# Effective Holographic Theories for Condensed Matter Systems

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- Motivation: Strange Metallic Behaviour and Effective Holographic Theories for Condensed Matter Systems
- 2 Uncharged Solutions
- 3 Exact Charged Solutions •  $\gamma \delta = 1$ 
  - $\gamma = \delta$
  - 4 Near-Extremal Scaling Solutions
- 5 Conclusions & Remarks

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#### Motivation

### Motivation: Strange Metallic Behaviour

- Non-Fermi Liquid Behaviour: High Tc Cuprates (e.g. La<sub>2-x</sub>Sr<sub>x</sub>CuO<sub>4</sub>), Heavy Fermion Systems (e.g. CeCu<sub>5.9</sub>Au<sub>0.1</sub>), Iron based High Tc compounds (e.g. SmFeAsO<sub>1-x</sub>F<sub>x</sub>)
- Properties of Strange Metal Region:
  - $\rightarrow$  Linear DC Resistivity:  $\rho = \rho_0 + AT$  (down to T = 0)
  - $\rightarrow$  AC conductivity scaling:  $\sigma(\omega) \sim \omega^{-2/3}$  for  $T \ll \omega \ll \Lambda$
  - $\rightarrow$  Electronic Specific Heat and Entropy:  $C_{v}, S \sim T$



#### Motivation

# Motivation: Effective Holographic Theories

- <u>EHT:</u> p-dim. strongly coupled EFT  $\leftrightarrow p$  + 1-dim. weak gravity
- Minimal Ingredients for a High Tc EHT:
  - Finite Carrier Density: Global U(1) current  $J^{\mu} \leftrightarrow$  bulk gauge field  $A_{\mu}$
  - Leading Relevant Uncharged Operator:  $\mathcal{O}_{\Phi} \leftrightarrow$  bulk scalar  $\Phi$
  - Energy-Momentum-Tensor:  $T_{\mu
    u} \leftrightarrow$  bulk metric  $g_{\mu
    u}$
- Two-Derivative Infrared Effective Holographic Action: (Sugra Intuition)

$$S = M_{Pl}^{p-1} \int d^{p+1} x \sqrt{-g} \left[ R - \frac{1}{2} (\partial \Phi)^2 - V(\Phi) - \frac{Z(\Phi)}{4} F_{\mu\nu}^2 \right]$$
$$V(\Phi) = 2\Lambda e^{-\delta \Phi}, \qquad Z(\Phi) = e^{\gamma \Phi}$$

- (Un)Charged Black Hole Solutions: Important Observables
  - Thermodynamics (Entropy, Specific Heat, Phase Transitions)
  - DC/AC conductivities
  - Other transport coefficients, ...

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### **On Naked Singularities**

- In holography naked singularities may be acceptable, since they might not necessarily signal a breakdown of predictability of the theory, as long as they do not affect the field theory physics severely.
- Gubser gave the first criterion for good singularities : They should be limits of solutions with a regular horizon. [Gubser (2000)]
- The second criterion amounts to having a well-defined spectral problem for fluctuations around the solution: The second order equations describing all fluctuations are Sturm-Liouville problems (no extra boundary conditions needed at the singularity). [Gursoy+E.K.+Nitti (2008)]
- The singularity is "repulsive" (like the Liouville wall). It has an overlap with the previous criterion. It involves the calculation of 'Wilson loops" [Gursoy+E.K.+Nitti (2008)]
- It is not known whether the list is complete.

Motivation

# Conductivity & Charged Excitation Spectra I

 Conductivity is a main characteristic transport coefficient in a finite density system.

$$J^{i}(\omega,\vec{k})=\sigma^{ij}(\omega,\vec{k}) E_{j}(\omega,\vec{k})$$

• Can be calculated from a Kubo formula (in linear response)

$$\sigma^{ij}(\omega, \vec{k}) = rac{G_R^{ij}(\omega, \vec{k})}{i\omega}$$

Various limits are of experimental importance

$$\vec{k} \to 0 \to \sigma^{ij}(\omega, T) \to AC$$
 conductivity  
 $\omega \to 0$  and  $\vec{k} \to 0 \to \sigma^{ij}(T) \to DC$  conductivity

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Motivation

### Conductivity & Charged Excitation Spectra II

The linear response calculation at  $\vec{k} = 0$  can, in holography, be recast as a Schrödinger problem [Horowitz, Roberts 2009]. In the backreacted background

$$ds^{2} = -D(r)dt^{2} + B(r)dr^{2} + C(r)(dx_{i}dx^{i}), \quad A'_{t} = q \frac{\sqrt{D(r)B(r)}}{Z(\phi)C(r)^{\frac{p-1}{2}}}$$

we must turn on perturbations

$$A_i = a_i(r)e^{i(\omega t)}, g_{ti}(r, t) = z_i(r)e^{i\omega t}$$

From Einstein and Maxwell, we obtain a second order equation

$$\partial_r \left( ZC^{\frac{p-3}{2}} \sqrt{\frac{D}{B}} a_i' \right) + ZC^{\frac{p-3}{2}} \left( \sqrt{\frac{B}{D}} \omega^2 - \frac{q^2 \sqrt{DB}}{ZC^{p-1}} \right) a_i = 0$$

It can be mapped to a Schrödinger equation

$$-\frac{d^2\Psi}{dz^2} + V_{eff}\Psi = \omega^2\Psi \,, \quad V_{eff} = \frac{q^2D}{ZC^{p-1}} + \frac{1}{4}\left(\frac{\partial_z\bar{Z}}{\bar{Z}}\right)^2 + \frac{1}{2}\partial_z\frac{\partial_z\bar{Z}}{\bar{Z}} \,, \quad \bar{Z} = ZC\frac{p-3}{2}$$

via a coordinate and wave function redefinition

$$rac{dz}{dr} = \sqrt{rac{B}{D}} \;, \quad a_i = rac{\Psi}{\sqrt{Z}} \;.$$

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# Conductivity & Charged Excitation Spectra III

The frequency dependent conductivity is given by [Roberts+Horowitz (2009), Kachru et.al. (2009)]

$$\sigma(\omega) = \frac{1 - \mathcal{R}}{1 + \mathcal{R}} - \frac{i}{2\omega} \frac{\dot{Z}}{Z} \Big|_{\text{boundary}}$$

The second term always vanishes in our cases. Several situations can now occur:

- $V \rightarrow 0$  in IR: Conductor, continuous spectrum
- $V \rightarrow V_0$  in IR: Conductor with continuous spectrum and mass gap
- V → ∞ in IR, but not UV: Insulator with continuous spectrum (for p = 3 and Δ ≥ 1, with gap for Δ = 1)
- $V \rightarrow \infty$  in both IR and UV: Insulator with discrete spectrum and mass gap (confined phase) (for p > 3 or p = 3 and  $\Delta < 1$ )

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Motivation: Strange Metallic Behaviour and Effective Holographic Theories for Condensed Matter Systems

#### Uncharged Solutions

- 3 Exact Charged Solutions •  $\gamma \delta = 1$ •  $\gamma = \delta$
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#### **Uncharged Solutions**

## Uncharged Solutions

Two Solutions with "good" IR singularities: ۲

[Gubser hep-th/0002160]

Small Planar Black Holes:

$$ds^{2} = r^{\frac{4}{(p-1)\delta^{2}}} \left(-f(r)dt^{2} + d\bar{x}^{2}\right) + \frac{dr^{2}}{f(r)}$$

$$f(r) = 1 - \left(\frac{r_{0}}{r}\right)^{\frac{2p}{(p-1)\delta^{2}} - 1}, \quad e^{\delta\Phi} = \frac{|\Lambda|\delta^{4}}{\frac{2p}{(p-1)} - \delta^{2}}r^{2}$$

$$ure: \qquad \left[ 4\pi T = \left(\frac{2p}{(p-1)\delta^{2}} - 1\right)r_{0}^{\frac{2}{(p-1)\delta^{2}} - 1} \right]$$

$$\frac{2}{p-1}: T \xrightarrow{r_{0} \to 0} 0$$

$$2 \to T \xrightarrow{r_{0} \to 0} p-1$$

- Temperatu
- $\rightarrow$  A)  $\delta^2 < \frac{1}{2}$
- $\rightarrow$  B)  $\delta^2 = \frac{2}{p-1}$ :  $T \stackrel{r_0 \rightarrow v}{\rightarrow} \frac{p-1}{4\pi}$ ightarrow C)  $\delta^2 > rac{2}{D-1}$ :  $T \stackrel{r_0 
  ightarrow 0}{
  ightarrow} \infty$
- Thermal Gas Solution:  $f(r) \equiv 1$
- AdS UV completion:  $V_{full}(\Phi) \stackrel{IR}{\sim} e^{-\delta\Phi}$ , AdS Minimum in UV  $\rightarrow$  Large Planar Black Hole Branch for large  $r_0$  with  $T \sim r_0$

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**EHTCMS** 

# **Uncharged Solutions**



- A)  $\delta^2 < \frac{2}{p-1}$ : continuous PTs at  $T = 0_+$ , continuous & gapless spectrum
- B) δ<sup>2</sup> = <sup>2</sup>/<sub>p-1</sub>: continuous PT at T<sub>min</sub> = <sup>p-1</sup>/<sub>4π</sub> depending on subleading terms in V(Φ) (n<sup>th</sup> order or KT) [U. Gursoy's talk], continuous & gapped
- C)  $\delta^2 > \frac{2}{\rho-1}$ : 1st order deconfinement transition, Insulator/Conductor
- D)  $\frac{2p}{p-1} > \delta^2 > \frac{2(p+2)}{3(p-1)}$  Graviton fluctuation problem unacceptable
- E)  $\delta^2 > \frac{2p}{p-1}$  : Gubser's criterion violated

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# Uncharged Solutions: AC/DC Conductivity

• Probe DBI calculation yields  $(g^S = e^{k\Phi}g^E, \kappa = \frac{4\pi}{\left(\frac{2p}{(2-1)\kappa^2}-1\right)})$ 

$$\sigma_{xx} = e^{-k\phi_0}(\kappa T)^{\frac{2k(\rho-1)\delta+4}{(\rho-1)\delta^2-2}} \sqrt{\langle J^t \rangle^2 + e^{2(\gamma+k)\phi_0}(\kappa T)^{\frac{4(\rho-1)[1+(\gamma+k)\delta]}{2-(\rho-1)\delta^2}}}$$

- At low densities :  $\rho \sim T^{\frac{2(\rho-1)\gamma\delta+2(\rho-3)}{(\rho-1)\delta^2-2}}$ ,  $\gamma_{\text{linear}} \equiv \frac{\delta^2 \frac{2(\rho-2)}{(\rho-1)}}{2\delta}$  (indep. of *k*)
- At high densities :  $\rho \sim T^{\frac{2k(p-1)\delta+4}{2-(p-1)\delta^2}} \langle J^t \rangle$  (cannot be linear in the range  $\delta^2 < 2(p+2)/(3(p-1))$ )

 $\clubsuit$  Interpolation between  $\ \rho \sim {\it T}$  (NFL ) and  $\rho \sim {\it T}^2$  (FL ) for  $\delta = -{\it k}$ 

• Dragging string :  $\rho \simeq \frac{T_{t}g_{xx}^{\mathcal{E}}(r_{h})e^{k\Phi(r_{h})}}{J^{t}}, \quad S \sim g_{xx}^{\mathcal{E}}(r_{h})^{\frac{p-1}{2}}$  $\Rightarrow \left[\rho \sim S \sim C_{V} \quad \text{for} \quad p = 3, k = 0\right]$ 

• AC conductivity at zero temperature ( $\omega \gg T,\, p=3$ ): [Kachru et. al. (2009)]

$$\sigma \sim \omega^n, n = -\frac{2}{3} \text{ for } \gamma = \left\{ \frac{\delta^2 - 1}{3\delta}, \frac{2(\delta^2 - 1)}{3\delta} \right\}$$

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 $\gamma \delta = 1$ 

### An exact charged solution for $\gamma\delta=1$ [Charmousis et.al. 0905.3337]

$$\begin{split} \mathrm{d}s^2 &= -e^{\Phi/\delta} \, V(r) \frac{\mathrm{d}t^2}{r^2} + e^{\delta\Phi} \frac{\mathrm{d}r^2}{V(r)} + r^{3-\delta^2} e^{\frac{\delta^2-1}{2\delta}\Phi} (\mathrm{d}x^2 + \mathrm{d}y^2) \\ V(r) &= r^2 - 2mr^{\delta^2-1} + \frac{(1+\delta^2)q^2}{4\delta^2(3-\delta^2)^2} r^{2\delta^2-4} \,, \quad e^{\Phi} = r^{2\delta} \left[ 1 - \left(\frac{r_-}{r}\right)^{3-\delta^2} \right]^{\frac{4\delta(\delta^2-1)}{(3-\delta^2)(1+\delta^2)}} \\ (r_{\pm})^{3-\delta^2} &= m \pm \sqrt{m^2 - \frac{(1+\delta^2)q^2}{4\delta^2(3-\delta^2)^2}} \,, \quad A_t = \mu - \frac{q}{(3-\delta^2)r^{3-\delta^2}} \end{split}$$

- For Black Hole Solutions:  $\Lambda < 0$  and  $\delta^2 < 3$  (cf. Gubser)
- AdS completable for all  $\delta$
- AdS-Schwarzschild Limit:  $\delta \rightarrow 0, q \rightarrow 0$

 $\gamma \delta = 1$ 

# $\gamma \delta = 1$ : Solution Branches

• Temperature: 
$$T = (3 - \delta^2) r_+^{1 - \delta^2} \left[ 1 - \frac{4(1 + \delta^2)Q^2}{\delta^2 (3 - \delta^2)^2 r_+^{2(3 - \delta^2)}} \right]^{1 - \frac{2(\delta^2 - 1)^2}{(1 + \delta^2)(3 - \delta^2)}}$$

- A)  $\delta^2 \leq 1$ : T = 0 at Extremality
- B)  $1 < \delta^2 < 1 + \frac{2}{\sqrt{3}}$ : Two branches (SBH/LBH) + AdS-LBHs
- C)  $1 + \frac{2}{\sqrt{3}} \le \delta^2 < 3$ : *T* diverges at extremality, but AdS-LBHs take over



#### $\gamma \delta = 1$

# $\gamma \delta = 1$ : Phases and Low-Temperature Scaling



#### • Finite Q, $T \rightarrow 0_+$ :

- 2<sup>nd</sup> order phase transition for  $1 \frac{2}{\sqrt{5}} < \delta^2 < 1 + \frac{2}{\sqrt{5}}$
- 3<sup>rd</sup> or higher order otherwise
- No residual entropy at zero temperature

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# $\gamma \delta = 1$ : Transport Properties

• In the first two regimes  $0 \le \delta^2 \le 1 + \frac{2}{\sqrt{3}}$ , the AC conductivity is



DC resistivity from the dragging string calculation: At low temperatures

$$\rho_{\text{leading}} \sim \frac{T_f}{Jt} \left(\frac{q}{\ell}\right)^{\frac{2\delta(\delta(3-\delta^2)+(1+\delta^2)k)}{1+6\delta^2-3\delta^4}} \left(\ell T\right)^{\frac{2(\delta^2-1)(\delta^2-1+2k\delta)}{1+6\delta^2-3\delta^4}}$$

This is linear in temperature at k = 0 for  $\delta^2 = 1 \pm \frac{2}{\sqrt{5}}$ . We thus find

 $m{S} \sim m{C}_{m{Q}} \sim m{T} \sim 
ho_{m{D}m{C}}$ 

[Loram et. al. PRL 71 11 (1993)]

 The Karch/O'Bannon probe DBI calculation does not make sense in backreacted backgrounds.

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#### $\gamma = \delta$

# $\gamma = \delta$ : RN-Like Charged Dilaton Black Holes

$$\begin{aligned} ds^2 &= -V(r)dt^2 + e^{\delta\Phi} \frac{dr^2}{V(r)} + r^2(dx^2 + dy^2) \\ V(r) &= r^2 - 2mr^{\delta^2 - 1} + \frac{q^2}{4(1 + \delta^2)r^2}, \quad e^{\Phi} = r^{2\delta}, \quad r_e^4 = \frac{q^2}{4(3 - \delta^2)} \\ A_t &= \mu - \frac{q}{(1 + \delta^2)r^{1 + \delta^2}} \end{aligned}$$

- For Black Hole Solutions:  $\Lambda < 0$  and  $\delta^2 < 3$  [Charmousis et.al. 0905.3337]
- AdS completable for all δ
- Coincides with  $\gamma \delta = 1$  for  $\gamma = \delta = 1$
- AdS-RN Limit:  $\delta, \gamma \rightarrow 0$
- <u>Residual Entropy</u>:  $S = r_e^2 \simeq r_e^2 + \frac{T}{2} \frac{r_e^{1+\delta^2}}{3-\lambda^2} + \dots$

$$C_Q \simeq rac{T}{2} rac{r_e^{1+\delta^2}}{3-\delta^2} + \dots, \quad 
ho_{DC} \simeq 
ho_0 + AT + \dots$$

Image: A matrix

Exact Charged Solutions  $\gamma = \delta$ 

## $\gamma = \delta$ : Solution Branches and Phase Diagram



•  $Q > 0, T \rightarrow 0_+$ : EBHs dominate, no phase transition

•  $T > 0, Q \rightarrow 0_+$ :

- $\delta^2 < 1$  : 2<sup>nd</sup> order phase transition to NBHs
- $\delta^2 > 1$ : 1<sup>st</sup> order phase transition to EBHs
- AC Conductivity :  $\sigma \sim \omega^2 \; \forall \delta$

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- Motivation: Strange Metallic Behaviour and Effective Holographic Theories for Condensed Matter Systems
- 2 Uncharged Solutions
- 3 Exact Charged Solutions •  $\gamma \delta = 1$ •  $\gamma = \delta$
- 4 Near-Extremal Scaling Solutions
- 5 Conclusions & Remarks

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# **Near-Extremal Scaling Solutions**



• Near-Extremal Approximations to  $\gamma \delta = 1$  and  $\gamma = \delta$ 

• Lifshitz cases:  $\delta = 0, z = 1 + \frac{4}{\gamma^2}$ 

[M. Taylor, 0812.0530]

- Entropy:  $S \sim (2m)^{\frac{2(\gamma-\delta)^2}{WU}} \rightarrow 0$  except for  $\gamma = \delta$
- Multitude of continuous  $T \rightarrow 0_+$  Phase Transitions

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**Near-Extremal Scaling Solutions** 

## Spin 2 and Spin 1 constraints



Reliability constraints on the spin-2 fluctuations for p = 3 and p = 4. The blue region depicts the part of the  $(\gamma, \delta)$  plane which satisfies Gubsers constraint. The yellow-brownish and purple regions are the allowed regions from spin 2 fluctuations. The purple region furthermore is thermodynamically unstable, and will be stabilised after AdS completion. This touches the Gubser allowed region for p = 3 in the point  $(\pm 1, \mp 1)$ , and has a small overlap for p = 4.

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Near-Extremal Scaling Solutions

#### The extremal AC conductivity in p = 3



Contour plot of the scaling exponent *n* in the  $(\gamma, \delta)$  upper half plane for p = 3 Contours correspond to n = 1.52, ..., 8.36, starting with n = 1.52 in the upper right corner and increasing in steps of 0.76. The black solid line  $\gamma = \delta$  is n = 2, and brighter colors correspond to larger *n*. The yellow region is thermodynamically unstable. The scaling exponent diverges to  $+\infty$  along the dashed black line.

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#### Mott-like States for p = 3



Left: The region on the  $(\gamma \delta)$  plane where the IR black holes are unstable and the c > 0. Here the extremal finite density system has a mass gap and a discrete spectrum of charged excitations, when  $\Delta < 1$ . This resembles a Mott insulator and the figure provides the Mott insulator "islands" in the  $(\gamma, \delta)$  plane. Right: The region where the IR black holes are unstable, and c < 0. In this region the extremal finite density system has a gapless continuous spectrum at zero temperature. In both figures the horizontal axis parametrizes  $\gamma$ , whereas the vertical axis  $\delta$ .

# Near-Extremal DC Conductivity

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For massive charge carriers the drag force calculation yields (p = 3)

$$ho \sim T^m$$
,  $m = rac{4k(\delta - \gamma) + 2(\delta - \gamma)^2}{4(1 - \delta(\delta - \gamma)) + (\delta - \gamma)^2}$ 

• The exponent becomes unity for two values of  $\gamma$ 

$$\gamma_{\pm} = 3\delta + 2k \pm 2\sqrt{1 + (\delta + k)^2}$$

• For a non-dilatonic scalar, k = 0 and the temperature dependence of the entropy and the resistivity are the same in p = 3. Therefore, the entropy also scales linearly with T.

• For the Lifshitz solutions, we must take  $\delta = 0$  and  $\gamma = -\sqrt{\frac{4}{(z-1)}}$ . In this case we obtain that  $m = \frac{2+k\sqrt{4(z-1)}}{z}$ . When k = 0 this is in agreement with [Hartnoll+Polchinski+Silverstein+Tong]

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# **Domination of Charged Entropy**





In the shaded region, the entropy at finite charge density dominates the one for zero density at very low temperatures. In the rest of the  $(\gamma, \delta)$  diagram, the comparison cannot be made as both entropies are expected to be of  $\mathcal{O}(1)$  instead of  $\mathcal{O}(N^2)$ . The vertical axis represents the value of  $\delta$ , while the horizontal axis the value of  $\gamma$ .

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# **Conclusions & Remarks**

- Einstein-Maxwell-Neutral Scalar Theory is a good laboratory for strange metallic behaviour
- Describe: Interactions between a charged carrier sector and a leading relevant operator ("glue")  $\rightarrow$  substrate in the High  $T_c$  normal state
- Interesting phase structures of uncharged and (fully backreacted) charged solutions of our system
- Uncharged solutions: 1<sup>st</sup> order (De)confinement phase transitions appear with discrete and gapped spectra [Kiritsis, Gursoy, Nitti, Mazzanti on AdS/QCD]
- Charged solutions: Generically conductors, appearance of Mott-like insulating (confining) behaviour delayed to higher values of  $\delta$
- Well-definedness of the fluctuation problem poses interesting constraints on parameter space
- Continuous phase transitions at zero temperature (and charge)
- No residual entropy except for  $\gamma = \delta$ , charged entropy dominance
- Correlation between low-temperature scaling of entropy, (electronic) specific heat and DC resistivity observed in experiments = + = = - > = - > 

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