

Hyperscaling-violating Lifshitz hydrodynamics from black holes

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Introduction

The holography gives the correspondences between black holes and fluids.

[Son-Starinets, 02] [Policastro-Son-Starinets, 02] [Bhattacharyya-Hubeny-Minwalla-Rangamani, 07]

Lifshitz spacetimes give holographic description of the Lifshitz scaling invariant theories.

[Son, 08] [Balasubramanian-McGreevy, 08]

$$ds^2 = -r^{2z} dt^2 + r^2(dx^i)^2 + \frac{dr^2}{r^2}$$

[Kachru-Liu-Mulligan, 08] [Taylor, 08]

Boundary theories in the Lifshitz spacetimes have the Newton-Cartan geometry as a background.

[Christensen-Hartong-Obers-Rollier, 14]

[Hartong-Kiritsis-Obers, 14, 14, 15]

Here we would like to see that

Fluid/gravity correspondence for the Lifshitz spacetime give a holographic description of fluids in Newton-Cartan background.

Einstein-Maxwell-Dilaton model

The action

$$S = \frac{1}{16\pi G} \int d^{d+1}x \sqrt{-g} \left(R - 2\Lambda - \frac{1}{4} e^{\lambda\phi} F^2 - \frac{1}{2} (\partial\phi)^2 \right)$$

Lifshitz spacetime solution

$$ds^2 = -r^{2z} dt^2 + \frac{dr^2}{r^2} + \sum_i r^2 (dx^i)^2$$

$$A = ar^{z+d-1} dt$$

$$e^{\lambda\phi} = \mu r^{2(1-d)}$$

$$\lambda^2 = 2 \frac{d-1}{z-1}$$
$$\Lambda = -\frac{(z+d-1)(z+d-2)}{2}$$

$$\mu a^2 = \frac{2(z-1)}{z+d-1}$$

The metric has Lifshitz scaling symmetry

$$t \rightarrow c^z t$$

$$x^i \rightarrow c x^i$$

$$r \rightarrow c^{-1} r$$

The gauge field A and dilaton ϕ breaks the scaling symmetry

Hydrodynamic ansatz

The black hole solution in the Eddington-Finkelstein coordinate

$$ds^2 = -r^{2z} f(r) dt^2 + 2r dr dt + r^2 (dx^i)^2$$

$$A = ar^{z+d-1} f(r) dt - ar^2 dr \quad e^{\lambda\phi} = \mu r^{2(1-d)}$$

$$f(r) = 1 - \frac{r_0^{z+d-1}}{r^{z+d-1}}$$

Introduce Galilean boost ($x^i \rightarrow x^i - v^i t$) and x^μ -dependence

$$ds^2 = -r^{2z} f(r) dt^2 + 2r dr dt + r^2 (dx^i - v^i(x) dt)^2$$

$$A = a(x) r^{z+d-1} f(r) dt - a(x) r^2 dr + \mathcal{A}_i(x) (dx^i - v^i(x) dt)$$

$$e^{\lambda\phi} = \mu(x) r^{2(1-d)} \quad f(r) = 1 - \frac{r_0^{z+d-1}(x)}{r^{z+d-1}}$$

This is not a solution of EOM



We have to introduce correction terms

Derivative expansion

We introduce the correction terms

$$ds^2 = -r^{2z} f(r) dt^2 + 2r dr dt + r^2 (dx^i - v^i(x) dt)^2 + h_{\mu\nu} dx^\mu dx^\nu$$

$$A = a(x) r^{z+d-1} f(r) dt - a(x) r^2 dr + \mathcal{A}_i(x) (dx^i - v^i(x) dt) + a_\mu dx^\mu$$

$$e^{\lambda\phi} = \mu(x) r^{2(1-d)} e^{\lambda\varphi}$$

Hydrodynamic regime: variation of x^μ -dependence is very slow.



derivative expansion (with respect to ∂_μ)

$$r_0(x) = r_0(0) + x^\mu \partial_\mu r_0(0) + \dots$$

$$v^i(x) = v^i(0) + x^\mu \partial_\mu v^i(0) + \dots$$

$$a(x) = a(0) + x^\mu \partial_\mu a(0) + \dots$$

Derivative expansion

We introduce the correction terms

$$ds^2 = -r^{2z} f(r) dt^2 + 2r dr dt + r^2 (dx^i - v^i(x) dt)^2 + h_{\mu\nu} dx^\mu dx^\nu$$

$$A = a(x) r^{z+d-1} f(r) dt - a(x) r^2 dr + \mathcal{A}_i(x) (dx^i - v^i(x) dt) + a_\mu dx^\mu$$

$$e^{\lambda\phi} = \mu(x) r^{2(1-d)} e^{\lambda\varphi}$$

First order EOM

Correction terms $\psi = (h_{\mu\nu}, a_\mu, \varphi)$ are treated as first order in ∂_μ .

$$\psi''(r) + F_1(r)\psi'(r) + F_2\psi(r) = F_3(\partial_\mu r_0, \partial_\mu v^i, \dots)$$



Linear terms of
correction terms $h_{\mu\nu}, a_\mu, \varphi$



Linear terms of
source terms $\partial_\mu r_0, \partial_\mu v^i, \partial_\mu a, \dots$

We solve these linear inhomogeneous differential equations

First order solution (for $d = 4, z = 2$)

$$ds^2 = -r^4 f(r) dt^2 + 2r dr dt + r^2 (dx^i - v^i dt)^2 \\ + \frac{2}{3} r^2 \partial_i v^i dt^2 - r^2 F(r) \sigma_{ij} (dx^i - v^i dt)(dx^j - v^j dt)$$

$$A = a \left(r^5 f(r) - \frac{1}{3} r^3 \partial_i v^i \right) dt - ar^2 dr + \mathcal{A}_i (dx^i - v^i dt)$$

$$e^{\lambda\phi} = \mu r^{2(1-d)}$$

$$f(r) = 1 - \frac{r_0^5}{r^5}$$

$$F(r) = \int dr \frac{r^3 - r_0^3}{r(r^5 - r_0^5)}$$

$$\sigma_{ij} = \partial_i v^j + \partial_j v^i - \frac{2}{3} \delta_{ij} \partial_k v^k$$

It must satisfy the following constraints

$$0 = \partial_t a + v^i \partial_i a - a \partial_i v^i$$

$$0 = \partial_t r_0 + v^i \partial_i r_0 + \frac{1}{3} r_0 \partial_i v^i$$

$$\mu(x) a^2(x) = \frac{2(z-1)}{z+d-1}$$

$$0 = \partial_t \mathcal{A}_i + v^j \partial_j \mathcal{A}_i + \mathcal{A}_j \partial_i v^j + a \partial_i r_0^5$$

Boundary stress-energy tensor for Lifshitz

Induced metric in terms of the vielbein

$$\gamma_{\mu\nu} = -r^{2z} f \tau_\mu \tau_\nu + r^2 e_\mu^a e_\nu^a \quad \gamma^{\mu\nu} = -r^{-2z} f^{-1} \hat{v}^\mu \hat{v}^\nu + r^{-2} e_a^\mu e_a^\nu$$

Gauge field in this frame: $\hat{A}_0 = \hat{v}^\mu A_\mu$ $\hat{A}_a = e_a^\mu A_\mu$

Variation of the action in these variables

$$\delta S = \int d^d x \left(-\hat{S}_\mu^0 \delta \hat{v}^\mu + \hat{S}_\mu^a \delta e_a^\mu + \hat{j}^0 \delta \hat{A}_0 + \hat{j}^a \delta \hat{A}_a + \mathcal{O}_\phi \delta \phi \right)$$

The stress-energy tensor and current are defined by

$$\hat{T}^\mu{}_\nu = \hat{S}_\nu^0 \hat{v}^\mu - \hat{S}_\nu^a e_a^\mu \quad J^\mu = \hat{j}^0 \hat{v}^\mu + \hat{j}^a e_a^\mu$$

The stress-energy tensor is related to Brown-York tensor $T^\mu{}_\nu$ as

$$\hat{T}^\mu{}_\nu = T^\mu{}_\nu + J^\mu A_\nu \quad \text{where} \quad T^{\mu\nu} = \frac{1}{8\pi G} (\gamma^{\mu\nu} K - K^{\mu\nu})$$

$\hat{T}^\mu{}_\nu$ is asymmetric tensor

$K_{\mu\nu}$: extrinsic curvature

Stress-energy tensor for Lifshitz fluid

We introduce the counter term

$$S_{\text{ct}} = \int d^4x \sqrt{-\gamma} \left[-(5+z) + \frac{z+d-1}{2} e^{\lambda\phi} \gamma^{\mu\nu} A_\mu A_\nu \right]$$

The renormalized stress-energy tensor

$$\tilde{T}^0_0 = \frac{1}{8\pi G} \frac{1}{r^{z+3}} \left[-\frac{3}{2} r_0^{z+3} - \frac{z-1}{a} v^i \mathcal{A}_i \right] + \mathcal{O}(r^{-(z+4)})$$

$$\tilde{T}^i_0 = \frac{1}{8\pi G} \frac{1}{r^{z+3}} \left[-\frac{z+3}{2} r_0^{z+3} v^i + \frac{z(z+3)}{4(z-1)} r_0^{2z} \partial_i r_0 - \frac{z-1}{a} v^i v^j \mathcal{A}_j + \frac{1}{2} r_0^3 \sigma_{ij} v^j \right] + \mathcal{O}(r^{-(z+4)})$$

$$\tilde{T}^0_i = \frac{1}{8\pi G} \frac{1}{r^{z+3}} \frac{z-1}{a} \mathcal{A}_i + \mathcal{O}(r^{-(z+4)})$$

$$\tilde{T}^i_j = \frac{1}{8\pi G} \frac{1}{r^{z+3}} \left[\frac{z}{2} r_0^{z+3} - \frac{1}{2} r_0^3 \sigma_{ij} + \frac{z-1}{a} v^i \mathcal{A}_j \right] + \mathcal{O}(r^{-(z+4)})$$

Stress-energy tensor for Lifshitz fluid

Since the volume form behaves as

$$\sqrt{-\gamma} \sim r^{z+3}$$

The leading terms of the following stress-energy tensor gives regular contributions

$$\tilde{T}^0_0 = \frac{1}{8\pi G} \frac{1}{r^{z+3}} \left[-\frac{3}{2} r_0^{z+3} - \frac{z-1}{a} v^i \mathcal{A}_i \right] + \mathcal{O}(r^{-(z+4)})$$

$$\tilde{T}^i_0 = \frac{1}{8\pi G} \frac{1}{r^{z+3}} \left[-\frac{z+3}{2} r_0^{z+3} v^i + \frac{z(z+3)}{4(z-1)} r_0^{2z} \partial_i r_0 - \frac{z-1}{a} v^i v^j \mathcal{A}_j + \frac{1}{2} r_0^3 \sigma_{ij} v^j \right] + \mathcal{O}(r^{-(z+4)})$$

$$\tilde{T}^0_i = \frac{1}{8\pi G} \frac{1}{r^{z+3}} \frac{z-1}{a} \mathcal{A}_i + \mathcal{O}(r^{-(z+4)})$$

$$\tilde{T}^i_j = \frac{1}{8\pi G} \frac{1}{r^{z+3}} \left[\frac{z}{2} r_0^{z+3} - \frac{1}{2} r_0^3 \sigma_{ij} + \frac{z-1}{a} v^i \mathcal{A}_j \right] + \mathcal{O}(r^{-(z+4)})$$

Constraints and conservation law

The following constraints in EOM gives the conservation law on the boundary

$$n^\mu \gamma^{\nu\rho} R_{\mu\nu} = 8\pi G n^\mu \gamma^{\nu\rho} T_{\mu\nu} \quad n_\nu \nabla_\mu (e^{\lambda\phi} F^{\mu\nu}) = 0$$

This should agree with the conservation law for $\tilde{T}^\mu{}_\nu$.

The constraints do not agree with $D_\mu \tilde{T}^\mu{}_\nu = 0$,
but agree with the conservation law in the Newton-Cartan theory.

$$D_\mu \mathcal{E}^\mu = -\frac{1}{2} (D^\mu \hat{v}^\nu + D^\nu \hat{v}^\mu) \mathcal{T}_{\mu\nu}$$

$$D_\mu \mathcal{T}^\mu{}_i = \hat{v}^\mu D_i \mathcal{P}_\mu - D_\mu (\hat{v}^\mu \mathcal{P}_i)$$

$$D_\mu \mathcal{J}^\mu = 0 \quad \text{where} \quad \mathcal{J}^\mu = \frac{1}{a} \hat{v}^\mu$$

Energy flow \mathcal{E}^μ , momentum density \mathcal{P}_μ and stress tensor $\mathcal{T}^\mu{}_\nu$ are defined as

$$\mathcal{E}^\mu = -\tilde{T}^\mu{}_\nu \hat{v}^\nu \quad \mathcal{P}_\mu = \tilde{T}^\rho{}_\nu \tau_\rho P_\mu^\nu \quad \mathcal{T}^\mu{}_\nu = \tilde{T}^\rho{}_\sigma P_\rho^\mu P_\nu^\sigma$$

$$\hat{v}^\mu = (1, v^i) \quad P_\nu^\mu = e_a^\mu e_\nu^a: \text{Projection to spatial direction}$$

Newton-Cartan geometry

Non-relativistic theory of gravity

Described by Newton-Cartan data $(\tau_\mu, h^{\mu\nu}, \hat{v}^\mu, \hat{A}_\mu)$

1-form τ defines time direction.

$h^{\mu\nu}$ is inverse metric on time slice.

} Galilei data

Galilei data $(\tau_\mu, h^{\mu\nu})$ satisfies

$$\tau_\mu h^{\mu\nu} = 0$$

Galilei data is constant. For covariant derivative D_μ

$$D_\mu \tau_\nu = 0 \quad D_\rho h^{\mu\nu} = 0$$

Galilei connection is not unique.

In order to define connection, we have to introduce \hat{v}^μ and \hat{A}_μ .

Conservation law in Newton-Cartan theory

In the Newton-Cartan theory variation of the action is

$$\delta S = \int d^d x \sqrt{-\gamma} [J^\mu \delta \hat{A}_\mu - \mathcal{P}_\mu \delta \hat{v}^\mu - \mathcal{E}^\mu \delta \tau_\mu - \frac{1}{2} \mathcal{T}_{\mu\nu} \delta h^{\mu\nu}]$$

Under the infinitesimal coordinate transformation ξ^μ (for torsionless case)

$$\delta_\xi \tau_\mu = D_\mu (\xi^\nu \tau_\nu) \qquad \delta_\xi h^{\mu\nu} = h^{\mu\rho} D_\rho \xi^\nu + h^{\nu\rho} D_\rho \xi^\mu$$

$$\delta_\xi \hat{v}^\mu = \xi^\rho D_\rho \hat{v}^\mu - \hat{v}^\rho D_\rho \xi^\mu \qquad \delta_\xi \hat{A}_\mu = -\hat{F}_{\mu\nu} \xi^\nu$$

Since the covariant derivative of \hat{v}^μ does not vanish, the conservation law is different from ordinary cases.

$$D_\mu \mathcal{E}^\mu = -\hat{F}_{\mu\nu} \hat{v}^\mu J^\nu - \frac{1}{2} (D^\mu \hat{v}^\nu + D^\nu \hat{v}^\mu) \mathcal{T}_{\mu\nu}$$

$$D_\mu \mathcal{T}^\mu{}_\nu = \hat{v}^\mu D_\nu \mathcal{P}_\mu - D_\mu (\hat{v}^\mu \mathcal{P}_\nu) - \hat{F}_{\mu\nu} J^\mu$$

Fluid variables and fluid equations

Energy flow, momentum density and stress tensor takes the following form

$$\begin{aligned}\mathcal{E}^0 &= \mathcal{E} & \mathcal{E}^i &= \mathcal{E}v^i - \kappa\partial_i T \\ \mathcal{P}_i &= q\mathcal{A}_i & \mathcal{T}_{ij} &= P\delta_{ij} - \eta\sigma_{ij}\end{aligned}$$

The energy density, pressure and charge density are

$$\mathcal{E} = \frac{3}{16\pi G}r_0^{z+3} \quad P = \frac{z}{16\pi G}r_0^{z+3} \quad q = \frac{z-1}{a} \quad T = \frac{z+3}{4\pi}r_0^z$$

The shear viscosity and thermal conductivity are

$$\eta = \frac{1}{16\pi G}r_0^3 \quad \kappa = \frac{1}{8G(z-1)}r_0^{z+1}$$

Bulk viscosity vanishes.

Energy density and pressure satisfy the Lifshitz scaling condition

$$z\mathcal{E} = (d-1)P$$

Fluid variables and fluid equations

Energy flow, momentum density and stress tensor takes the following form

$$\mathcal{E}^0 = \mathcal{E} \quad \mathcal{E}^i = \mathcal{E}v^i - \kappa\partial_i T$$

$$\mathcal{P}_i = q\mathcal{A}_i \quad \mathcal{T}_{ij} = P\delta_{ij} - \eta\sigma_{ij}$$

The conservation law gives the fluid equations

$$D_\mu \mathcal{E}^\mu = -\frac{1}{2}(D^\mu \hat{v}^\nu + D^\nu \hat{v}^\mu)\mathcal{T}_{\mu\nu} \quad D_\mu \mathcal{J}^\mu = 0$$

$$D_\mu \mathcal{T}^\mu_i = \hat{v}^\mu D_i \mathcal{P}_\mu - D_\mu(\hat{v}^\mu \mathcal{P}_i)$$

The fluid equation is expressed as

$$0 = \partial_t \mathcal{E} + v^i \partial_i \mathcal{E} + (\mathcal{E} + P)\partial_i v^i - \frac{1}{2}\eta\sigma_{ij}\sigma_{ij} - \partial_i(\kappa\partial_i T)$$

$$0 = \partial_i P + q\partial_t \mathcal{A}_i + qv^j \partial_j \mathcal{A}_i + q\mathcal{A}_j \partial_i v^j - \partial_j(\eta\sigma_{ij})$$

$$0 = \partial_t q + \partial_i(qv^i)$$

Non-relativistic fluid equations

The fluid equations from the Lifshitz black hole are

$$0 = \partial_t \mathcal{E} + v^i \partial_i \mathcal{E} + (\mathcal{E} + P) \partial_i v^i - \frac{1}{2} \eta \sigma_{ij} \sigma_{ij} - \partial_i (\kappa \partial_i T)$$

$$0 = \partial_i P + q \partial_t \mathcal{A}_i + q v^j \partial_j \mathcal{A}_i + q \mathcal{A}_j \partial_i v^j - \partial_j (\eta \sigma_{ij})$$

$$0 = \partial_t q + \partial_i (q v^i)$$

Ordinary fluid equations are

$$0 = \partial_t \mathcal{E} + v^i \partial_i \mathcal{E} + (\mathcal{E} + P) \partial_i v^i - \frac{1}{2} \eta \sigma_{ij} \sigma_{ij} - \partial_i (\kappa \partial_i T)$$

$$0 = \partial_i P + \rho \partial_t v^i + \rho v^j \partial_j v^i - \partial_j (\eta \sigma_{ij})$$

$$0 = \partial_t \rho + \partial_i (\rho v^i)$$

Energy conservation (1st line) and continuity equation (3rd line) agree.
But the Navier-Stokes does not agree.

Fluid equations in Newton-Cartan theory

The Newton-Cartan theory gives ordinary fluid equations.

Torsionless Newton-Cartan connection is invariant under Milne boost

$$v^i \rightarrow v^i + V^i \quad \hat{A} \rightarrow \hat{A} + V^i(dx^i - v^i dt) - \frac{1}{2}V^2 dt$$

If $\hat{A}_\mu = 0$, for $v^i = 0$, we obtain after the boost to $v^i \neq 0$

$$\hat{A} = v^i dx^i - \frac{1}{2}v^2 dt$$

Then, the Navier-Stokes equation

$$0 = \partial_i P + \rho \partial_t v^i + \rho v^j \partial_j v^i - \partial_j (\eta \sigma_{ij})$$

is expressed as

$$\partial_i P - \partial_j (\eta \sigma_{ij}) = \hat{F}_{i\mu} J^\mu \quad \text{where} \quad J^\mu = \rho \hat{v}^\mu$$

Navier-Stokes from Lifshitz black hole

The fluid equation from the Lifshitz black hole

$$0 = \partial_i P + q \partial_t \mathcal{A}_i + q v^j \partial_j \mathcal{A}_i + q \mathcal{A}_j \partial_i v^j - \partial_j (\eta \sigma_{ij})$$

can also be expressed as

$$\partial_i P - \partial_j (\eta \sigma_{ij}) = \mathcal{F}_{i\mu} J^\mu \quad \text{where} \quad \mathcal{A} = \mathcal{A}_i (dx^i - v^i dt)$$

The Navier-Stokes equation in the Newton-Cartan theory is

$$\partial_i P - \partial_j (\eta \sigma_{ij}) = \hat{F}_{i\mu} J^\mu \quad \text{where} \quad \hat{A} = v^i dx^i - \frac{1}{2} v^2 dt$$

They are in the same form but have **different gauge fields** in r.h.s.

Another ansatz for gauge field

What happens if we use $\mathcal{A} = v^i dx^i - \frac{1}{2} v^2 dt$

instead of $\mathcal{A} = \mathcal{A}_i(dx^i + v^i dt)$ from the beginning?

We generalize \mathcal{A}_t to arbitrary

$$ds^2 = -r^{2z} f(r) dt^2 + 2r dr dt + r^2 (dx^i - v^i(x) dt)^2$$

$$e^{\lambda\phi} = \mu(x) r^{2(1-d)} \quad f(r) = 1 - \frac{r_0^{z+d-1}(x)}{r^{z+d-1}}$$

$$A = a(x) r^{z+d-1} f(r) dt - a(x) r^2 dr + \mathcal{A}_i(x) (dx^i - v^i(x) dt)$$



$$A = a(x) r^{z+d-1} f(r) dt - a(x) r^2 dr + \mathcal{A}_t(x) dt + \mathcal{A}_i(x) dx^i$$

The solution does not change but constraint for $\mathcal{A} = v^i dx^i - \frac{1}{2} v^2 dt$

$$\partial_i P - \partial_j (\eta \sigma_{ij}) = \mathcal{F}_{i\mu} J^\mu = -q \partial_t v^i - q v^j \partial_j v^i$$

gives the non-relativistic Navier-Stokes equation.

The stress-energy tensor becomes

$$\tilde{T}^0_0 = \frac{1}{8\pi G} \frac{1}{r^{z+3}} \left[-\frac{3}{2} r_0^{z+3} - \frac{z-1}{a} v^i \mathcal{A}_i \right] + \mathcal{O}(r^{-(z+4)})$$

$$\tilde{T}^i_0 = \frac{1}{8\pi G} \frac{1}{r^{z+3}} \left[-\frac{z+3}{2} r_0^{z+3} v^i + \frac{z(z+3)}{4(z-1)} r_0^{2z} \partial_i r_0 + \frac{z-1}{a} v^i \mathcal{A}_t \right. \\ \left. + \frac{1}{2} r_0^3 \sigma_{ij} v^j \right] + \mathcal{O}(r^{-(z+4)})$$

$$\tilde{T}^0_i = \frac{1}{8\pi G} \frac{1}{r^{z+3}} \frac{z-1}{a} \mathcal{A}_i + \mathcal{O}(r^{-(z+4)})$$

$$\tilde{T}^i_j = \frac{1}{8\pi G} \frac{1}{r^{z+3}} \left[\left(z r_0^{z+3} - \frac{z-1}{a} (\mathcal{A}_t + v^k \mathcal{A}_k) \right) \delta_{ij} - \frac{1}{2} r_0^3 \sigma_{ij} + \frac{z-1}{a} v^i \mathcal{A}_j \right] \\ + \mathcal{O}(r^{-(z+4)})$$

\tilde{T}^μ_ν cannot be identified with the fluid stress-energy tensor.

We define new stress-energy tensor T^μ_ν as

$$\tilde{T}^\mu_\nu = T^\mu_\nu + J^\mu \mathcal{A}_\nu - \delta^\mu_\nu J^\rho \mathcal{A}_\rho$$

Energy density \mathcal{E}^μ , pressure P and stress tensor \mathcal{T}^i_j does not change but momentum density vanishes $\mathcal{P}_i = 0$, for new T^μ_ν .

Now, we take $F_{\mu\nu}J^\nu$ term in the conservation into account.

$$D_\mu \mathcal{E}^\mu = -\mathcal{F}_{\mu\nu} \hat{v}^\mu J^\nu - \frac{1}{2} (D^\mu \hat{v}^\nu + D^\nu \hat{v}^\mu) \mathcal{T}_{\mu\nu}$$

$$D_\mu \mathcal{T}^\mu_\nu = \hat{v}^\mu D_\nu \mathcal{P}_\mu - D_\mu (\hat{v}^\mu \mathcal{P}_\nu) - \mathcal{F}_{\mu\nu} J^\mu$$

Then, these equation gives the non-relativistic fluid equations

$$0 = \partial_t \mathcal{E} + v^i \partial_i \mathcal{E} + (\mathcal{E} + P) \partial_i v^i - \frac{1}{2} \eta \sigma_{ij} \sigma_{ij} - \partial_i (\kappa \partial_i T)$$

$$0 = \partial_i P + q \partial_t v^i + q v^j \partial_j v^i - \partial_j (\eta \sigma_{ij})$$

$$0 = \partial_t q + \partial_i (q v^i)$$

Holographic entropy current

Entropy current J_S^μ  Volume form on time slice at the horizon

$$\epsilon_{\mu_1 \dots \mu_d} J_S^{\mu_1} dx^{\mu_2} \wedge \dots \wedge dx^{\mu_d}$$

In terms of the normal vector to the horizon $n_\mu = \partial_\mu \mathcal{S}$ where $\mathcal{S} = r - r_0(x)$.

$$J_S^\mu = \frac{\sqrt{h} n^\mu}{4G n^0}$$

The entropy current for Lifshitz black hole is calculated as

$$J_S^0 = \frac{1}{4G} r_0^3 \quad J_S^i = \frac{1}{4G} r_0^3 v^i - \frac{z}{8G(z-1)} r_0^z \partial_i r_0$$

Thermodynamic relation

Energy flow, pressure and temperature

$$\mathcal{E}^0 = \frac{1}{16\pi G} r_0^{z+3} \quad \mathcal{E}^i = \frac{1}{16\pi G} \left(r_0^{z+3} v^i - \frac{z(z+3)}{2(z-1)} r_0^{2z} \partial_i r_0 \right)$$

$$P = \frac{z}{16\pi G} r_0^{z+3} \quad T = \frac{z+3}{4\pi} r_0^z$$

Entropy current

$$J_S^0 = \frac{1}{4G} r_0^3 \quad J_S^i = \frac{1}{4G} r_0^3 v^i - \frac{z}{8G(z-1)} r_0^z \partial_i r_0$$

They satisfy the thermodynamic relation

$$T J_S^\mu = -\tilde{T}^\mu_\nu \hat{v}^\nu + P \hat{v}^\mu = \mathcal{E}^\mu + P \hat{v}^\mu$$

Second law of thermodynamics

Fluid equations give

$$0 = 3\partial_t r_0^{z+3} + 3v^i \partial_i r_0^{z+3} + (z+3)r_0^{z+3} \partial_i v^i \\ - \sigma_{ij} \sigma_{ij} - \frac{z(z+3)}{2(z-1)} \partial_i (r_0^{2z} \partial_i r_0)$$

Divergence of the entropy current becomes

$$\partial_\mu J_S^\mu = \frac{1}{4G} \left[\partial_t r_0^3 + \partial_i \left(r_0^3 v^i - \frac{z}{2(z-1)} r_0^z \partial_i r_0 \right) \right] \\ = 2\pi \frac{\eta}{r_0^z} \sigma_{ij} \sigma_{ij} + \frac{z^2}{8G(z-1)} r_0^{z-1} (\partial_i r_0)^2 \geq 0$$

Divergence of the entropy current is non-negative



2nd law of thermodynamics

Conclusion

- We have considered fluid/gravity correspondence for the Lifshitz black hole geometry.
- Naïve ansatz gives the stress-energy tensor which satisfy the conservation law of the Newton-Cartan theory.
- Energy conservation and continuity equations agree with those for ordinary non-relativistic fluids.
- The Navier-Stokes equation is different from ordinary one, but in terms of the gauge field, it agrees with that in the Newton-Cartan theory.
- If we take $\mathcal{A} = v^i dx^i - \frac{1}{2} v^2 dt$, the constraints agrees with the ordinary non-relativistic fluid equations.
- But the boundary stress-energy tensor cannot simply be identified to the fluid stress-energy tensor. We have to subtract gauge field terms.
- Entropy current is defined from the horizon area and satisfies the local thermodynamic relation and second law.

Thank you

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