

# Black Funnel and Droplets

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V. Hubeny, D. Marolf, M. Rangamani, arXiv:0908.2270  
J. Santos, B.W, arXiv:1207.4205  
J. Santos, B.W, arXiv:1405.2078

# AdS/CFT with Curved Boundary

To find gravity duals to holographic CFTs, we solve Einstein's Equations with negative cosmological constant

$$S = \int d^{d+1}x \sqrt{-g} (R - 2\Lambda)$$

Solutions have a boundary metric. Typically, this is

- Minkowski space (Poincaré AdS)
- Sphere (Global AdS)

We will take the boundary metric to contain a black hole (e.g. Schwarzschild).

# Holographic CFTs on Curved Spacetime

## Hawking Radiation

- Hawking radiation is usually discussed in the context of perturbative fields.
- What are the differences when the fields are strongly coupled?

# Holographic CFTs on Curved Spacetime

## Hawking Radiation

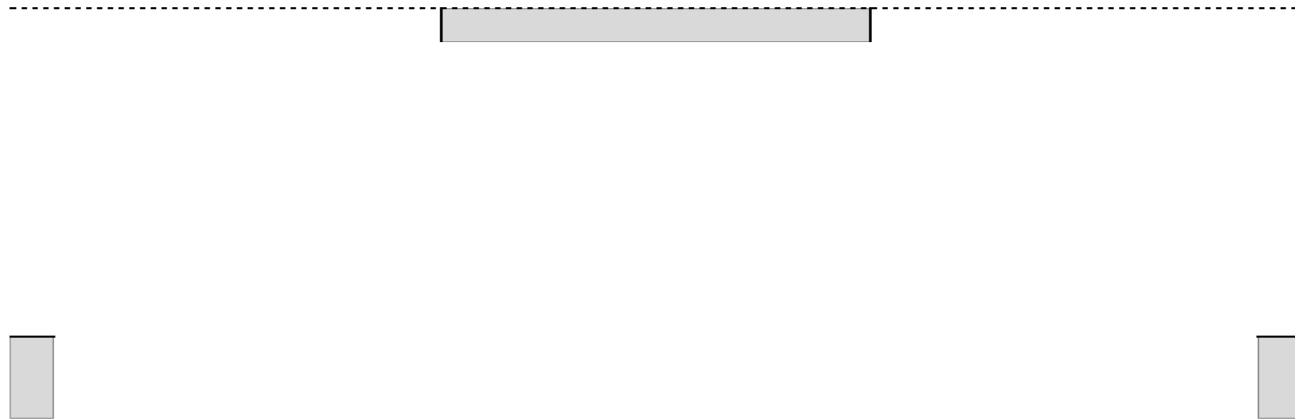
- Hawking radiation is usually discussed in the context of perturbative fields.
- What are the differences when the fields are strongly coupled?

## Heat Transport

- Black holes can play the role of a heat bath.
- Provides a natural setting to study heat transport.

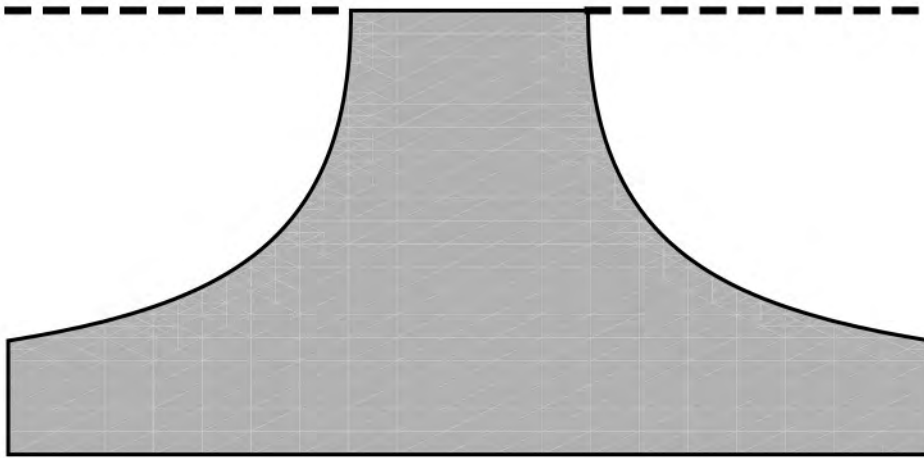
# AdS/CFT with Curved Boundary

Black hole metric (e.g. Schwarzschild) on boundary with finite temperature at infinity.



# Gravity Dual

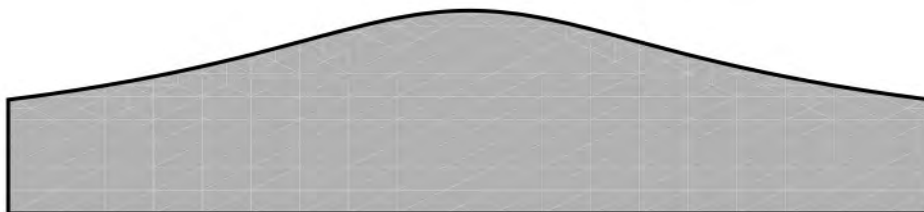
Horizons are either connected or disconnected.



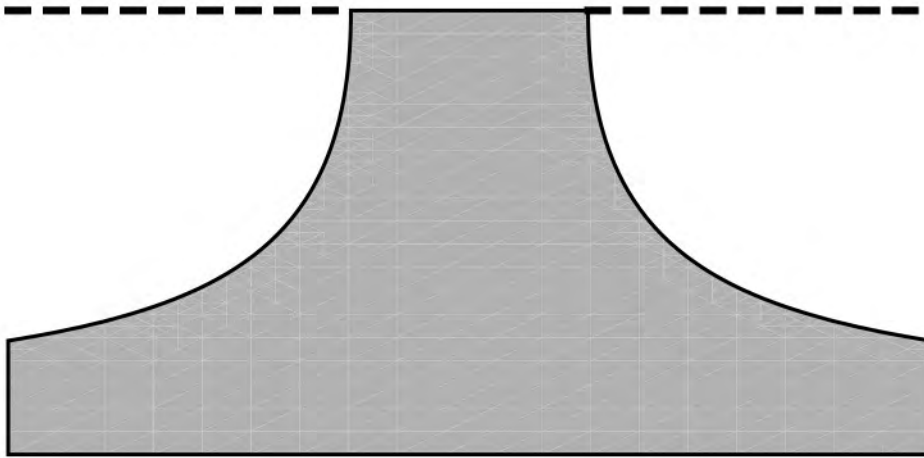
(Flowing) Black Funnel



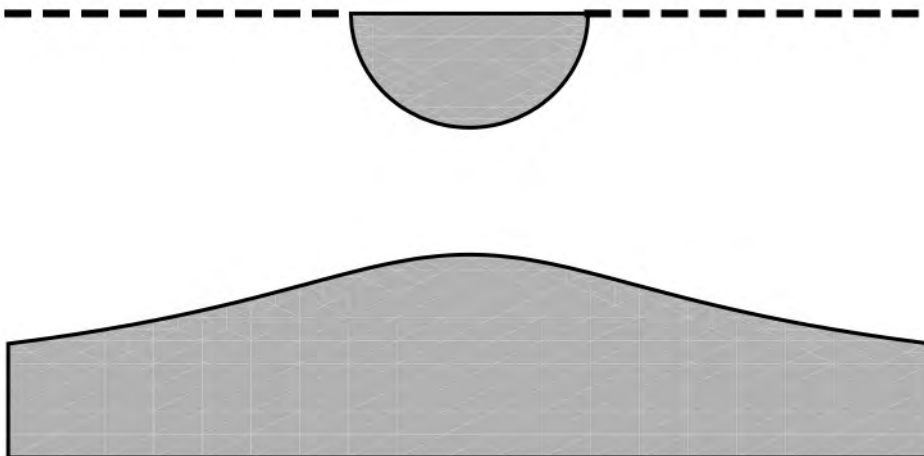
Black Droplet  
(Over Planar Black hole)



# Gravity Dual



Heat easily exchanged  
with infinity



Heat NOT easily  
exchanged with infinity

# Droplet Field Theory Interpretation

**From the field theory perspective, why are droplets static?**

One possible answer:

- Suppose CFT far from black hole admits a useful quasiparticle description, and these quasiparticles have some preferred size  $R_{\text{quasi}} \sim 1/T_{\infty}$ .
- Black holes of size  $R_{\text{BH}} \ll R_{\text{quasi}}$  might have trouble absorbing/emitting such quasiparticles.



# Droplet Field Theory Interpretation

**From the field theory perspective, why are droplets static?**

Another possible answer:

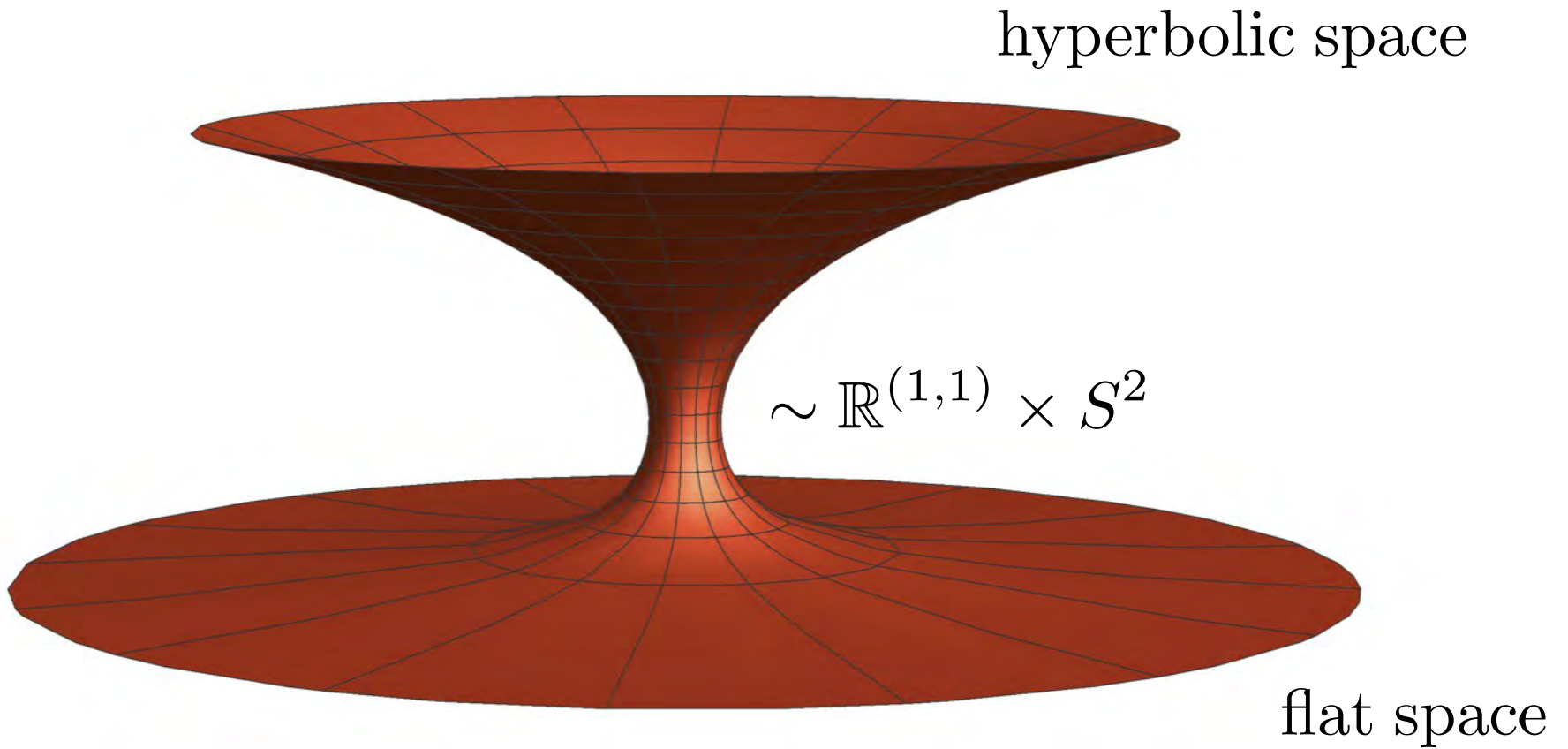
- Under a conformal transformation  $ds^2 \rightarrow ds^2 / f(r)$ , the boundary black hole horizon becomes a hyperbolic region.

$$-f dt^2 + \frac{dr^2}{f} + r^2 d\Omega$$

$$\rightarrow -dt^2 + \frac{dr^2}{f^2} + \frac{r^2}{f} d\Omega$$

$$\sim -dt^2 + \frac{4}{|f'|^2} \left( \frac{dz^2}{z^2} + \frac{d\Omega}{z^2} \right) + \dots$$

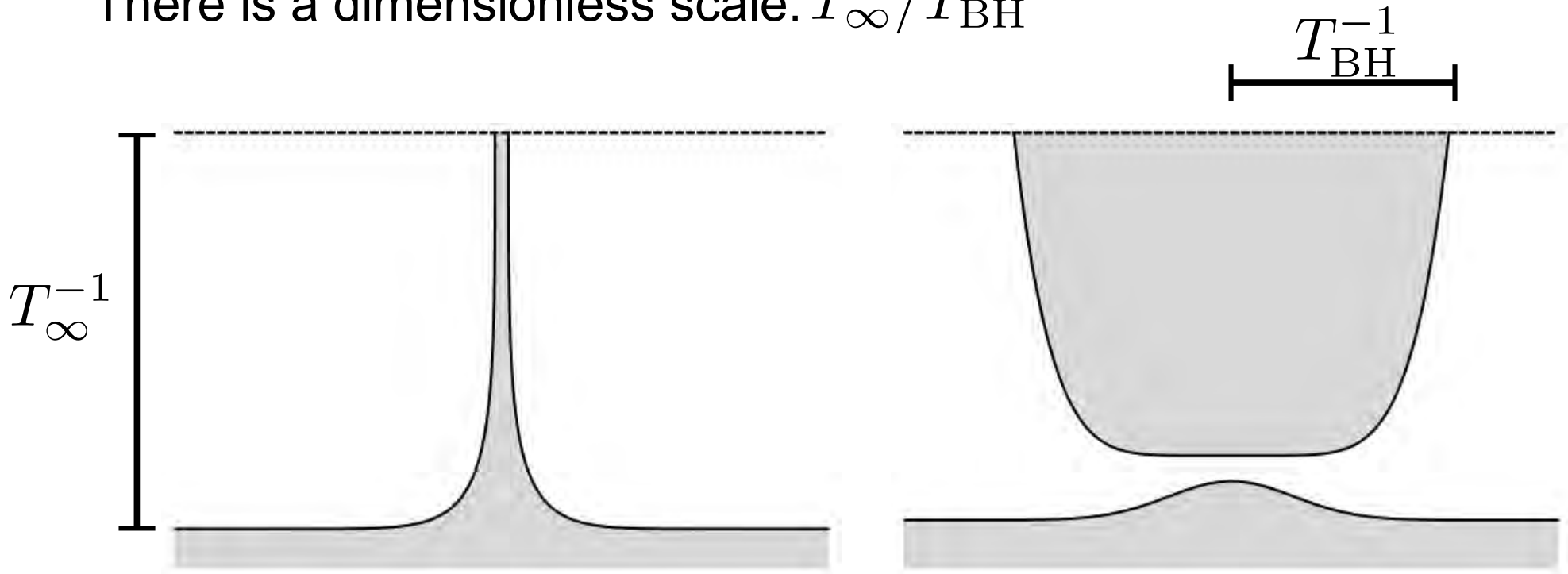
# Ultrastatic Frame



Confining region prevents interaction between heat baths.

# Phase Transition

There is a dimensionless scale:  $T_\infty/T_{\text{BH}}$



- These seem like 'unnatural' configurations.
- Conjecture: Funnel solutions preferred for  $T_\infty/T_{\text{BH}} \gg 1$ , and droplet solutions preferred for  $T_\infty/T_{\text{BH}} \ll 1$ .
- Cone Transition?

# Static Solutions

For simplicity, we will focus on static solutions:

- Boundary metric is asymptotically flat Schwarzschild.
- Funnels have  $T_\infty = T_{\text{BH}}$ .
- Droplets can have  $T_\infty \neq T_{\text{BH}}$ .
- No analytic solutions, must rely on numerics.

# The DeTurck Method

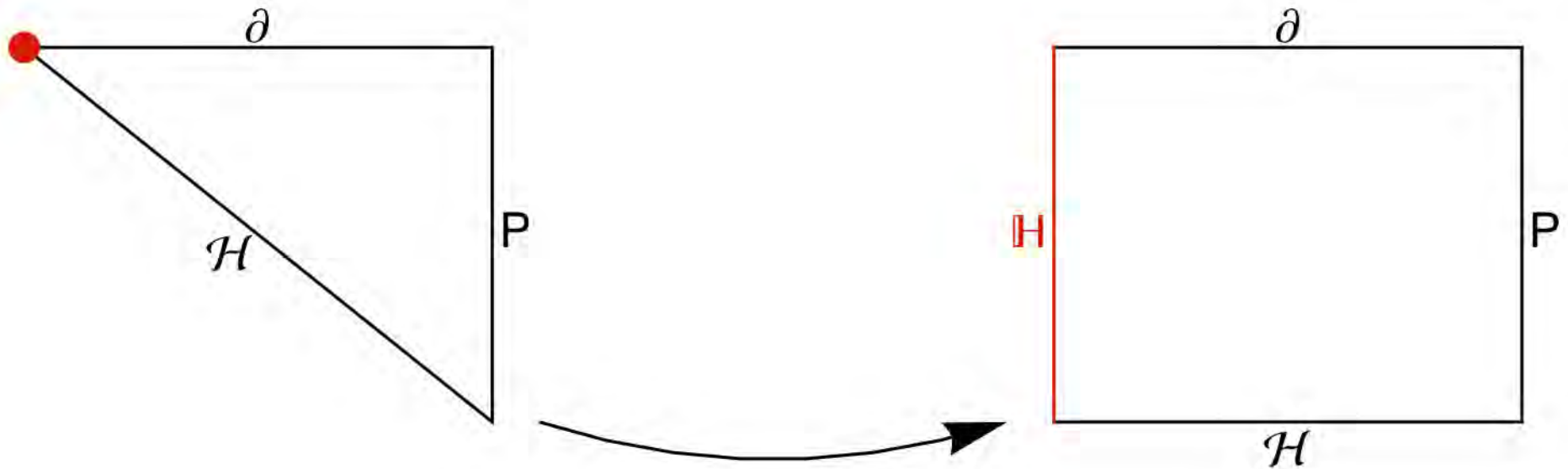
Instead of solving the usual Einstein's equations, solve the Einstein-DeTurck equations:

$$R_{\mu\nu} = -\frac{d}{\ell^2}g_{\mu\nu} + \nabla_{(\mu}\xi_{\nu)} \quad \xi^\mu = g^{\alpha\beta} \left( \Gamma_{\alpha\beta}^\mu - \bar{\Gamma}_{\alpha\beta}^\mu \right)$$

- Only a solution to Einstein's equations when  $\xi^\mu = 0$ .
- Can prove that solutions with  $\xi^\mu \neq 0$  do not exist (in our case).
- Do not need to fix a gauge a priori. Solving the equations will give solution in the gauge  $\xi^\mu = 0$ .
- Equations are elliptic.
- Quantity  $\chi \equiv \xi^\mu \xi_\mu$  can be used to monitor numerical error.

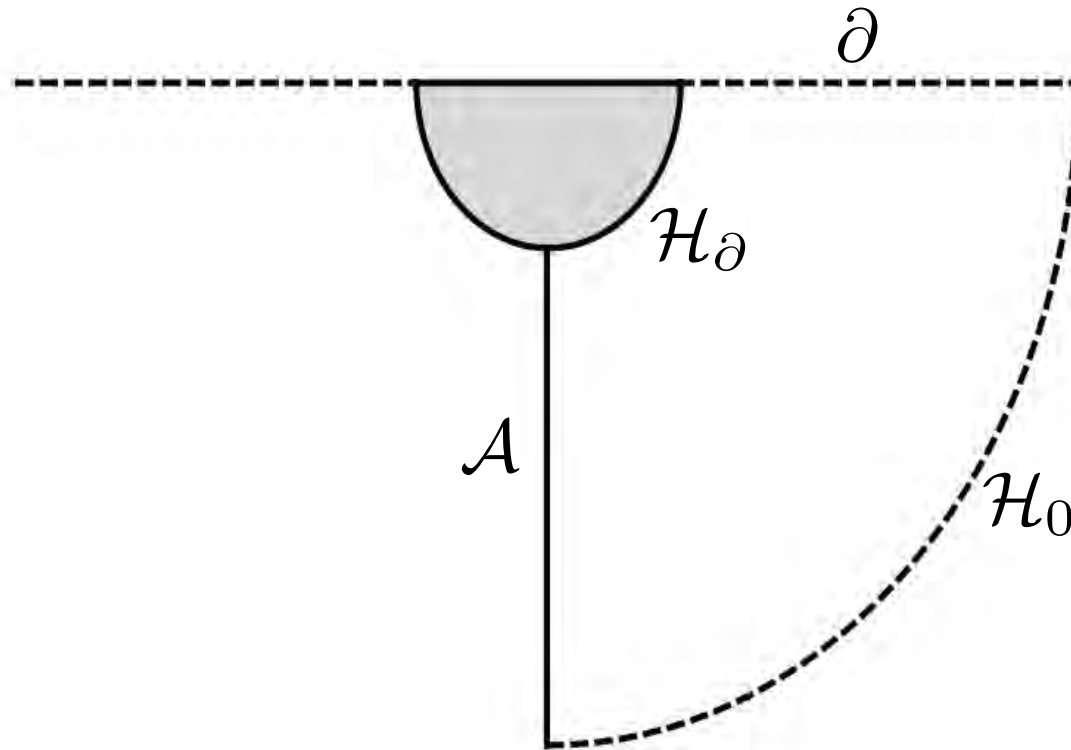
# Funnel Integration Domain

- Natural integration domain is a triangle.
- Expand point where boundary meets the horizon.
- Point becomes a hyperbolic black hole.



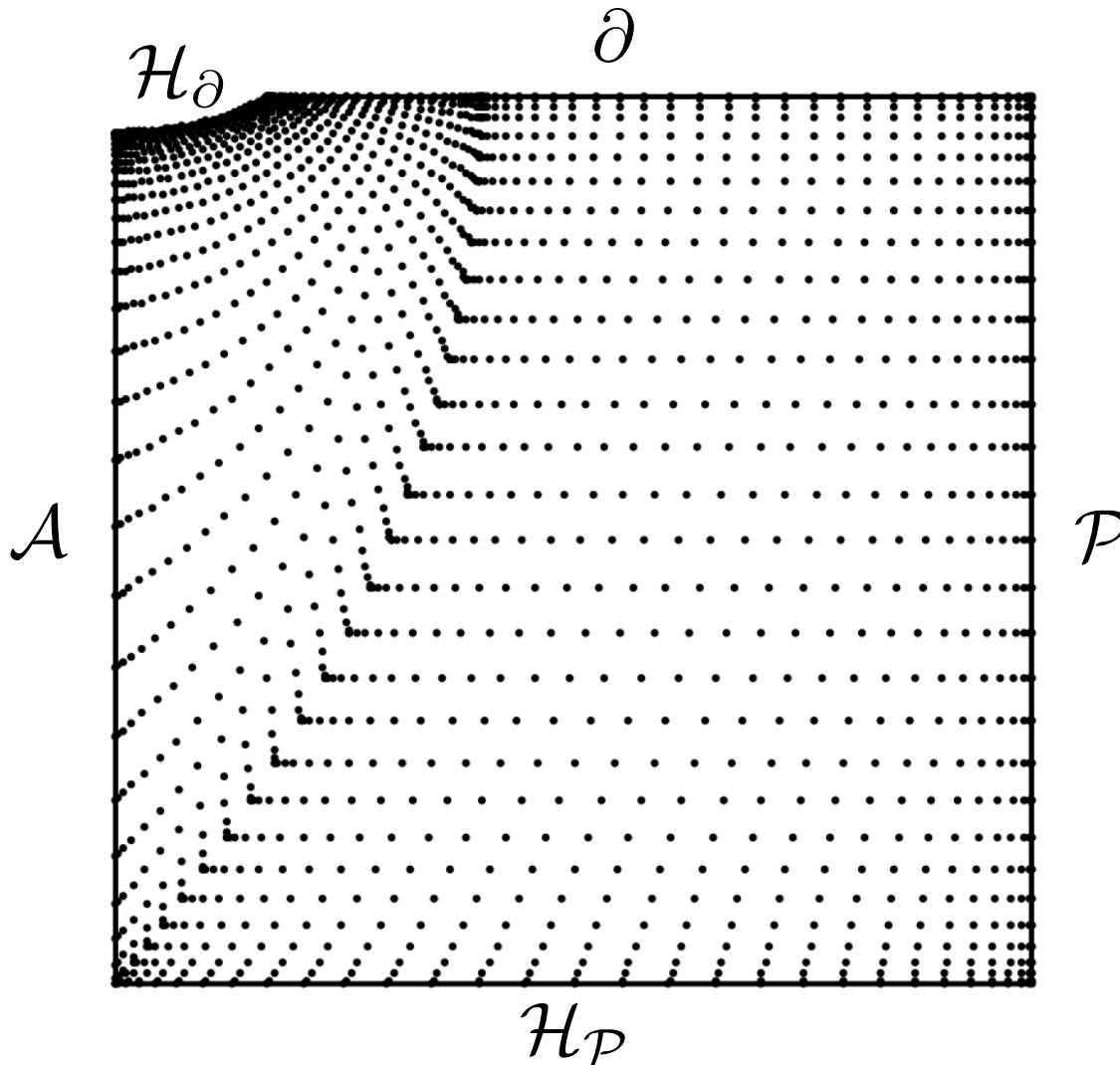
# Droplet Integration Domains

Droplets with  $T_\infty = 0$  fit in 'polar' coordinates.



# Droplet Integration Domains

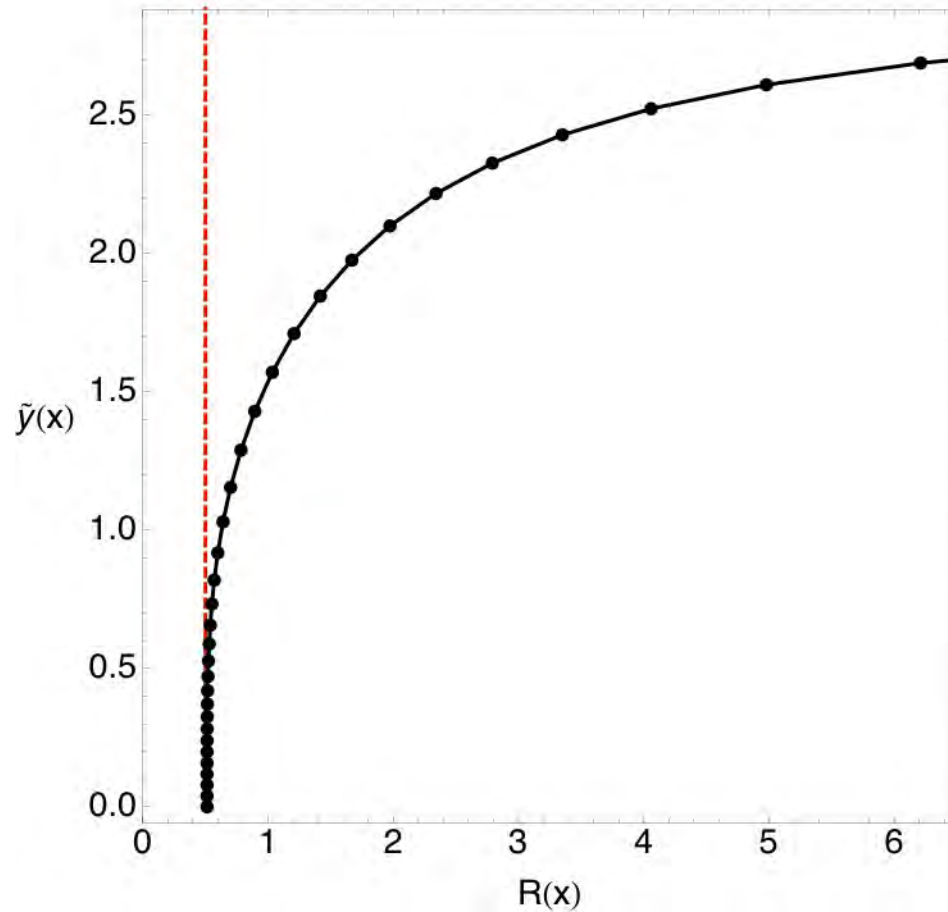
Use patching with transfinite interpolation for droplets with planar black hole ( $T_\infty \neq 0$ ) .





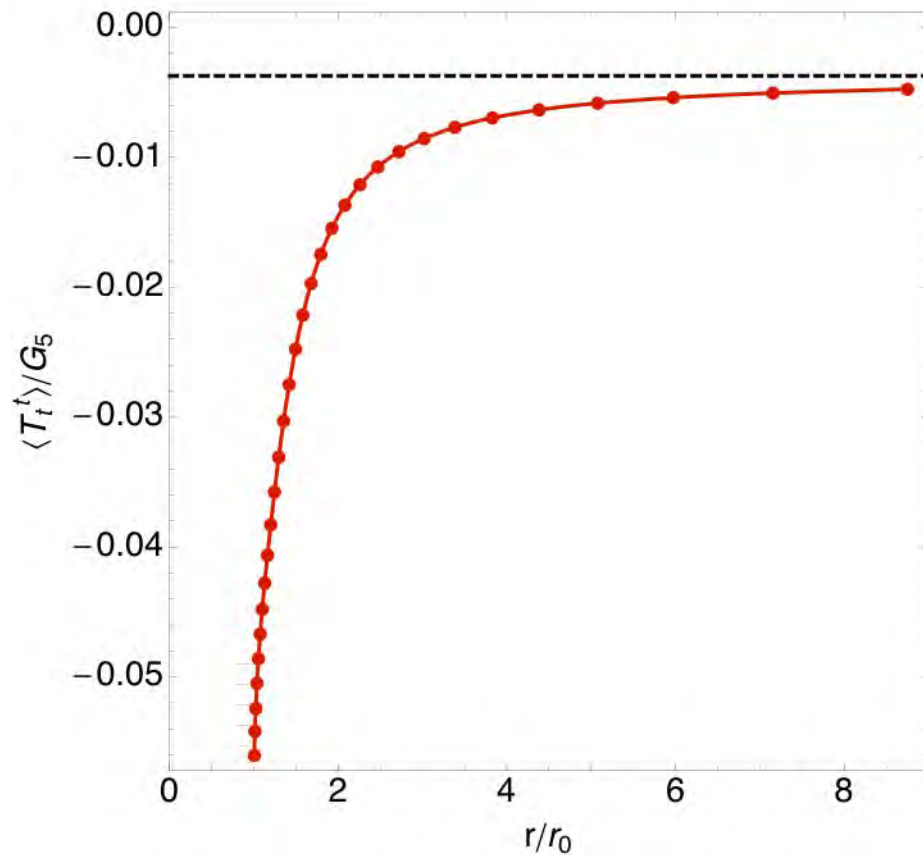
# Funnel Embedding

Embed the funnel horizon in hyperbolic space.



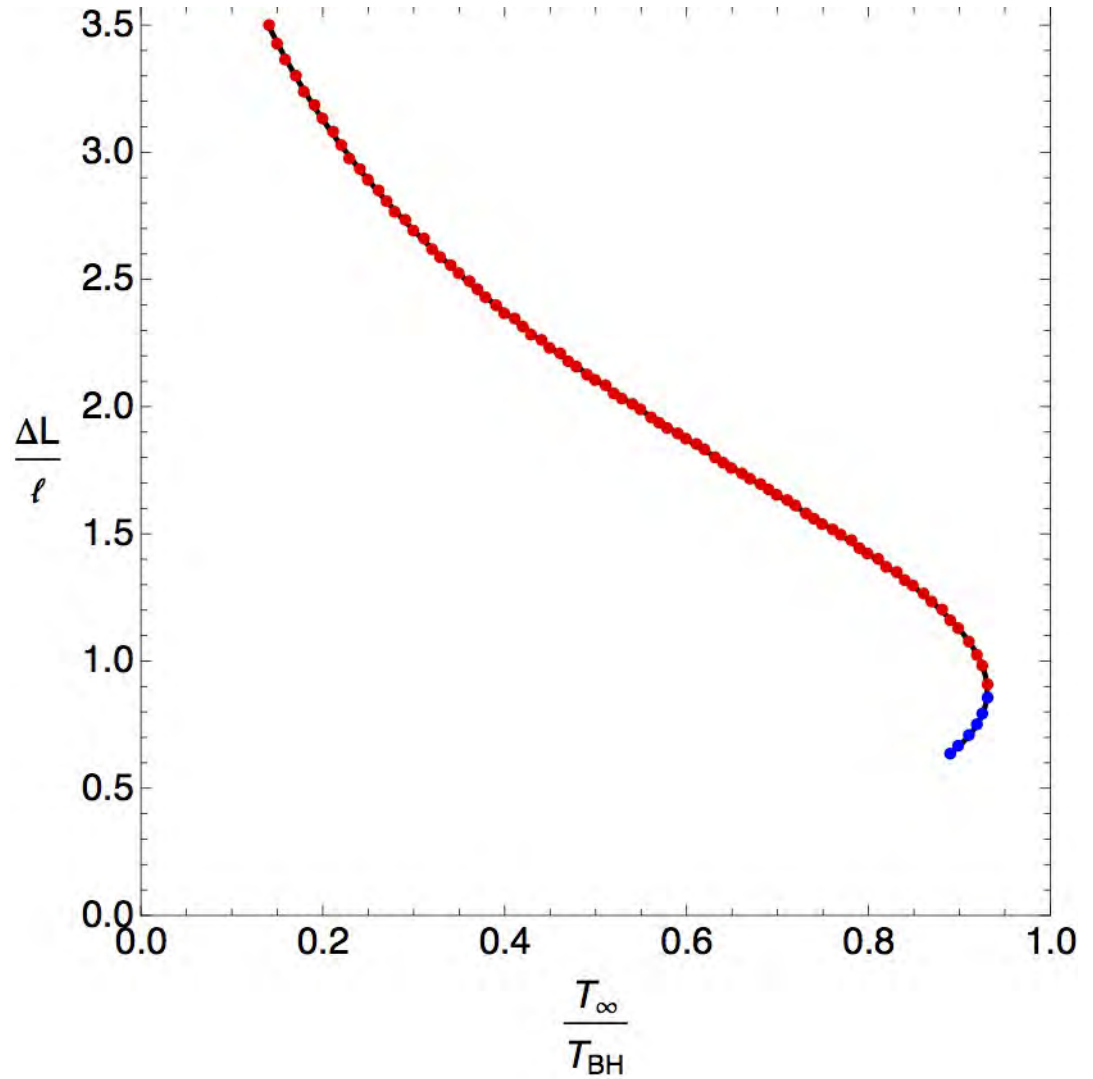
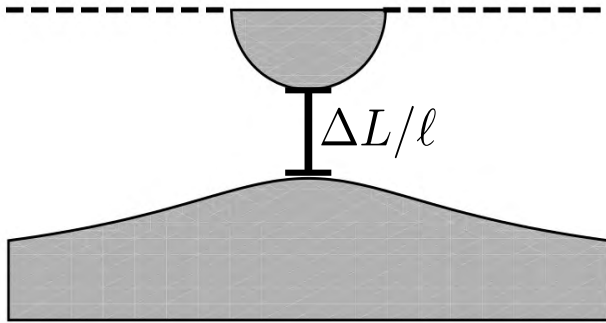
# Funnel Stress Tensor

Compute the stress tensor by expanding off the boundary in Fefferman-Graham coordinates.

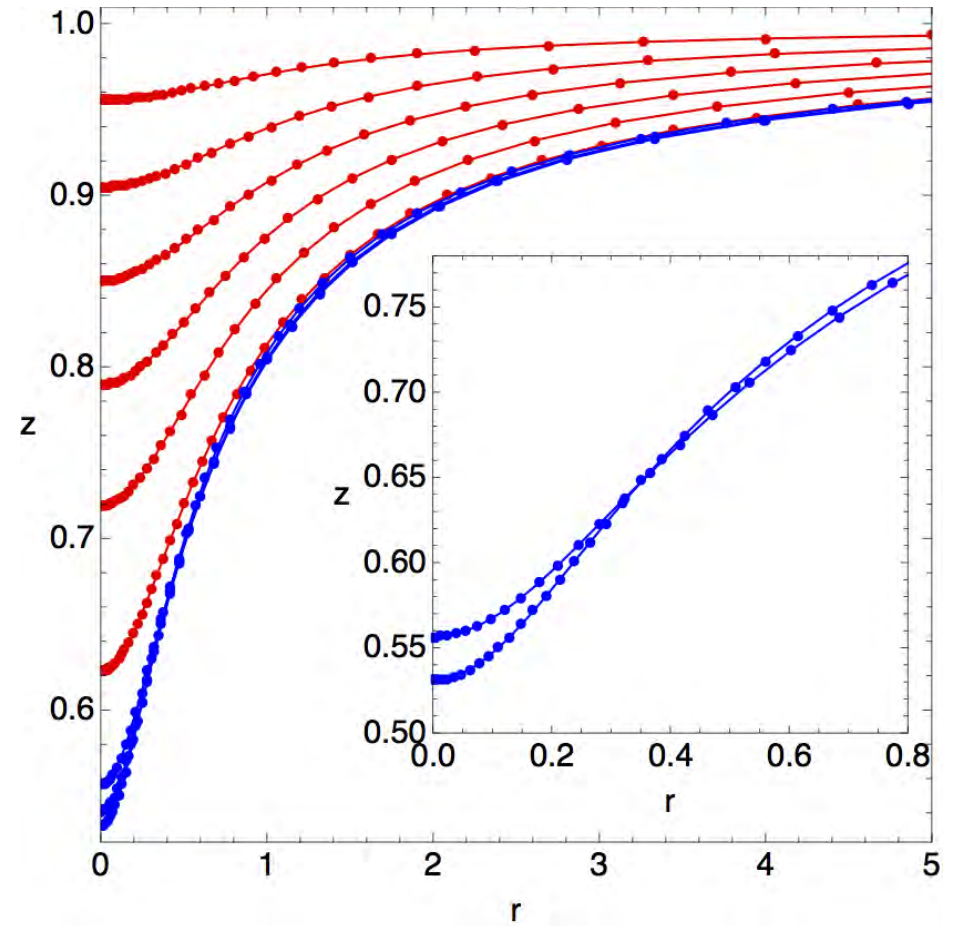
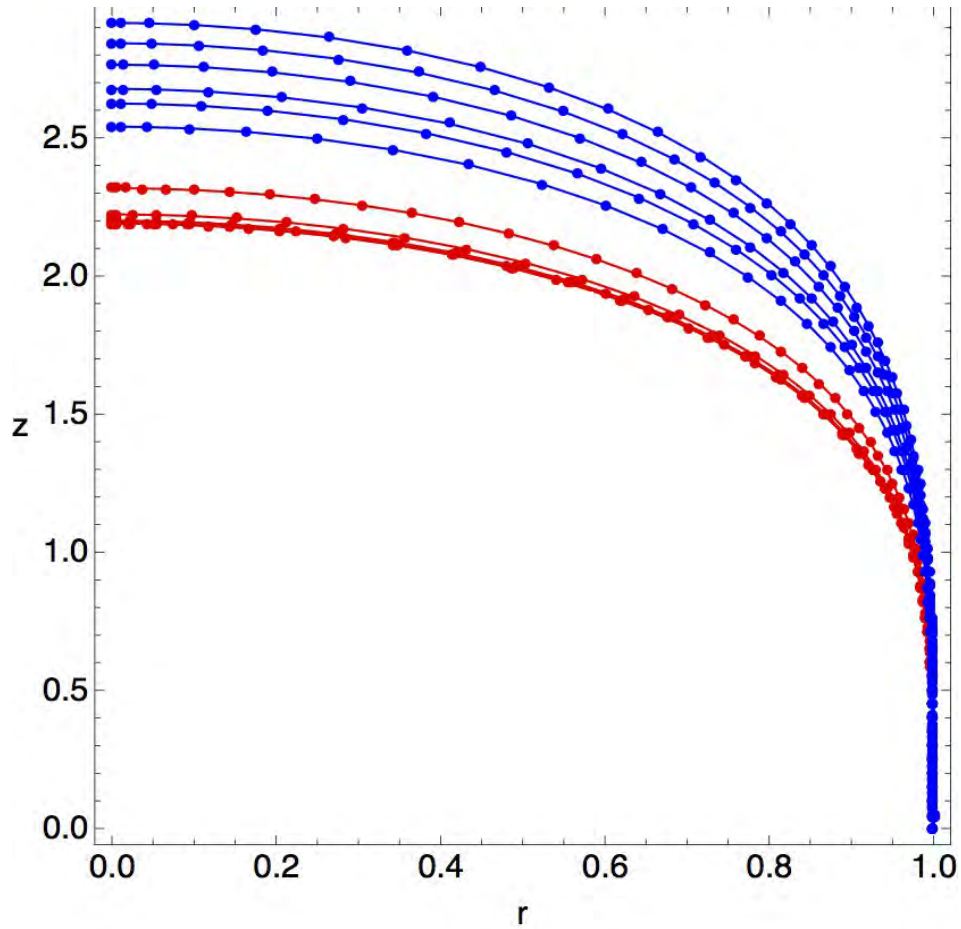


Energy-density positive, stress tensor falls off as  $1/r$ .

# Proper Distance

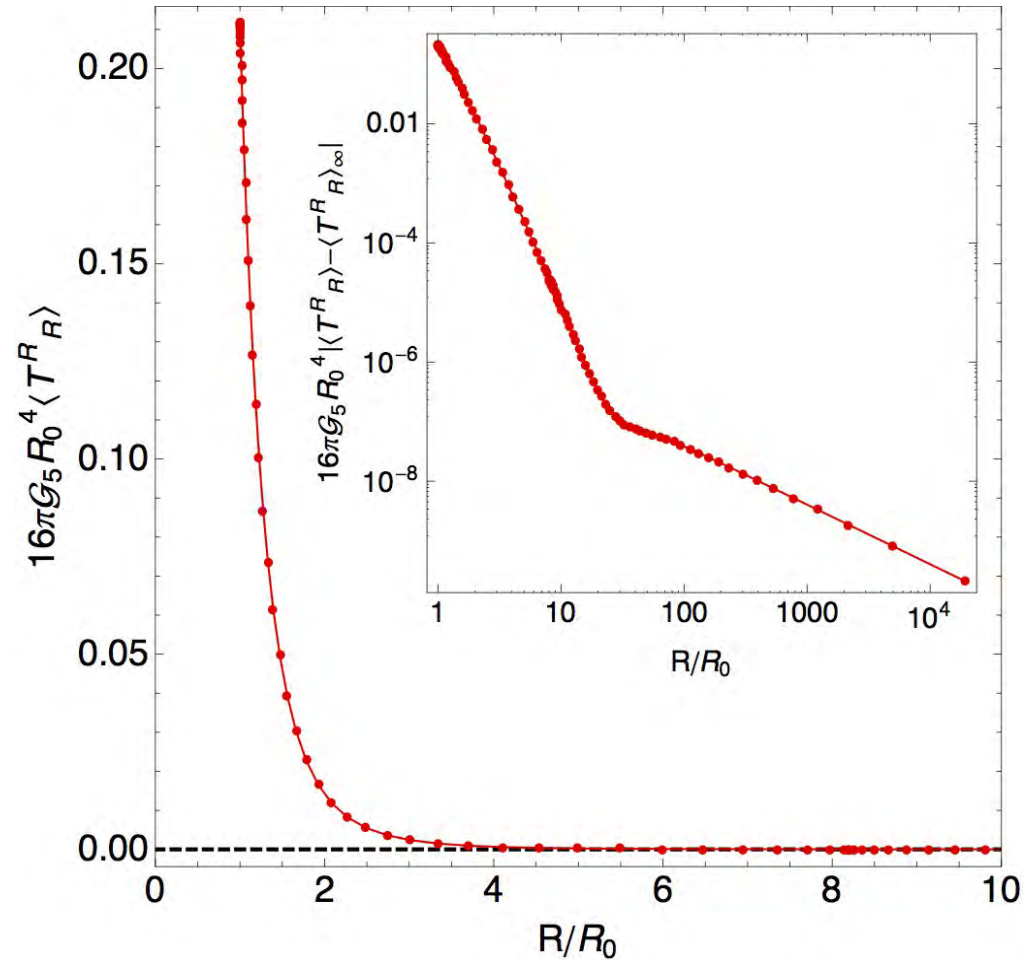
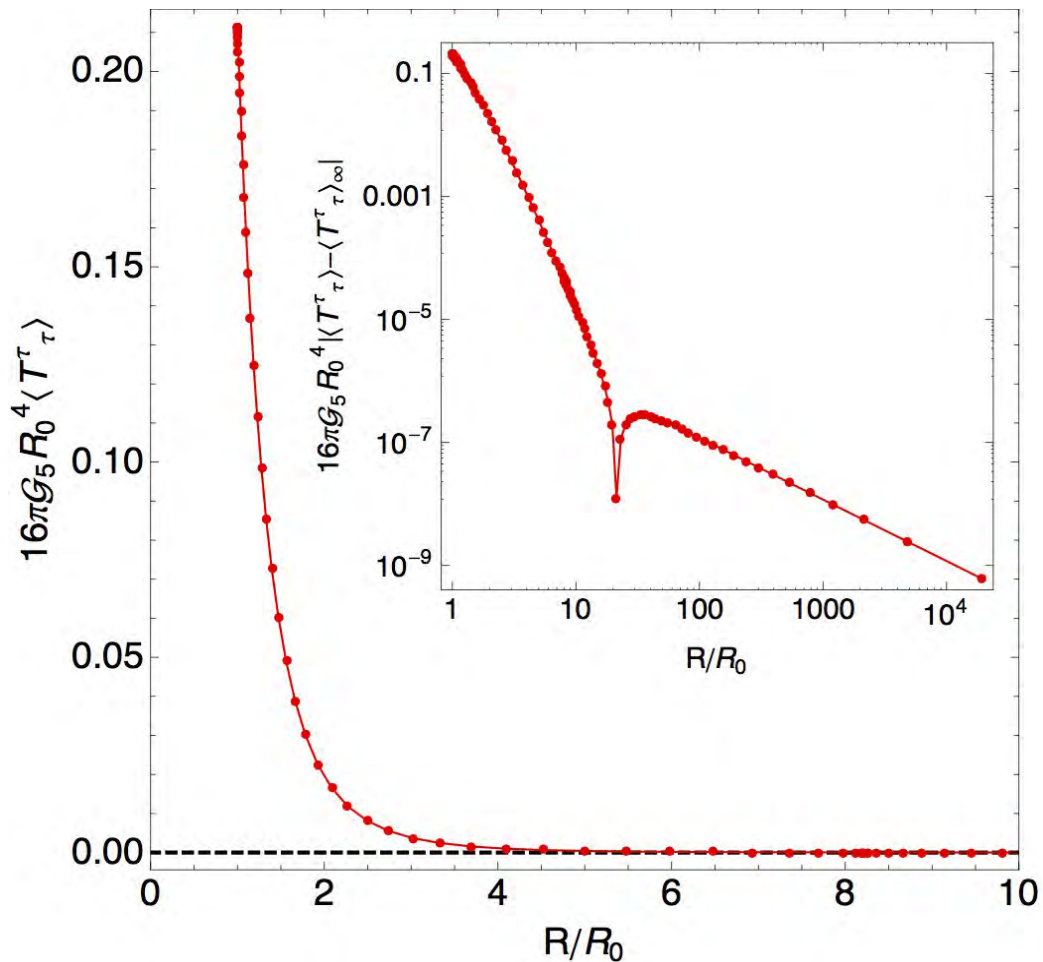


# Embedding in Hyperbolic Space



# Droplet Stress Tensor

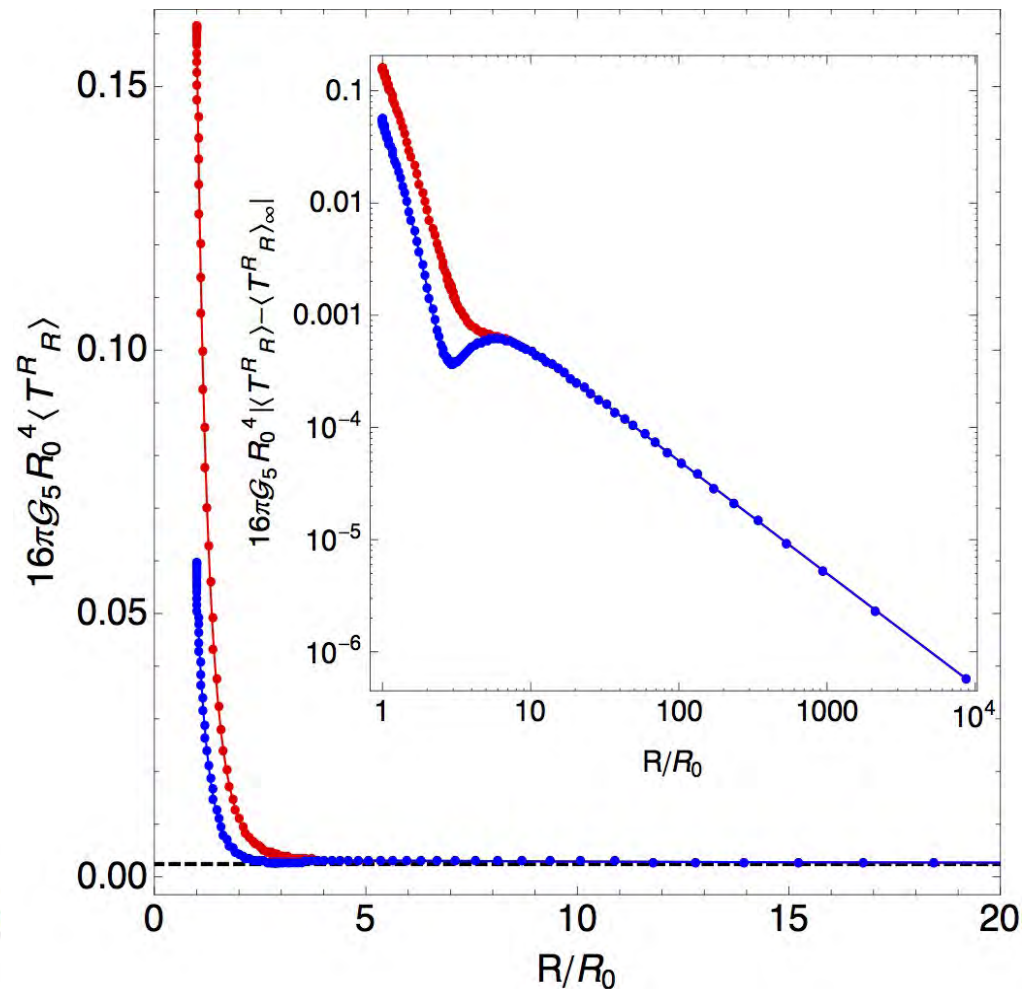
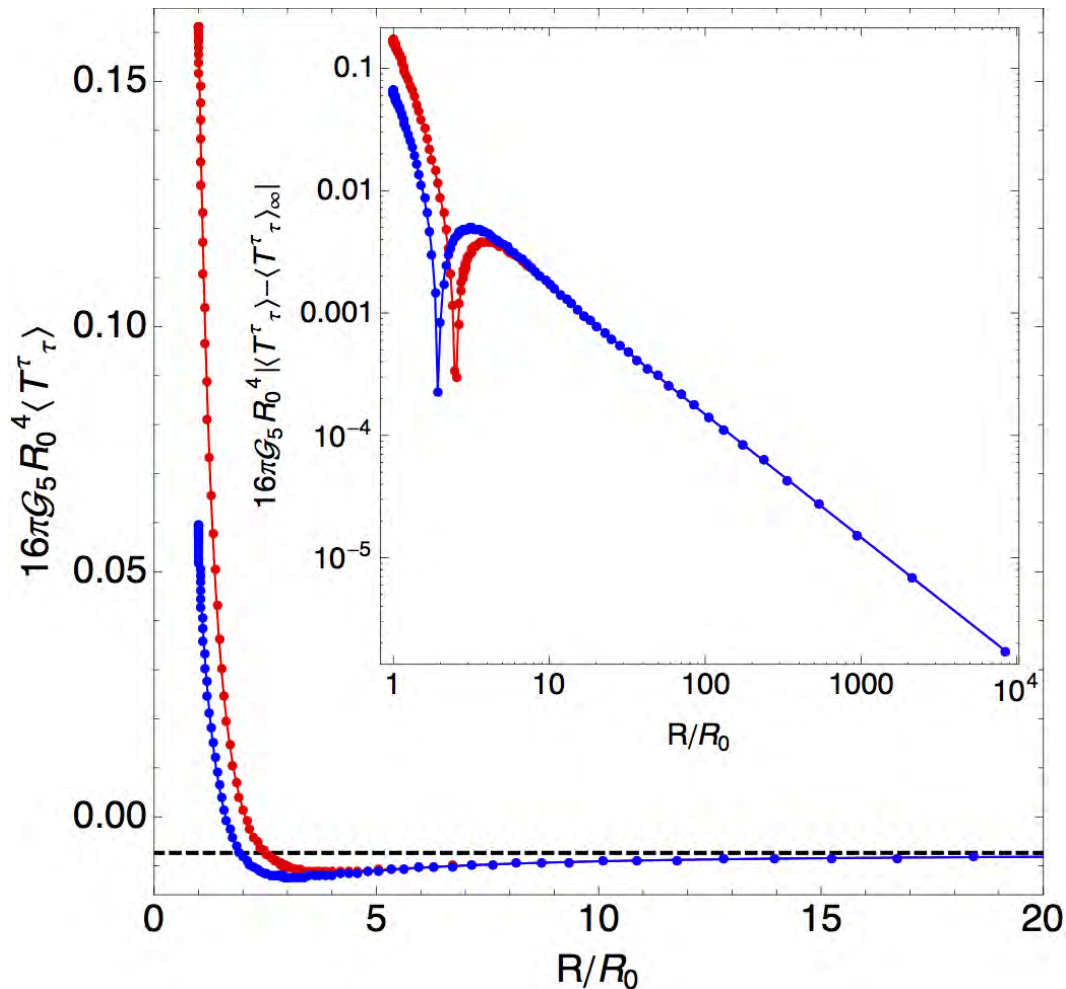
$$\frac{T_\infty}{T_{\text{BH}}} = 0.15$$



Stress tensor exhibits two power-law regimes.

# Droplet Stress Tensor

$$\frac{T_\infty}{T_{\text{BH}}} = 0.89$$



Falloff at infinity is universal.

# Summary

- Strongly coupled field theories on a black hole background can have two classes of solutions with different transport properties. These are black droplets and black funnels.
- In the 'droplet' phase, the field theory can be a poor conductor of energy-momentum and heat. These can be static, yet out of thermodynamic equilibrium.
- For a given  $T_\infty/T_{\text{BH}}$ , there can be two droplet phases: a long droplet or a short droplet.
- The Hartle-Hawking state of Schwarzschild is likely dual to a funnel, and funnels dominate for large  $T_\infty/T_{\text{BH}}$ .

# Future/Related Work

- Is there a cone transition? To what?
- Stability? Short droplets and equilibrium funnel likely stable. Long droplets might be unstable.
- Boulware state with  $T_{\text{bulk horizon}} = T_{\infty} = 0$ ? Quantum energy inequalities?



# Solutions with Non-Killing Horizons

- Flowing funnels needed to complete phase diagram.
- Existence of two kinds of funnels? (Wide neck/Narrow neck?)
- Extensions to Kerr/Rotating solutions.
- Equal-spinning Meyers-Perry droplets with  $T_\infty = 0$  found.
- Rotating funnels (Black twisters) unknown.
- Test fluid approximation.
- Are there turbulent instabilities?

Thank You