

Holographic Two-Point Functions of Conformal Gravity

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Based on:

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- **Conformal gravity action** is given by a Weyl squared term in the form

$$S_{CG} = \alpha_{CG} \int d^4x \sqrt{|g|} C^{\alpha\beta\gamma\delta} C_{\alpha\beta\gamma\delta},$$

which is invariant under the Weyl transformation $g_{\mu\nu} \rightarrow e^{2\Omega(x)} g_{\mu\nu}$.

- α_{CG} is a **dimensionless** coupling constant.
- A **positive definite Euclidean action** and an **acceptable Newtonian limit** is possible only for $\alpha_{CG} > 0$. ($\alpha_{CG} = 1$)
(B. Hasslacher and E. Mottola, 1981).
- The equation of motion extracted from this action is known as the **Bach equation** (Bach, 1921).

$$\left(\nabla^\delta \nabla_\gamma + \frac{1}{2} R_\gamma^\delta \right) C_{\alpha\delta\beta}^\gamma = 0.$$

- It is well-known that the **solutions of Einstein gravity**, and as a specific example the *AdS* space-time, **are also solutions of Bach equation**.
- Moreover because of the the higher derivative nature of CG, the **Bach equation admit also** solutions which **are not Einstein spaces**.
- The most general **spherically symmetric solution of CG** is given by the line-element ([R. J. Riegert, 1984](#))

$$ds^2 = -A(r)dt^2 + \frac{dr^2}{A(r)} + r^2 d\Omega^2,$$

where $d\Omega^2$ is the line element of the round 2-sphere and

$$A(r) = \sqrt{1 - 12aM} - \frac{2M}{r} - \Lambda r^2 + 2ar.$$

- In contrast with **Einstein gravity which is not renormalizable** (M. H. Goroff and A. Sagnotti, 1985),
- **CG is power-counting renormalizable** (K. S. Stelle, 1977)
 - in fact, asymptotically free (J. Julve and M. Tonin, 1978, E. S. Fradkin and A. A. Tseytlin, 1982)
 - and therefore is considered as a possible UV completion of gravity (S. L. Adler, 1983).
- CG is **emerged in the Gauge/Gravity duality as a counter term** (M. Henningson and K. Skenderis – H. Liu and A. A. Tseytlin, 1998).
- CG also arises in **twistor-string theory** (N. Berkovits and E. Witten, 2004).

- In recent years CG has been in the center of attention. It is used to explain galactic rotation curves without need for dark matter (P. D. Mannheim, 2010)
- CG is equal (at the linearized level) to Einstein gravity by imposing a special boundary condition (J. Maldacena , 2011).

- More recently, it is shown that the (D. Grumiller, M. Irakleidou, I. Lovrekovic and R. McNees, 2014, PRL)
 - On-shell action for the four dimensional conformal gravity is renormalized **without need to counterterm**.
 - It is observed that the free energy derived from the on-shell action is consistent with the Arnowitt-Deser-Misner mass and Walds definition of the entropy (H. Lu, Y. Pang, C. Pope, and J. F. Vazquez-Poritz, 2012).
 - Doing the near boundary analysis, it has been argued that the **first two coefficients** in the Fefferman-Graham (FG) expansion of the boundary metric can consistently be interpreted as **two independent sources** for two operators in the boundary theory.
 - The **one-point functions** of that operators which are the **Energy-Momentum (EM) tensor** and the **Partial Massless Response (PMR)** have been worked out.
- **Our work** is finding the **two-point functions** of that operators.

- In the FG-gauge

$$ds^2 = \frac{l^2}{\rho^2} \left(-\sigma d\rho^2 + \gamma_{ij} dx^i dx^j \right),$$

it can be proven (A.N and K.Skenderis, unpublished) that close to $\rho = 0$,

$$\gamma_{ij} = \gamma_{ij}^{(0)} + \frac{\rho}{L} \gamma_{ij}^{(1)} + \frac{\rho^2}{L^2} \gamma_{ij}^{(2)} + \frac{\rho^3}{L^3} \gamma_{ij}^{(3)} + \dots$$

- Specification of **boundary conditions** (D. Grumiller, M. Irakleidou, I. Lovrekovic and R. McNees, 2014)

$$\delta\gamma_{ij}^{(0)}|_{\partial M} = 2\lambda\gamma_{ij}^{(0)}, \quad \delta\gamma_{ij}^{(1)}|_{\partial M} = \lambda\gamma_{ij}^{(1)},$$

- **Consistency of the boundary conditions** checked by
 - Considering on-shell action and the variational principle.
 - On-shell action for that metric remains finite.
 - Free-energy from the on-shell action leads to consistent thermodynamics.
 - The entropy derived in this way is the same as the entropy derived using Wald's Noether charge technique.

- The first variation of CG action is (D. Grumiller, M. Irakleidou, I. Lovrekovic and R. McNees, 2014)

$$\delta S_{CG} = e.o.m + \int_{\partial M} d^3x \sqrt{-\gamma} (T^{ij} \delta \gamma_{ij} + P^{ij} \delta K_{ij}).$$

- T^{ij} and P^{ij} are holographic response functions conjugate to the sources γ_{ij}^0 and γ_{ij}^1 respectively
 - $T^{ij} \rightarrow$ Brown-York stress tensor.
 - $P^{ij} \rightarrow$ Partially massless response.
 - $\gamma_{ij}^{(1)}$ plugged the linearized CG-e.o.m around (A)dS background exhibits **partial masslessness** behaviour.

- In order to find two-point functions, one should calculate the second variation of the on-shell action with respect to the corresponding sources. Alternatively one can obtain it by **varying the one-point function with respect to the source**

$$\langle P_{ij}(x) P_{kl}(0) \rangle = \frac{i}{\sqrt{-\gamma(0)}} \frac{\delta \langle P_{ij}(x) \rangle}{\delta \gamma_{(1)}^{kl}(0)},$$

$$\langle T_{ij}(x) T_{kl}(0) \rangle = \frac{i}{\sqrt{-\gamma(0)}} \frac{\delta \langle T_{ij}(x) \rangle}{\delta \gamma_{(0)}^{kl}(0)},$$

$$\langle P_{ij}(x) T_{kl}(0) \rangle = \frac{i}{\sqrt{-\gamma(0)}} \frac{\delta \langle P_{ij}(x) \rangle}{\delta \gamma_{(0)}^{kl}(0)} = \frac{i}{\sqrt{-\gamma(0)}} \frac{\delta \langle T_{kl}(0) \rangle}{\delta \gamma_{(1)}^{ij}(x)},$$

- This is where the **information from inside the bulk comes into play**. That is, we should solve exactly the equations of motion by specifying convenient boundary conditions.

- To find the linearized equations of motion for metric fluctuations around the AdS_4 space-time we substitute $\gamma_{ij} = \eta_{ij} + f_{ij}$ in Bach equation. In this way

$$r\left(\frac{1}{3}\square f'' - \partial^i \partial^j f''_{ij}\right) - \frac{1}{2}\partial^a \partial^b f'_{ab} + \frac{1}{6}\square f' + \frac{1}{12}\square^2 f - \frac{1}{12}\partial^a \partial^b \square f_{ab} = 0,$$

$$\begin{aligned} & r^3\left(\frac{16}{3}f''''\eta_{ij} - 16f''''_{ij}\right) + r^2\left((16f'''' + \frac{8}{3}\square f'' - \frac{4}{3}\partial^a \partial^b f''_{ab})\eta_{ij} - 48f''''_{ij}\right. \\ & - 8\square f''_{ij} - \frac{4}{3}\partial_i \partial_j f'' + 4\partial_i \partial^a f''_{aj} + 4\partial_j \partial^a f''_{ai}) + r\left((4f'' + \frac{4}{3}\square f' - \frac{2}{3}\partial^a \partial^b f'_{ab}\right. \\ & - \frac{1}{3}\partial^a \partial^b \square f_{ab} + \frac{1}{3}\square^2 f)\eta_{ij} - 12f''_{ij} - 4\square f'_{ij} + 2\partial_i \partial^a f'_{aj} + 2\partial_j \partial^a f'_{ai} - \frac{2}{3}\partial_i \partial_j f' \\ & \left. + \partial_i \partial^a \square f_{aj} + \partial_j \partial^a \square f_{ai} - \frac{1}{3}\partial_i \partial_j \square f - \square^2 f_{ij} - \frac{2}{3}\partial_i \partial_j \partial^a \partial^b f_{ab}\right) = 0. \end{aligned}$$

$$2r^2(\partial^j f''_{ij} - \frac{1}{3}\partial_i f'''') + r(-\partial_i f'' + \frac{1}{2}\partial^a \square f'_{ia} - \frac{1}{6}\partial_i \square f' - \frac{1}{3}\partial_i \partial^a \partial^b f'_{ab}) = 0,$$

- The equations of motion are **looking complicated** to solve.
- However by using the **reparametrization invariance and the Weyl symmetry**, one can write a simple form for the linearized equations

$$\frac{3}{2}(\square - \frac{4}{3}\Lambda)(\square - \frac{2}{3}\Lambda)\tilde{h}_{\mu\nu} = 0,$$

- In this way, solutions are classified into **two different modes** with the following differential equation

$$(\square + a^2)\tilde{h}_{\mu\nu} = 0, \quad a^2 = 2, 4.$$

- To construct a time-like mode it is possible to choose $p_i = E\delta_i^t$ as a time-like three-momentum.
- The e.o.m becomes

$$4r^2\tilde{h}''_{rr} + 6r\tilde{h}'_{rr} + (rE^2 + a^2 - 4)\tilde{h}_{rr} = 0,$$

$$4r^2\tilde{h}''_{ir} + 6r\tilde{h}'_{ir} + (rE^2 + a^2 - 4)\tilde{h}_{ir} - 4r\partial_i\tilde{h}_{rr} = 0,$$

$$4r^2\tilde{h}''_{ij} + 6r\tilde{h}'_{ij} + (rE^2 + a^2 - 4)\tilde{h}_{ij} - 4r\partial_i\tilde{h}_{rj} - 4\partial_j\tilde{h}_{ri} + 8r\eta_{ij}\tilde{h}_{rr} = 0.$$

- Solving the above equations with appropriate boundary conditions and after that using below transformation to **convert the solution to FG gauge**:

$$\xi^\mu(r, x^i) = e^{-iEt}\xi^\mu(r),$$

$$r^{-1}f_{\mu\nu} = \tilde{h}_{\mu\nu} + \nabla_\mu\xi_\nu + \nabla_\nu\xi_\mu.$$

- The **Ghost modes** are solutions for $a^2 = 4$.
- Imposing proper boundary conditions
 - We consider the "**infalling boundary condition in the Bulk**". That means that we require the behavior of solution around the Poincare horizon of the bulk, i.e. at $r \rightarrow \infty$, to be of the form $e^{\pm iE(t-\sqrt{r})}$, so that the fluctuation modes go toward the horizon as time passes and do not come out of it.
 - The expansion of these modes **near the boundary at $r = 0$** contains **no r_0 terms**. We impose this condition in order to make sure that while computing the two-point functions by varying one-point functions with respect to the source $\tilde{h}_{ij}^{(1)}$, the source $\tilde{h}_{ij}^{(0)}$ is automatically turned off in the last step of calculation.

- After imposing proper boundary condition, the covariant time-like solution in terms of three-momentum p_i as follows

$$f_{ij}^{G(1,2)} = i(1 - e^{i|\rho|\sqrt{r}}) \frac{1}{|\rho|} (p_i \epsilon_j^{1,2} + p_j \epsilon_i^{1,2}),$$

$$f_{ij}^{G(3,4)} = \sqrt{r} e^{i|\rho|\sqrt{r}} M_{ij}^{1,2},$$

$$f_{ij}^{G(5)} = \frac{2i}{|\rho|} (1 - e^{i|\rho|\sqrt{r}}) (\eta_{ij} - 2 \frac{p_i p_j}{p^2}) - \sqrt{r} e^{i|\rho|\sqrt{r}} (\eta_{ij} - \frac{p_i p_j}{p^2}) + i|\rho|r \frac{p_i p_j}{p^2},$$

where $\epsilon_i^{1,2}$ are two transverse polarizations from which, $M_{ij}^{1,2}$ are constructed

$$M_{ij}^1 = \epsilon_i^1 \epsilon_j^1 - \epsilon_i^2 \epsilon_j^2, \quad M_{ij}^2 = \epsilon_i^1 \epsilon_j^2 + \epsilon_i^2 \epsilon_j^1.$$

- The Einstein modes are solutions for $a^2 = 2$.
- After imposing proper boundary conditions,

$$f_{ij}^{E(1,2)} = (p_i \epsilon_j^{1,2} + p_j \epsilon_i^{1,2}),$$

$$f_{ij}^{E(3,4)} = e^{i|\rho|\sqrt{r}}(1 - i|\rho|\sqrt{r})M_{ij}^{1,2},$$

$$f_{ij}^{E(5)} = p_i p_j, \quad f_{ij}^{E(6)} = -2\eta_{ij} + r p_i p_j.$$

- The precise form of one-point functions are

$$P_{ij} = -\frac{4\sigma}{\ell} E_{ij}^{(2)},$$

$$T_{ij} = \sigma \left[\frac{2}{\ell} (E_{ij}^{(3)} + \frac{1}{3} E_{ij}^{(2)} f^{(1)}) - \frac{4}{\ell} E_{ik}^{(2)} \psi_j^{(1)k} + \frac{1}{\ell} f_{ij}^{(0)} E_{kl}^{(2)} \psi_{(1)}^{kl} \right. \\ \left. + \frac{1}{2\ell^3} \psi_{ij}^{(1)} \psi_{kl}^{(1)} \psi_{(1)}^{kl} - \frac{1}{\ell^3} \psi_{kl}^{(1)} (\psi_i^{(1)k} \psi_j^{(1)l} - \frac{1}{3} f_{ij}^{(0)} \psi_m^{(1)k} \psi_{(1)}^{lm}) \right] \\ - 4 D^k B_{ijk}^{(1)} + i \leftrightarrow j,$$

- where $\psi_{ij}^{(n)}$ is defined as the **traceless part** of $f_{ij}^{(n)}$

$$\psi_{ij}^{(n)} = f_{ij}^{(n)} - \frac{1}{3} f_{ij}^{(0)} f^{(n)}, \quad f^{(n)} = f_{(0)}^{ij} f_{ij}^{(n)}.$$

and

$$E_{ij}^{(2)} = -\frac{1}{2l^2} \psi_{ij}^{(2)} + \frac{\sigma}{2} (R_{ij}^{(0)} - \frac{1}{3} f_{ij}^{(0)} R^{(0)}) + \frac{1}{8l^2} f^{(1)} \psi_{ij}^{(1)} \\ B_{ijk}^{(1)} = \frac{1}{2\ell} (D_j \psi_{ik}^{(1)} - \frac{1}{2} f_{ij}^{(0)} D^l \psi_{kl}^{(1)}) - j \leftrightarrow k,$$

and $E_{ij}^{(3)}$ has complicated expression!!

- The correlation functions of two Partial Massless and two Stress Tensor operators are respectively

$$\begin{aligned} \langle P_{ij} P_{kl} \rangle &= i \frac{\delta \langle P_{ij} \rangle}{\delta f_{(1)}^{kl}} = \frac{1}{2|p|^3} (p_i p_k \Theta_{jl} + p_i p_l \Theta_{jk} + p_j p_k \Theta_{il} + p_j p_l \Theta_{ik}) \\ &\quad + \frac{1}{|p|^3} (\Theta_{ik} \Theta_{jl} + \Theta_{il} \Theta_{jk} - \Theta_{ij} \Theta_{kl}), \end{aligned}$$

$$\langle T_{ij} T_{kl} \rangle = i \frac{\delta T_{ij}}{\delta f_{(0)}^{kl}} = \frac{1}{|p|} (\Theta_{ik} \Theta_{jl} + \Theta_{il} \Theta_{jk} - \Theta_{ij} \Theta_{kl}),$$

where

$$\Theta_{ij} = \eta_{ij} p^2 - p_i p_j.$$

- And for mixed correlation functions we obtain

$$\langle T_{ij} P_{kl} \rangle = i \frac{\delta \langle T_{ij} \rangle}{\delta f_{(1)}^{kl}} = i \frac{\delta \langle P_{ij} \rangle}{\delta f_{(0)}^{kl}} = 0.$$

- This result is consistent with this fact that the correlation function of two operators with different scaling dimensions in a CFT vacuum is zero.
- Actually in the calculation of correlation functions, a relevant question is that, in which vacuum one is performing the computations.
- To answer this question, remind that we considered the linearisation of metric around the AdS_4 vacuum and observed that the on-shell perturbations respect the asymptotically AdS_4 form. According to asymptotically AdS_4 form of solutions, the dual field theory is a CFT.

- To further check the result one can use the Ward identities which come from the constraints on one-point functions in presence of sources

$$f_{ij}^{(0)} T^{ij} + \frac{1}{2} \psi_{ij}^{(1)} P^{ij} = 0, \quad f_{ij}^{(0)} P^{ij} = 0.$$

$$2D_i T^{ij} + 2D_i P^i_k h_{(1)}^{kj} + 2P^i_k D_i h_{(1)}^{kj} = P^{ik} D^j h_{ik}^{(1)}.$$

- Our correlation functions satisfy for example

$$\frac{\delta f_{ij}^{(0)}}{\delta f_{(0)}^{kl}} P^{ij} + f_{ij}^{(0)} \frac{\delta P^{ij}}{\delta f_{(0)}^{kl}} = 0,$$

and other constraints.

- Note that P_{ij} is traceless. According to the York decomposition of a traceless tensor A_{ij} we have

$$A_{ij} = \nabla_i V_j + \nabla_j V_i + P_{ij}^{TT} + (\nabla_i \nabla_j - \frac{1}{3} \eta_{ij} \nabla^2) S,$$

with V_i being a transverse vector ($\nabla^i V_i = 0$), P_{ij}^{TT} a transverse-traceless tensor and S being a scalar.

- Moreover, PMR modes are actually massive gravitons with $M^2 = \frac{2\Lambda}{3}$.
- Massive gravitons with this special amount of mass contain only spin two and spin one parts (S.Deser and A.Waldron, 2012). Thus

$$P_{ij} = \nabla_i V_j + \nabla_j V_i + P_{ij}^{TT}.$$

- Our boundary correlation functions also in full agreement with above fact in the bulk side.

- These are our results for two-point correlators which **all match the expectations for a CFT**

$$\langle T_{ij} T_{kl} \rangle = A_T \hat{\Theta}_{ij,kl} \frac{1}{|x|^2}, \quad \langle T_{ij} P_{kl} \rangle = 0$$

$$\langle P_{ij}^{TT} | P_{kl}^{TT} \rangle = A_P \hat{\Theta}_{ij,kl} \log|x|, \quad \langle V_i(x) | V_j(0) \rangle = A_V \frac{x_i x_j}{|x|^4},$$

where

$$\hat{\Theta}_{ij,kl} = \hat{\Theta}_{ik} \hat{\Theta}_{jl} + \hat{\Theta}_{il} \hat{\Theta}_{jk} - \hat{\Theta}_{ij} \hat{\Theta}_{kl},$$

and

$$A_T = 4\pi, \quad A_V = -4\pi, \quad A_P = -4\pi.$$

- The amplitudes A_V and A_P are **negative**.
- This observation leads to the conclusion that the **dual field theory is non-unitary**.

- We calculate the two-point functions for EM and PMR operators that have been identified as two response functions for two independent sources in the dual CFT.
- The correlation function of EM with PMR tensors turns out to be zero which is expected according to the conformal symmetry.
- The two-point function of EM is that of a transverse and traceless tensor, and the two-point function of PMR which is a traceless operator contains two distinct parts, one for a transverse-traceless tensor operator and another one for a vector field, **both of which fulfill criteria of a CFT.**
- We also observe the **absence of the scalar part in PMR.**
- It was claimed that the CG is unitary (C.M.bender and P.Mannheim, 2007,PRL). Our result shows that **CG is not unitary at least in the linearized level**

Thank You