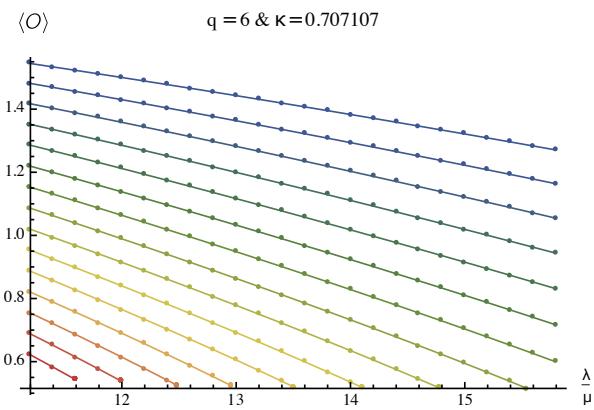
- ▶ Expansion about IR geometry

$$w = w_0 + w_1 r^{4/3}(1 + c_w r^\delta) \quad \rho = \rho_0 + \rho_1 r^{4/3}(1 + c_\rho r^\delta) \quad a = a_0 r^{5/4}(1 + c_a r^\delta)$$

$$\nu_1 = \nu_{10} + \log(r^{-1/3}) + c_1 r^\delta \quad \nu_2 = \nu_{20} + \log(r^{2/3}) + c_2 r^\delta \quad U = \frac{18}{5} r^2(1 + c_U r^\delta)$$

$$\nu_3 = \nu_{30} + \log(r^{1/3}) + c_3 r^\delta$$

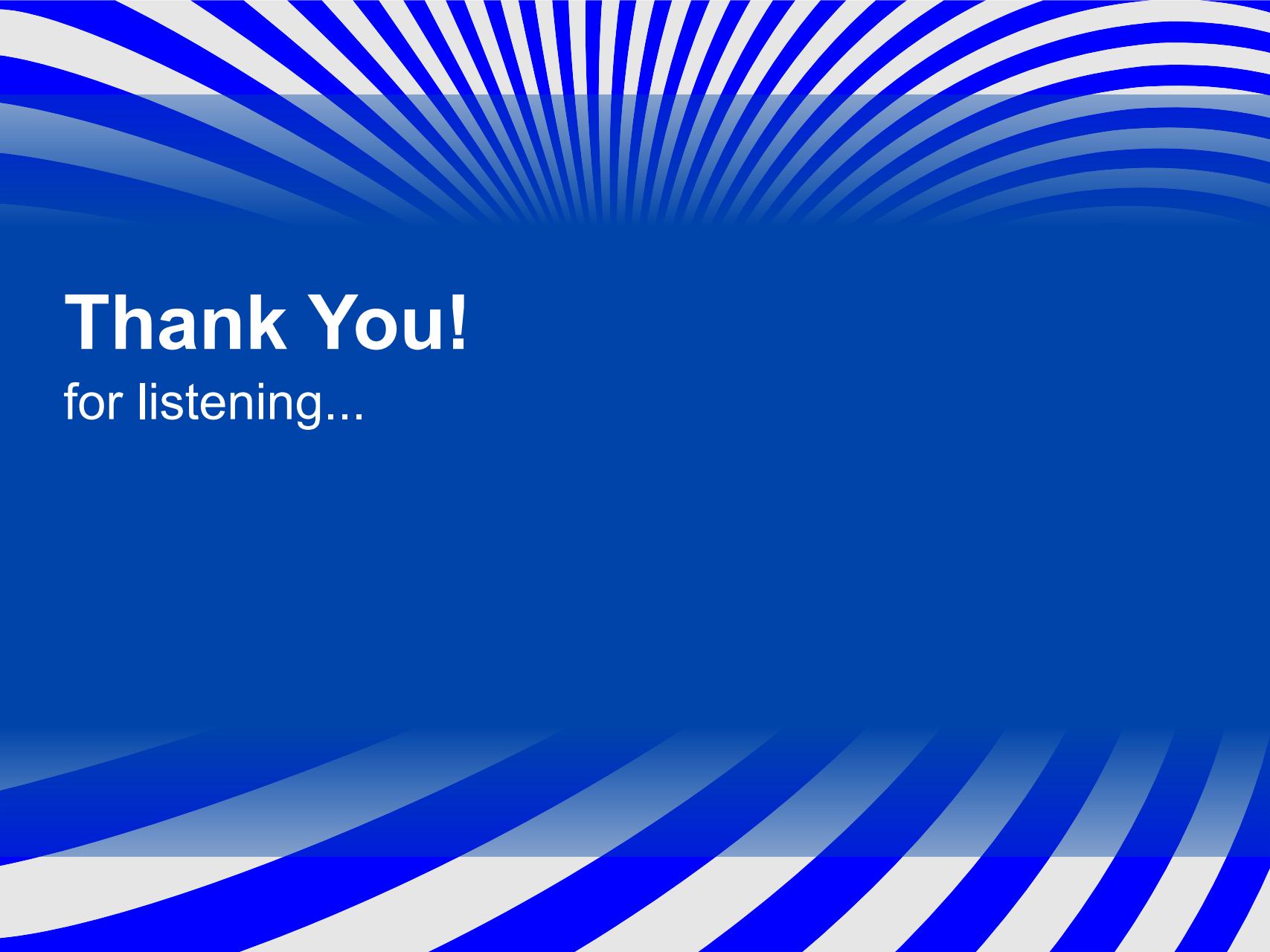
Zero Temperature Thermodynamics



- \bullet $\frac{p}{\mu} = 0.17$ \bullet $\frac{p}{\mu} = 0.33$
- \bullet $\frac{p}{\mu} = 0.19$ \bullet $\frac{p}{\mu} = 0.35$
- \bullet $\frac{p}{\mu} = 0.21$ \bullet $\frac{p}{\mu} = 0.37$
- \bullet $\frac{p}{\mu} = 0.23$ \bullet $\frac{p}{\mu} = 0.39$
- \bullet $\frac{p}{\mu} = 0.25$ \bullet $\frac{p}{\mu} = 0.41$
- \bullet $\frac{p}{\mu} = 0.27$ \bullet $\frac{p}{\mu} = 0.43$
- \bullet $\frac{p}{\mu} = 0.29$ \bullet $\frac{p}{\mu} = 0.45$
- \bullet $\frac{p}{\mu} = 0.31$

Outlook

- ▶ Check other relations such as Uemura's or Tanner's relation
- ▶ Scaling behavior of optical conductivity with temperature and frequency
- ▶ Zero temperature solutions deserve closer investigation
 - ↳ Determine zero temperature quantities e.g. ρ , n_s , σ , ...
 - ↳ Compute zero temperature normal phase diagram

The background of the slide features a dynamic, abstract design. It consists of several layers of curved, wavy lines in varying shades of blue, creating a sense of depth and motion. In the center, there is a cluster of straight, radial lines that fan outwards, also in shades of blue. The overall effect is reminiscent of a rising sun or a stylized landscape.

Thank You!
for listening...