Chiral Transport and the Hydrodynamic Frames

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Based on works:

[1508.06879 1509.08878 1612.0864 170?.????]

In collaboration with

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0. Outline

- Relativistic hydro
- Motivation for hydrodynamic chiral transport
- Non-dissipative feature of chiral transport
- Spectrum of chiral hydro modes
- Chiral hydro modes from Kinetic theory
- Frame choice and hydro modes

1. Hydrodynamics

Response of the system to perturbations:

at low energy and Long wave-length limit

Equations of hydro are local conservation equations

$$\partial_{\mu}T^{\mu\nu} = F^{\mu\nu}J_{\nu}$$
 variables
$$T(x), \ \mu(x), \ u^{\mu}(x)$$

$$\partial_{\mu}J^{\mu} = 0$$

$$(u^{\mu}u_{\mu} = -1)$$

Main idea of hydro: 14 variables in terms of 5 variables

2. Dual gravity picture?

Early studies:

Membrane Paradigm: Fluid on the horizon

[Damur; Thorne Price Macdonald 1970's]

New viewpoint:

Fluid-Gravity Duality: Fluid on the boundary

[Bhattacharyya Hubeny Minwalla Rangamani 2007]

3. Extensions of Fluid–Gravity:

- 1. Forced Fluid
- 2. Non-Relativistic Fluid
- 3. (Chirally) Charged Fluid

[Bhattacharyya Loganayagam Minwalla Nampuri Trividi 2007]

[Bhattacharyya Minwalla Wadia 2008]

[Erdmenger Haack Kaminski Yarom 2008]

[Banerjee Bhattacharyya Bhattacharyya Dutta Loganayagam Surowka 2008]

$$S = \frac{1}{16\pi G_5} \int \sqrt{-g_5} \left[R + 12 - F_{AB}F^{AB} - \frac{4\kappa}{3} \epsilon^{LABCD} A_L F_{AB} F_{CD} \right]$$
 anomaly

$$T_{\mu\nu} = p(\eta_{\mu\nu} + 4u_{\mu}u_{\nu}) - 2\eta\sigma_{\mu\nu} + \dots$$

$$J_{\mu} = n \ u_{\mu} - \mathfrak{D} \ P_{\mu}^{\nu} \mathcal{D}_{\nu} n + \xi \ l_{\mu} + \dots$$

vorticity: $l^{\mu} \equiv \epsilon^{\nu\lambda\sigma\mu} u_{\nu} \partial_{\lambda} u_{\sigma}$

5. Hydro with triangle anomalies:

Motivated by Fluid / Gravity:

In the presence of anomalies one may add parity odd terms:

$$J^{\mu} = nu^{\mu} - \sigma T P^{\mu\nu} \partial_{\nu} \left(\frac{\mu}{T} \right) + \sigma E^{\mu} + \xi \omega^{\mu} + \xi_B B^{\mu}$$

$$\partial_{\mu} j^{\mu} = C E^{\mu} B_{\mu},$$
$$\partial_{\mu} T^{\mu\nu} = F^{\nu\lambda} j_{\lambda}$$

[Son Surowka 2009]

[Kharzeev Yee ; Neiman Oz 2011]

[Jensen Loganayagam Yarom 2012]

In Landau_Lifhitz frame

$$\xi = \mathcal{C}\mu^{2} \left(1 - \frac{2}{3} \frac{\bar{n}\mu}{\bar{\epsilon} + \bar{p}} \right) + \mathcal{D}T^{2} \left(1 - \frac{2\bar{n}\mu}{\bar{\epsilon} + \bar{p}} \right)$$
$$\xi_{B} = \mathcal{C}\mu \left(1 - \frac{1}{2} \frac{\bar{n}\mu}{\bar{\epsilon} + \bar{p}} \right) - \frac{\mathcal{D}}{2} \frac{\bar{n}T^{2}}{\bar{\epsilon} + \bar{p}}$$

6. Chiral transport is non-dissipative:

Ohm Law:

$$oldsymbol{J} = \sigma oldsymbol{E}$$
 (-) (-)(+)

dissipative

London 2nd eq.:

$$\frac{\partial \boldsymbol{J}_s}{\partial t} = \sigma_s \boldsymbol{E}$$
(+) (+)(+)

non-dissipative

Transport in system of Single right-handed fermions:

$$J = \xi_B B$$
(-) (+)(-)

non-dissipative

7. A more realistic model:

Chiral hydrodynamics with both vector and axial currents:

$$\partial_\mu T^{\mu
u} = F^{
u \lambda} J_\lambda$$

$$\partial_\mu J^\mu = 0 \qquad \text{with}$$

$$\partial_\mu J^\mu_5 = \mathcal{C} E_\mu B^\mu$$

$$T^{\mu\nu} = (\epsilon + p)u^{\mu}u^{\nu} + p \eta^{\mu\nu}$$
$$J^{\mu} = nu^{\mu} + \xi \omega^{\mu} + \xi_{B}B^{\mu}$$
$$J^{\mu}_{5} = n_{5}u^{\mu} + \xi_{5}\omega^{\mu} + \xi_{B5}B^{\mu}$$

[Gao Liang Pu Wang Wang 2012]

[Landsteiner Megias Pena-Benitez 2013]

$$\xi = 2\mathcal{C} \left(\mu \mu_5 - \frac{n\mu_5}{3w} \left(3\mu^2 + \mu_5^2 \right) \right) - 2\mathcal{D} \frac{n\mu_5}{w} T^2$$

$$\xi_5 = \mathcal{C} \left(\mu^2 + \mu_5^2 - \frac{2n_5\mu_5}{3w} \left(3\mu^2 + \mu_5^2 \right) \right) + \mathcal{D} \left(1 - \frac{2n_5\mu_5}{w} \right) T^2$$

$$\xi_B = \mathcal{C} \mu_5 \left(1 - \frac{n\mu}{w} \right)$$

$$\xi_{5B} = \mathcal{C} \mu \left(1 - \frac{n_5\mu_5}{w} \right)$$

8. Hydro excitations in a chiral system:

Long-wavelength excitations around equilibrium:

$$u^{\mu} = (1, \mathbf{\Omega} \times \mathbf{x})$$
 $\Omega r \ll 1,$ $T = Const., \ \mu = Const., \ \mu_5 = Const.$ $\mathbf{B} = Const.$

$$\partial_{\mu} T^{\mu\nu} = F^{\nu\lambda} J_{\lambda}$$
$$\partial_{\mu} J^{\mu} = 0$$
$$\partial_{\mu} J^{\mu}_{5} = \mathcal{C} E_{\mu} B^{\mu}$$

perturbing the hydro fields:

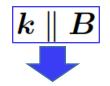
$$\phi_a = (T, \, \boldsymbol{\pi_i}, \, \mu, \, \mu_5), \qquad a = 1, 2, ..., 6$$

$$M_{ab}^{B\Omega}(\mathbf{k},\omega)\delta\phi_a(\mathbf{k},\omega)=0$$

9. Fluid coupled to weak magnetic field

$$\begin{bmatrix} -i\alpha_{1}\omega & ik_{j} & -i\alpha_{2}\omega & -i\alpha_{3}\omega \\ i\alpha_{1}v_{s}^{2}k^{i} & -i\omega\delta_{j}^{i} - i\frac{\xi}{2\bar{w}}\left(\boldsymbol{B}\cdot\boldsymbol{k}\delta_{j}^{i} - B_{j}k^{i}\right) - \frac{\bar{n}}{\bar{w}}\epsilon^{i}{}_{jl}B^{l} & i\alpha_{2}v_{s}^{2}k^{i} & i\alpha_{3}v_{s}^{2}k^{i} \\ -i\beta_{1}\omega + \left(\frac{\partial\xi_{B}}{\partial T}\right)i\boldsymbol{B}\cdot\boldsymbol{k} & \frac{\bar{n}}{\bar{w}}ik_{j} - \frac{\xi_{B}}{\bar{w}}i\omega B_{j} & -i\beta_{2}\omega + \left(\frac{\partial\xi_{B}}{\partial\mu}\right)i\boldsymbol{B}\cdot\boldsymbol{k} & -i\beta_{3}\omega + \left(\frac{\partial\xi_{B}}{\partial\mu_{5}}\right)i\boldsymbol{B}\cdot\boldsymbol{k} \\ -i\gamma_{1}\omega + \left(\frac{\partial\xi_{5B}}{\partial T}\right)i\boldsymbol{B}\cdot\boldsymbol{k} & \frac{\bar{n}_{5}}{\bar{w}}ik_{j} - \frac{\xi_{5B}}{\bar{w}}i\omega B_{j} & -i\gamma_{2}\omega + \left(\frac{\partial\xi_{5B}}{\partial\mu}\right)i\boldsymbol{B}\cdot\boldsymbol{k} & -i\gamma_{3}\omega + \left(\frac{\partial\xi_{5B}}{\partial\mu_{5}}\right)i\boldsymbol{B}\cdot\boldsymbol{k} \end{bmatrix}$$

6 linear coupled equations give 6 hydro modes:



1,2: Chiral-Magnetic-Heat wave:

$$\omega_{1,2}(k) = -\frac{\mathcal{A}_1 \pm \sqrt{\mathcal{A}_1^2 - \mathcal{A}_2 \mathcal{E}}}{\mathcal{E}} \boldsymbol{B}.\boldsymbol{k}$$

3,4: Ordinary sound wave:

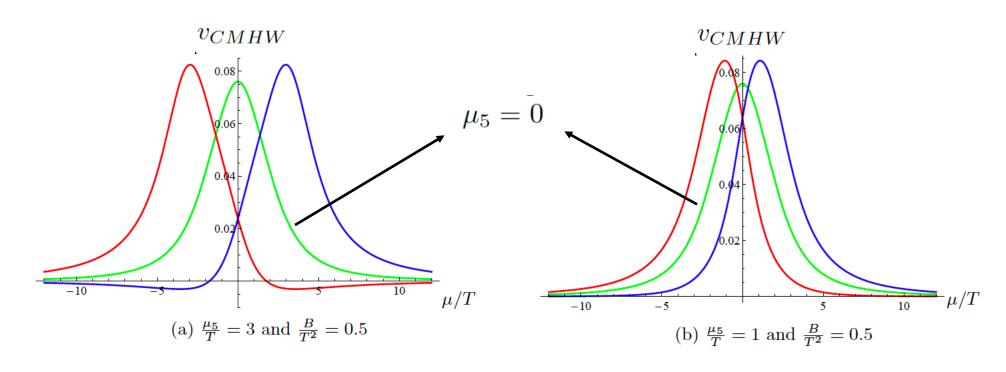
$$\omega_{3,4}(k) = \pm c_s k$$

5,6: Chiral Alfven wave:

$$\omega_{5,6}(k) = \pm \frac{\xi}{2w} Bk$$

10. Chiral –Magnetic–Heat Wave

Equation of state
$$\epsilon = 3p = \frac{7\pi^2}{60}T^4 + \frac{1}{2}(\mu^2 + \mu_5^2)T^2 + \frac{1}{4\pi^2}(\mu^4 + 6\mu^2\mu_5^2 + \mu_5^4)$$



[N.A Allahbakhshi Davody Taghavi 2016]

11. CMHW in QGP:

It is well_understood that the QGP produced in HIC in initially non_chiral: $\mu_5=0$

$$\omega_{1,2}(k) = -\frac{\mathcal{A}_1 \pm \sqrt{\mathcal{A}_1^2 - \mathcal{A}_2 \mathcal{E}}}{\mathcal{E}} B.k$$

$$\mu_5 = 0$$

CMW: [Kharzeev, Yee. 2012]

$$\omega_{1,2} = \pm \frac{Bk}{2\pi^2 \chi} \frac{1 - \frac{\mu n}{w}}{\sqrt{1 - \frac{\mu n}{w} - \frac{n}{\chi w} \left(\frac{n}{c_s^2} - \chi \mu\right)}}$$

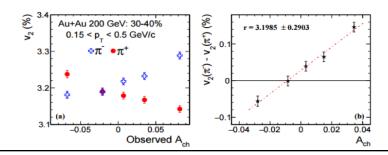
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Observation of Charge Asymmetry Dependence of Pion Elliptic Flow and the Possible Chiral Magnetic Wave in Heavy-Ion Collisions

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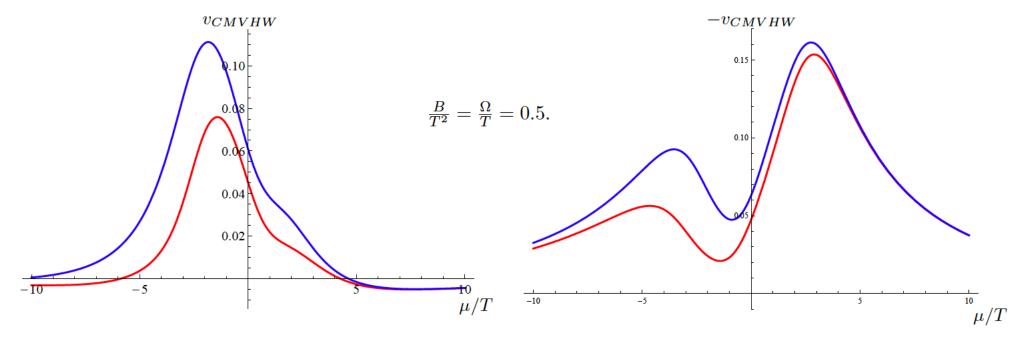
(STAR Collaboration)

We present measurements of π^- and π^+ elliptic flow, v_2 , at midrapidity in Au + Au collisions at $\sqrt{s_{\rm NN}} = 200, 62.4, 39, 27, 19.6, 11.5,$ and 7.7 GeV, as a function of event-by-event charge asymmetry, $A_{\rm ch}$, based on data from the STAR experiment at RHIC. We find that π^- (π^+) elliptic flow linearly increases (decreases) with charge asymmetry for most centrality bins at $\sqrt{s_{\rm NN}} = 27$ GeV and higher. At $\sqrt{s_{\rm NN}} = 200$ GeV, the slope of the difference of v_2 between π^- and π^+ as a function of $A_{\rm ch}$ exhibits a centrality dependence, which is qualitatively similar to calculations that incorporate a chiral magnetic wave effect. Similar centrality dependence is also observed at lower energies.



12. **Rotating** Fluid coupled to weak magnetic field

Scalar sector: mixed chiral-magnetic-vortical-heat wave



blue curve: $v_{sum}:=v_{CMHW}+v_{CVHW}=v_{CMVHW}|_{\Omega=0}+v_{CMVHW}|_{B=0}$ $v_{CMVHW}\neq v_{sum}$

13. Out of equilibrium dynamics from Kinetic theory

Chiral particles in kinetic theory: Chiral Kinetic Theory

$$I = \int_{t_i}^{t_f} (\boldsymbol{p} \cdot \dot{\boldsymbol{x}} + \boldsymbol{A} \cdot \dot{\boldsymbol{x}} - \Phi - |\boldsymbol{p}| - \boldsymbol{a_p} \cdot \dot{\boldsymbol{p}}) dt$$

[Stephanov Yee 2012]

[Son Yamamoto 2012]

$$|\hat{m{a}}_p| = rac{1}{p}$$
 Berry Curvature

 $|\hat{m{a}}_p| = rac{1}{p}$

$$\boldsymbol{b} = \boldsymbol{\nabla_p} \times \boldsymbol{a_p}$$

Berry Flux

Kinetic equation:

$$\frac{df}{dt} \equiv \frac{\partial f}{\partial t} + \frac{\partial f}{\partial \boldsymbol{x}} \dot{\boldsymbol{x}} + \frac{\partial f}{\partial \boldsymbol{p}} \dot{\boldsymbol{p}} = C[f]$$

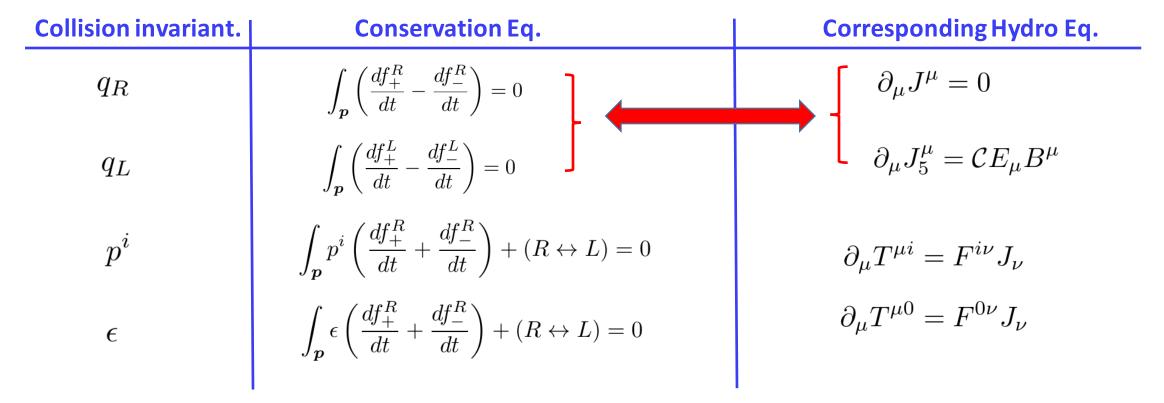
$$\sqrt{G}\dot{x} = \hat{p} + E \times b + B(\hat{p} \cdot b).$$

$$\sqrt{G}\dot{p} = E + \hat{p} \times B + b(E \cdot B).$$

14. Conservation equations in CKT:

distribution function
$$f_{\pm}^{R,L}\left(p;\;\beta(x),\boldsymbol{u}(x),\mu_{R,L}(x)\right) = \frac{1}{e^{\beta(x)\left(\epsilon-\boldsymbol{p},\boldsymbol{u}(x)\mp\mu_{R,L}(x)\right)}+1}$$

Integrating the kinetic equations after multiplying with collision invariant quantities:



15. Hydro modes from CKT

For slowly-varying macroscopic fields $\beta(x)$, u(x), $\mu_{R,L}(x)$

we expand the distribution function around the thermo distribution:

$$f_{\pm}^{R,L}(p) = \frac{1}{e^{\beta(\epsilon \mp \mu_{R,L})} + 1}$$

to find the linearized equations:

$$M_{ab} \delta \phi_a(\omega, \mathbf{k}) = 0$$

with: $\phi_a(x) = (\beta(x), \pi(x), \mu_R(x), \mu_L(x))$

16. Comparison between Hydro in LL and CKT

For a fluid of single right-handed fermions at $\mu = 0$:



relativistic hydro

Type of mode	No-drag fram	Landau-lifshitz frame
CMHW	$v_1 = \left(\frac{\xi}{2w} + \left(\frac{\partial \xi_B}{\partial n}\right)_{\epsilon}\right) B$	$v_1 = \left(\frac{\partial \xi_B}{\partial n}\right)_{\epsilon} B$
Sound	$v_{2,3} = \pm c_s - \frac{c_s^2 - 1}{2w} \xi B$	$v_{2,3} = \pm c_s$
CAW	$v_{4,5} = 0$	$v_{4,5} = -\frac{\xi}{2w} \mathbf{B}$

$$\xi = \mathcal{D}T^2$$

$$\mathcal{D} = \frac{1}{12}$$

17. Hydrodynamic frames:

Beyond the zero order in derivative expansion, there is an ambiguity in defining hydro variables. One can make transformations



while physical quantities $T^{\mu
u}$ J^{μ} and J^{μ}_{5}

Constructed by independent one-derivative data

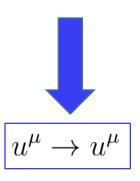
remaining invariant.

$$\nabla_{\mu}u^{\mu}, \; \partial_{\mu}^{\perp}\frac{\mu}{T}, \; B^{\mu}, \; \omega^{\mu}, \dots$$

Each choice of frame is equivalent to one way of fixing this ambiguity.

18. In the absence of dissipation

Thermodynamic frame



$$T \to T$$

$$\mu \to \mu + \mu_5$$

$$\mu_5 \to \mu_5$$

$$T^{\mu\nu} = wu^{\mu}u^{\nu} + pg^{\mu\nu} + \sigma_{\epsilon}^{\mathcal{B}}u^{(\mu}B^{\nu)} + \sigma_{\epsilon}^{\mathcal{V}}u^{(\mu}\omega^{\nu)}$$
$$j^{\mu} = nu^{\mu} + \sigma^{\mathcal{V}}\omega^{\mu} + \sigma^{\mathcal{B}}B^{\mu}$$

$$\delta u^{\mu} = -\frac{1}{\epsilon + p} \left(\sigma_{\epsilon}^{\mathcal{B}} B^{\mu} + \sigma_{\epsilon}^{\mathcal{V}} \omega^{\mu} \right)$$

Landau-Lifshitz

$$u_{\mu}T^{\mu\nu}_{(1)} = 0$$



$$u^{\mu} \to u^{\mu} + (\#B^{\mu})$$

$$T^{\mu\nu} = (\epsilon + p)u^{\mu}u^{\nu} + p\eta^{\mu\nu}$$
$$J^{\mu} = nu^{\mu} + \xi\omega^{\mu} + \xi_B B^{\mu}$$



19. Hydrodynamic modes are **not** frame invariant

Thermodynamic (Lab)	Landau-Lifshitz.	Boost
$u^{\mu}=(1,0)$	$u^{\mu} = (1, 0)$	$v_{LL} = rac{v_{Lab} - ilde{v}_{rel}}{1 - v_{Lab} ilde{v}_{rel}}$ $ ilde{v}_{rel} = -rac{\sigma_{\epsilon}^{\mathcal{B}} B + \sigma_{\epsilon}^{\mathcal{V}} \omega}{w}$
$u^{\mu} = (1, 0)$	$u^{\mu} = (1, \tilde{\boldsymbol{v}}_{rel}), u^{\mu}u_{\mu} = -1 + O(\tilde{\boldsymbol{v}}_{rel}^2)$	$v_{LL} = v_{Lab}$

Physical frame, consistent with Vilenkin 89 results.

"No-drag frame: [Stephanov Yee 2015]"

20. QCD fluid:

In single chirality fluid: CAW does not exist.

[Yamamoto 2105]

[N.A Davody Rezaei Hejazi 2016]

In QCD type fluid: CAW does exist:

$$v_{CAW} = \left(\mathcal{C}\mu\mu_5 - \frac{2n}{3w}\mathcal{C}\mu_5(3\mu^2 + \mu_5^2) - \frac{2n}{w}\mathcal{D}T^2\right)B$$

$$\mathcal{C} = \frac{1}{2\pi^2}, \qquad \mathcal{D} = \frac{1}{6}$$

Gauge-gravitational anomaly in QGP.

21. Open questions

1. How would change the situation when the magnetic field is dynamical?

2. Will be the CAW propagate when considering the back-reaction effects?

3. The full spectrum of chiral magneto hydrodynamics?

Thank you