Entanglement Tsunami

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HL and Josephine Suh, 1305.7244, PRL 112, 011601 (2014) HL and Josephine Suh, 1311.1200, PRD 89, 066012 (2014) Casini, Hubeny, Maxfield, HL, Mezei, Suh, to appear



Quantum entanglement

A quantum system: divide into A+B

Hilbert space: $\mathcal{H} = \mathcal{H}_A \otimes \mathcal{H}_B$

Wave function:
$$\Psi = \sum_n \psi_n(A) \otimes \chi_n(B)$$

Simplest measure: entanglement entropy

$$\rho_A = \mathrm{Tr}_B |\Psi\rangle \langle \Psi|$$
$$S_A = -\mathrm{Tr}\rho_A \log \rho_A$$

Entanglement and phases of Matter

Traditional (Landau) paradigm of phases:

Different orders characterized by different symmetries Phase transitions: symmetry breaking

crystals, superconductors, magnets,

Such a classification **not** adequate:

gapped phases: FQH, spin liquids ...

topological order

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gapless systems:

Non-Fermi liquids, gapless spin liquids, novel quantum critical points Long range Quantum Entanglement (???)



Gapped phase: $\gamma\,$ provides a diagnostic of topological order.

Gapless systems:

Lorentz invariant QFTs



 γ provides a measure of number of degrees of freedom and decreases along RG

C-theorem or F-theorem



A new paradigm ?

Quantum gravity



Quantum information





In this talk I discuss our recent exploration of evolution of quantum entanglement in equilibration processes.

How a system equilibrates:

foundation of quantum statistical physics, condensed matter, QCD,

- Entanglement: nonlocal probes of equilibration processes
- Equilibration provides a dynamical setting to study the generation of entanglement

Entanglement generation



$$\psi(t = 0) = \psi_A \otimes \psi_B$$
$$\psi(t) = e^{-iHt}\psi(0)$$
$$H = H_A + H_B + H_{AB}$$

How fast can entanglement be generated?

In most physical systems: Local Hamiltonian

$$H_{AB} = H_{CD}$$
 δ : UV cutoff

Small incremental entangling conjecture/theorem



Dur, Vidal et al, Bravyi, Kitaev Bennett et al, Van Acoleyen, Marien, Verstraete

$$H = H_A + H_B + H_{CD}$$

For spin systems:

$$\frac{dS_A}{dt} \le c||H||\log d,$$

$$d = \min(d_C, d_D)$$

 $d_{\rm C}\,$: dimension of Hilbert space of C



- Finite dimensional Hilbert space: not applicable in continuum limit.
- Gapped systems

For more general quantum systems ????

Neither ||H|| nor d can be precisely defined

How do we compare systems of different number of dof, different shapes, sizes of A etc ?

Equilibration processes in quantum field theories provide a good laboratory for studying such question.

A simple setup: global quenches

- 1. Start with a QFT in the ground state.
- 2. At t=0 in a very short time interval inject a uniform energy density
 - initial state homogeneous, isotropic, entanglement properties as vacuum



3. The system evolves to (thermal) equilibrium

The system is in a pure state throughout.



Long ranged entangled d.o.f. are measure zero.

Entanglement in equilibrium state

The system behaves macroscopically as a thermal state, with entanglement entropy disguised as thermal entropy:

$$S_A^{\text{long,eq}} = s_{\text{eq}} V_A$$

 s_{eq} : equilibrium entropy density

V_A: volume of region A

Essentially all d.o.f. inside A becomes long ranged entangled with those outside A.



How? at what rate is entanglement generated?

$$\Delta S_A(t) = S_A(t) - S_A(t=0)$$

Previous results in (1+1)-d CFTs



Special techniques in one spatial dimension do not apply to higher dimensions:

• Equilibration processes: complicated nonequilibrium many-body dynamics, generally out of theoretical control.

• Entanglement entropy is notoriously difficult to calculate even for simple regions in the vacuum of a free theory, not to mention for general regions in interacting theories far from equilibrium.



String theory to the rescue!

Important earlier work:

Hubeny, Rangamani, Takayanagi: arXiv:0705.0016

Abajo-Arrastia, Aparicio and Lopez, arXiv:1006.4090

Albash and Johnson, arXiv:1008.3027

Balasubramanian, Bernamonti, de Boer, Copland, Craps, Keski-Vakkuri, Muller and Schafer, Shigemori, Staessens arXiv:1012.4753, arXiv:1103.2683

Aparicio and Lopez, arXiv:1109.3571

Caceres and A. Kundu, arXiv:1205.2354

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Holographic description of quench



quench: thin shell collapse to form a black hole.

Holographic Entanglement entropy

Ryu, Takayanagi Hubeny, Rangamani, Takayanagi





R: characteristic size of the region

Interested in long-distance physics: $R \to \infty$

Gravity description



Large size and critical extremal surfaces



In general a rather complicated problem to determine time evolution of extremal surfaces

Critical extremal surfaces determine large R, large time behavior

Four scaling regimes in general dimensions

In the large size R limit: $R \gg 1/T$



Linear growth

For $R \gg t \gg 1/T$

See also Hartman, Maldacena

$$\Delta S_A(t) = v_E \, s_{\rm eq} \, A_{\Sigma} t + \cdots$$

 $s_{\rm eq}$: Equilibrium entropy density

independent of shape, holographic theories under consideration, the nature of equilibrium state, also likely thermalization processes

v_E: dimensionless number characterizing final eq state.

Critical extremal surface for linear growth



The critical extremal surface runs along a constant radial slice inside the horizon

$$ds^{2} = \frac{L^{2}}{z^{2}} \left(-hdt^{2} + \frac{1}{f}dz^{2} + d\vec{x}^{2} \right)$$
$$z_{m}: \text{ minimum of } \frac{h(z)}{z^{2D}}$$

D: # of spatial dimensions

 $v_E = (z_h/z_m)^D \sqrt{-h(z_m)}$ z_h : horizon size

Entanglement Tsunami

$$\Delta S_A(t) = v_E \, s_{\text{eq}} \, A_\Sigma \, t = s_{\text{eq}} \, \left(V_A - V_{A-v_E t} \right)$$



suggests a picture of tsunami wave of entanglement, with a sharp wave front.

d.o.f. in the region covered by the wave is now entangled with those outside A

natural with evolution from a local Hamiltonian



Tsunami velocity

$$\Delta S_A(t) = v_E \, s_{\rm eq} \, A_\Sigma t + \cdots$$

Neutral system (AdS Schwarzschild):

$$v_E^{(S)} = \frac{(\eta - 1)^{\frac{1}{2}(\eta - 1)}}{\eta^{\frac{1}{2}\eta}} = \begin{cases} 1 & D = 1\\ \frac{\sqrt{3}}{2^{\frac{4}{3}}} = 0.687 & D = 2\\ \frac{\sqrt{2}}{3^{\frac{3}{4}}} = 0.620 & D = 3\\ \frac{\sqrt{2}}{3^{\frac{3}{4}}} = 0.620 & D = 3\\ \frac{1}{2} & D = \infty \end{cases}$$
$$\eta \equiv \frac{2D}{D+1}$$

Turning on chemical potential reduces v_E.

Upper bound on v_E ?

 v_E should be constrained by causality.

In all gravity examples:

$$v_E \le v_E^{(S)} = \frac{(\eta - 1)^{\frac{1}{2}(\eta - 1)}}{\eta^{\frac{1}{2}\eta}} \quad \eta \equiv \frac{2D}{D + 1}$$

Null energy condition important

Comparing with free particle streaming



Assume:

- At t=0, there is a uniform density of "photons" with only local entanglement correlations.
- Entanglement spreads when photons propagate.

Leading to shape independent linear growth,

$$\Delta S_{\Sigma}(t) = v_E s_{\rm eq} A_{\Sigma} t + \cdots$$



In strongly coupled systems, entanglement tsunami propagates faster than those from free particles traveling at speed of light !

$$D \to \infty: \ v_E^{(S)} \to \frac{1}{2}, \ v_{\text{streaming}} \to \sqrt{\frac{2}{\pi(D+1)}} \to 0$$

Bound on entanglement growth?

For any non-equilibrium processes:

$$\Re_A(t) \equiv \frac{1}{s_{\rm eq} A_{\Sigma}} \frac{dS_A}{dt}$$

dimensionless, can be compared among region A of different shapes, sizes, and systems of different number of d.o.f.

Indications from gravity: after local equilibration (t >>1/T)

$$\Re_A(t) \le v_E^{(S)}$$

Comparing with small incremental entangling conjecture/theorem:

$$\frac{dS_A}{dt} \le v_E^{(S)} s_{eq} A_{\Sigma}$$

$$\frac{dS_A}{dt} \le c ||H| \log d, \quad d = \min(d_C, d_D)$$



Future directions

• More examples:

Both holographic and field theoretical

• Continuum limit of small incremental conjecture

• Implications for black hole physics

Thank You

Supplementary slides

Local equilibration scale

The system should locally equilibrate first at a time scale ℓ_{eq} after which thermodynamics should apply locally.

Gravity side: horizon formation

Holographic systems:
$$\ell_{\rm eq} \sim {1 \over T} ~(\mu=0)$$

At $t \sim \ell_{\rm eq}$ nonlocal observables, such as $S_A(t)$ with $R \gg \ell_{\rm eq}$ may still be far from their equilibrium values.

Pre-local-equilibration evolution

For $t \ll \ell_{
m eq}$

$$S(t) - S(0) = \frac{\pi}{D} \epsilon A_{\Sigma} t^2 + \cdots$$

 ϵ : energy density A_{Σ} : area

independent of shape, independent of theories under consideration (for relativistic theories), and independent of the nature of final equilibrium state.

Saturation as a "phase transition"

Saturation: the tsunami covers the whole region:

In the limit quench is sharp, there is a sharp saturation time $t_s(\Sigma)$.

• Discontinuous saturation: first derivative of S(t) is discontinuous.

 Continuous saturations are characterized by critical behavior at saturation

$$S_{\Sigma}(t) - S_{\Sigma}^{(\mathrm{eq})} \propto -(t_s - t)^{\gamma}$$

Saturation: simple geometries



strip: discontinuous, linear growth until saturation

Strip :
$$t_s = \frac{R}{v_E}$$



Generic shape: curvature effects should be important, e.g. for a sphere, continuous saturation,

$$t_S = \frac{1}{c_{\text{sphere}}} R - \frac{D-1}{4\pi T} \log R + O(R^0)$$

Memory loss regime

For $t_s \gg t_s - t \gg \ell_{
m eq}$ with Σ a sphere of radius R

$$S(R,t) - S_{\rm eq}(R) = -s_{\rm eq}\lambda \left(t_s(R) - t\right)$$

 λ : volume of region not yet entangled Independent of the size of the region: size information lost

Speculation for generic shape :

Both shape and size will be forgotten at late times, flow to a "fixed point."



Gravity description



Gravity description

