

Phase Structure of Strongly Coupled Gauge Theories

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Crete, May 2010

Introduction — equilibrium computations

Probe brane duals of 3+1d (& 2+1d) gauge theories

Chiral symmetry breaking dynamics

T & μ phase diagrams – chiral + meson melting transitions

First order transitions

Second order transitions

BKT transitions

Non-mean field second order transitions

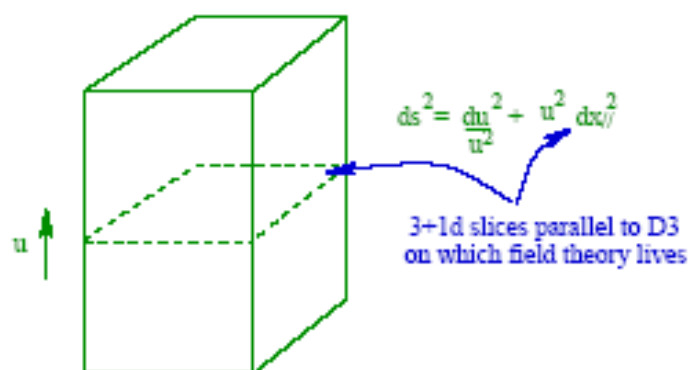
Motive:

QCD

(CM)

4d strongly coupled $\mathcal{N}=4$ SYM = IIB strings on $\text{AdS}_5 \times \text{S}^5$

Pretty well established by this point!



u corresponds to energy (RG) scale in field theory

The SUGRA fields act as sources

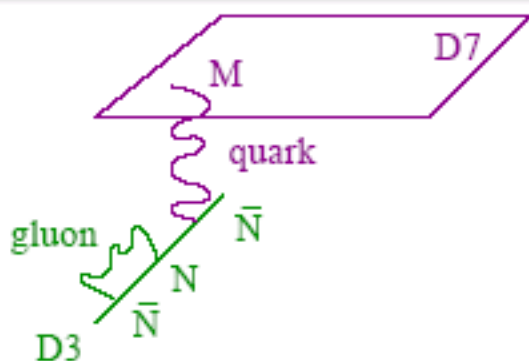
$$\int d^4x \Phi_{\text{SUGRA}}(u_0) \lambda \lambda$$

eg asymptotic solution ($u \rightarrow \infty$) of scalar

$$\varphi \simeq \frac{m}{u} + \frac{\langle \lambda \lambda \rangle}{u^3}$$

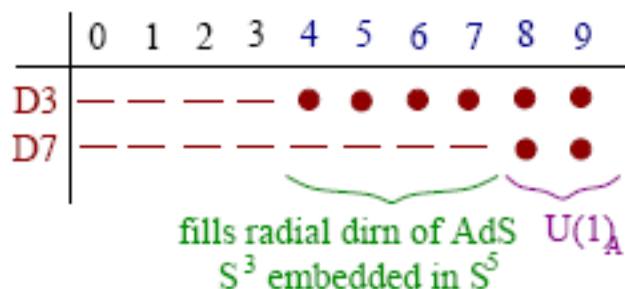
Adding Quarks

Bertolini, DiVecchia...; Polchinski, Grana; Karch, Katz...



Quarks can be introduced via D7 branes in AdS

The brane set up is



We will treat D7 as a probe - quenching in the gauge theory.

Minimize D7 world volume with DBI action

$$S_{D7} = -T_7 \int d\xi^8 \sqrt{P[G_{ab}]}, \quad P[G_{ab}] = G_{MN} \frac{dx^M}{d\xi^a} \frac{dx^N}{d\xi^b}$$

Quarks In AdS

Myers et al

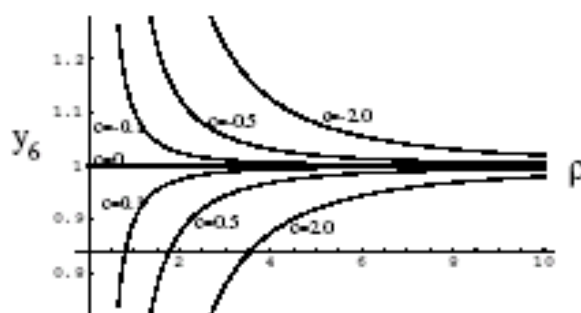
$$S_{D7} = -T_7 \int d^8\xi \epsilon_3 \rho^3 \sqrt{1 + \frac{g^{ab}}{\rho^2 + u_5^2 + u_6^2} (\partial_a u_5 \partial_b u_5 + \partial_a u_6 \partial_b u_6)}$$

EoM is:
$$\frac{d}{d\rho} \left[\frac{\rho^3}{\sqrt{1 + \left(\frac{du_6}{d\rho}\right)^2}} \frac{du_6}{d\rho} \right] = 0$$

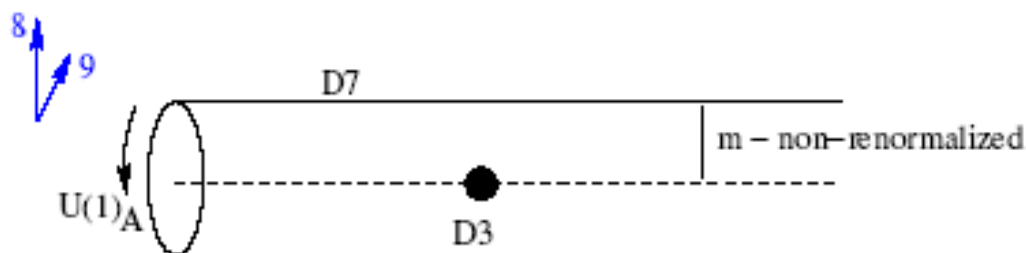
UV asymptotic solution is

$$u_6 = m + \frac{c}{\rho^2} + \dots$$

m is the quark mass, c the $\langle \bar{q}q \rangle$ condensate



In AdS regular D7 solution is flat brane



The D7 lie flat in AdS. We can consider fluctuations that describe R-chargeless mesons

$$W_6 + iW_5 = d + \delta(\rho) e^{ik \cdot x}$$

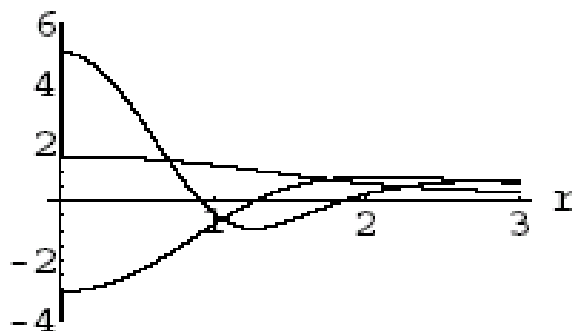
δ satisfies a linearized EoM

$$\partial_\rho^2 \delta + \frac{3}{\rho} \partial_\rho \delta + \frac{M^2}{(\rho^2 + 1)^2} \delta = 0$$

and the mass spectrum is

$$M = \frac{2d}{R^2} \sqrt{(n+1)(n+2)} \sim \frac{2m}{\sqrt{\lambda_{YM}}}$$

Tightly bound - meson masses suppressed relative to quark mass



Orthonormal wave functions

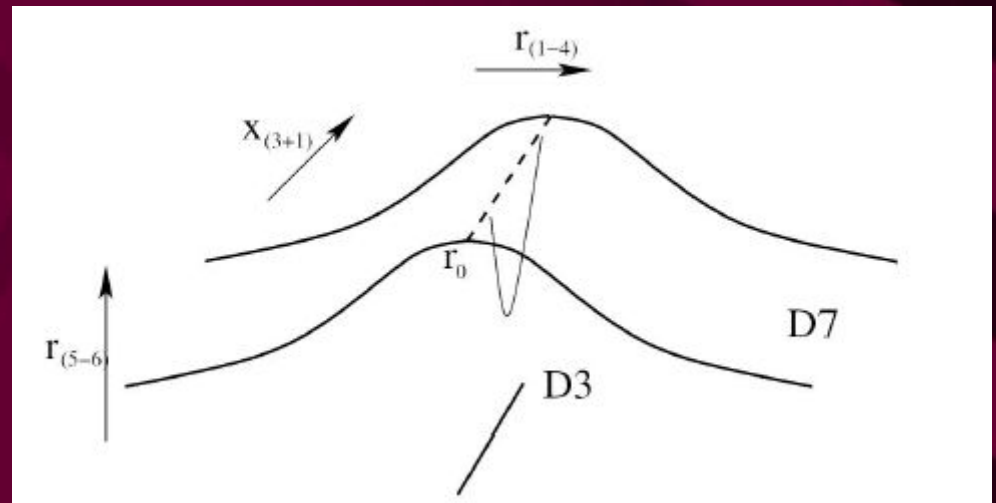
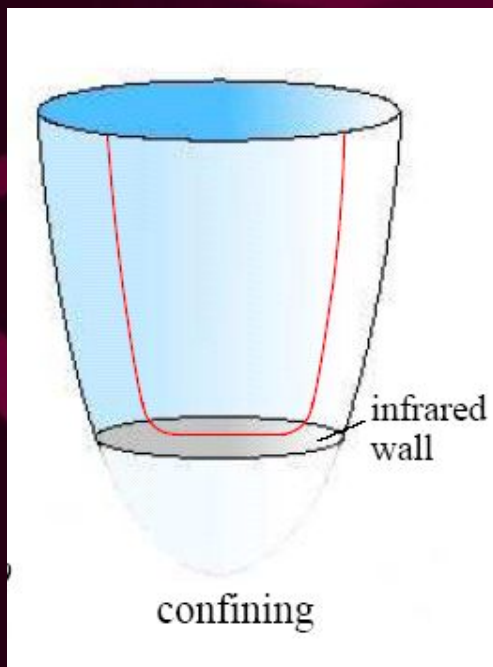
Add Confinement and Chiral Symmetry Breaking

$$ds^2 = \frac{r^2}{R^2} A^2(r) dx_{3+1}^2 + \frac{R^2}{r^2} dr^2,$$

$$A(r) = \left(1 - \left(\frac{r_w}{r}\right)^8\right)^{1/4}, \quad e^\phi = \left(\frac{1 + (r_w/r)^4}{1 - (r_w/r)^4}\right)^{\sqrt{3/2}}$$

Dilaton Flow Geometry: Gubser, Sfetsos

Here, this is just a simple, back reacted, repulsive, hard wall....



BEEGK, Ghoroku..

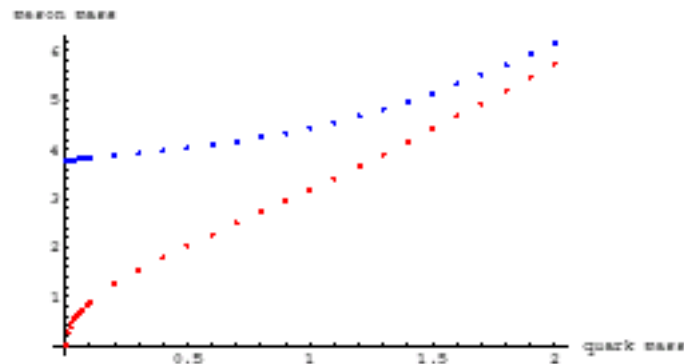
Pion Physics

Seek pion solutions of the form

$$\pi(x, r) = f(\rho)e^{ikx}, \quad k^2 = -M^2$$

$f(\rho)$ must be smooth - normalizable - at all ρ

The pion and sigma masses can thus be computed as a function of quark mass



There is a Goldstone in the massless limit.

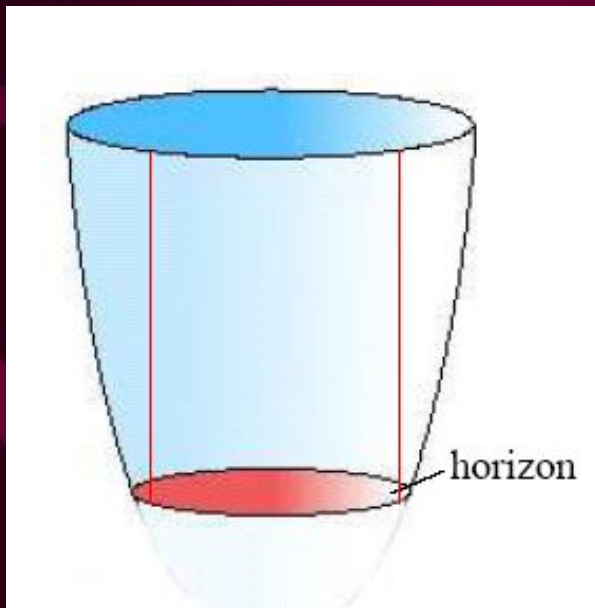
Expected \sqrt{m} behaviour

Finite T - AdS-Schwarzschild

$$ds^2 = \frac{r^2}{R^2}(-f dt^2 + d\vec{x}^2) + \frac{R^2}{r^2 f} dr^2 + R^2 d\Omega_5^2$$

where $R^4 = 4\pi g_s N \alpha'^2$ and

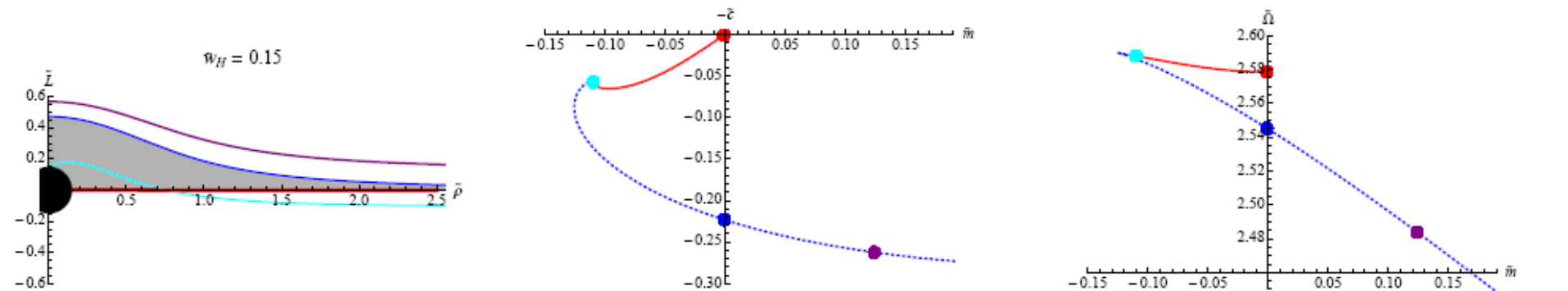
$$f := 1 - \frac{r_H^4}{r^4}, \quad r_H := \pi R^2 T.$$



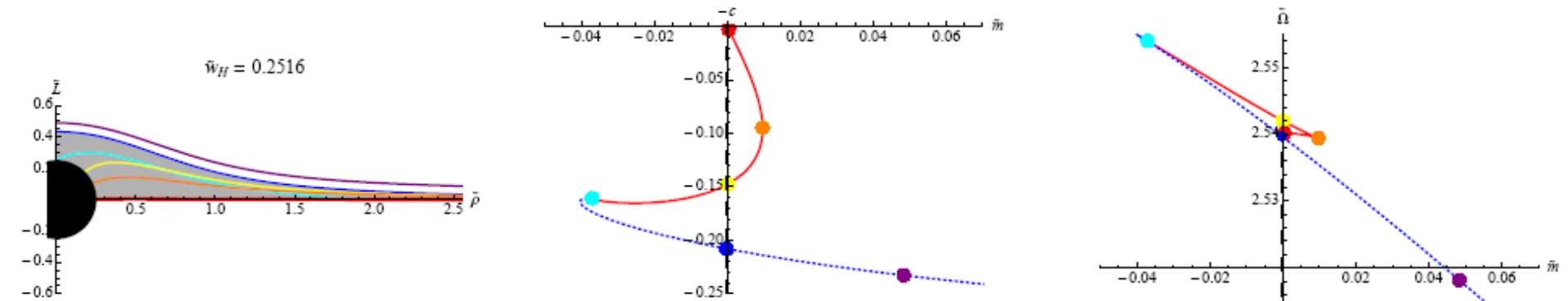
Quarks are
screened by
plasma

Asymptotically
AdS, SO(6)
invariant at all
scales... horizon
swallows
information at r_H
.... Witten
interpreted as
finite
temperature...
black hole... has
right
thermodynamic
properties...

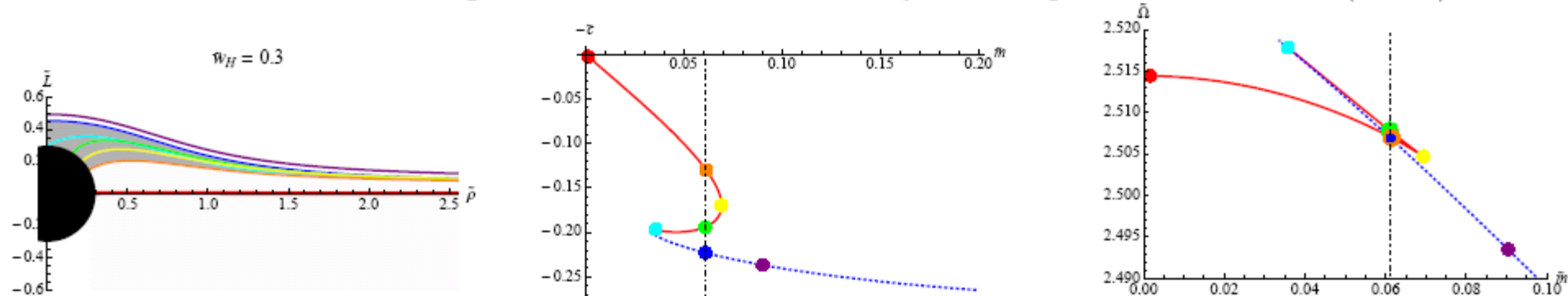
First Order Chiral Symmetry Restoration – B + T



(a) Low temperature - $\tilde{w}_H = 0.15$. Here we see chiral symmetry breaking with the blue embedding thermodynamically preferred over the red at $\tilde{m} = 0$.



(b) Transition temperature - $\tilde{w}_H = 0.2516$. This shows the point where the first order chiral symmetry phase transition occurs from the blue to the red embedding. The transition can be identified by considering Maxwell's construction (Middle) or the



(c) Above the transition - $\tilde{w}_H = 0.3$. This is the chiral restored phase with the $\tilde{m} = 0$ curve lying along the $\tilde{\rho}$ axis (red).

Quasi-normal modes & meson melting

BEEGK... Sonnenschein... Hoyos... Myers, Mateos...

Linearized fluctuations in eg the scalars on the D7 brane must now enter the black hole horizon...

Quasi-normal modes are those modes that near the horizon have only in-falling pieces...

The mass of the bound states become complex – they decay into the thermal bath...

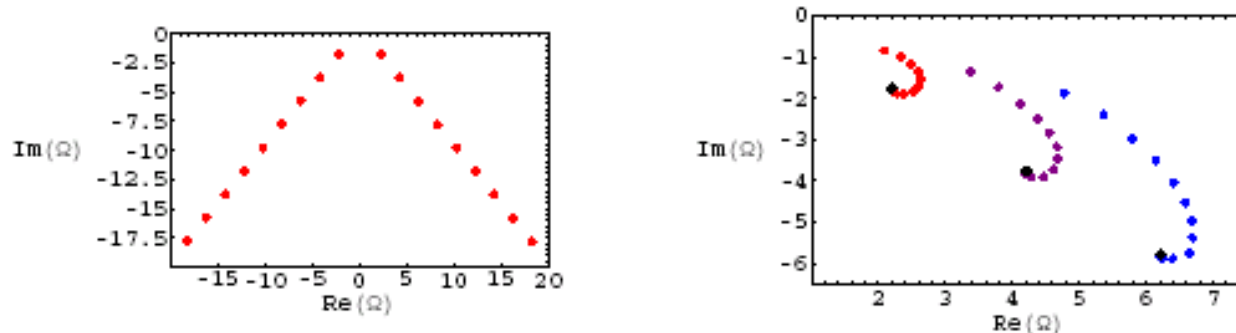
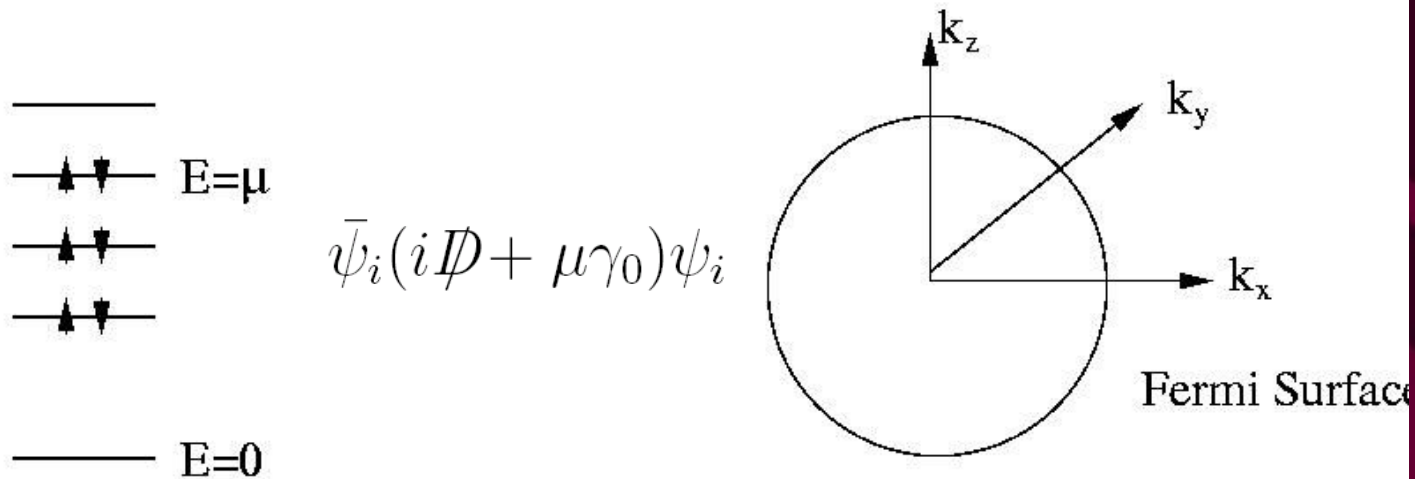


Figure 7.4: The lowest quasi-normal modes for $m_q = 0$ on the left and the three lowest quasi-normal modes for increasing m_q on the right. The black points on the right show the limiting values for $m_q = 0$.

Chemical Potential

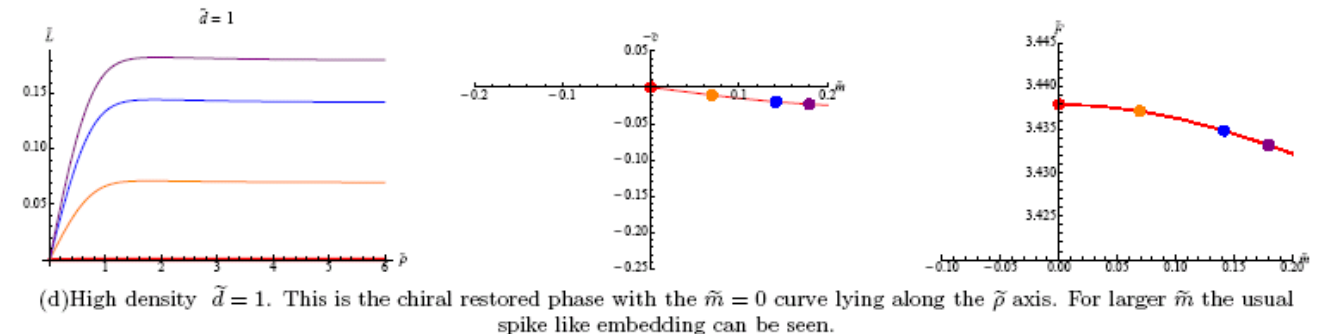
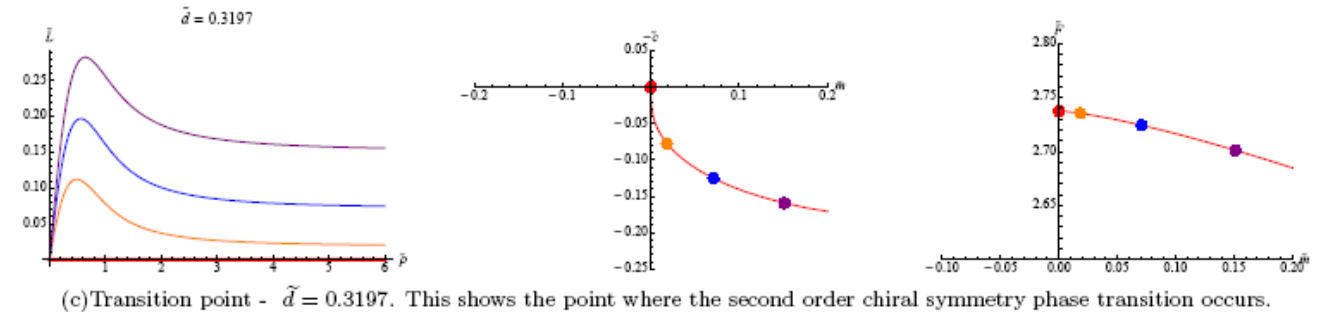
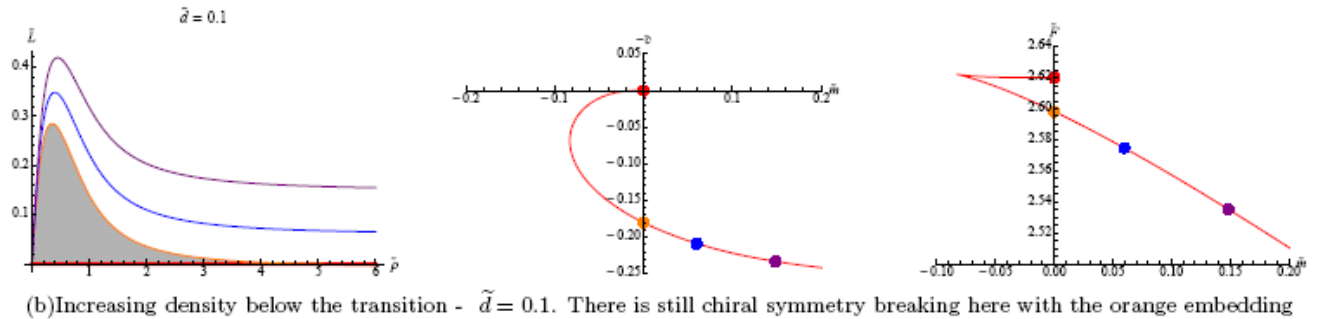
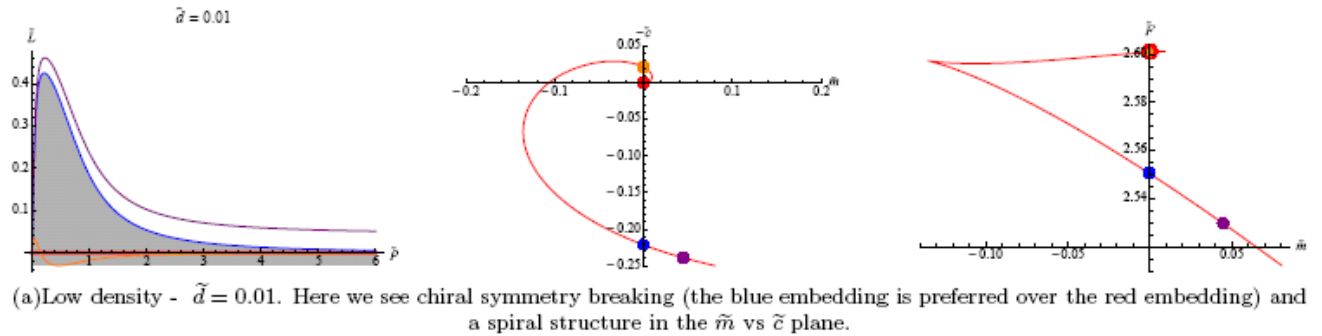
At finite density the Fermi-sea of quarks fills up to an energy called the chemical potential



We can think of μ as a background vev for the temporal component of the photon...

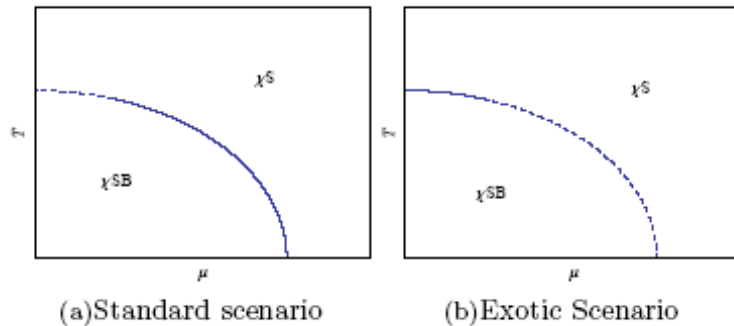
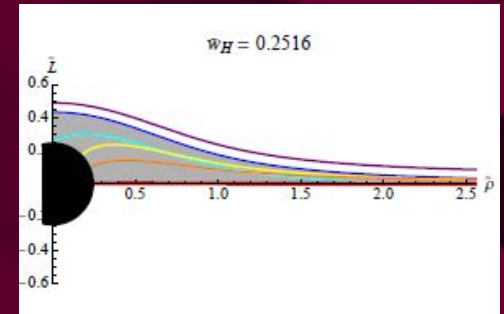
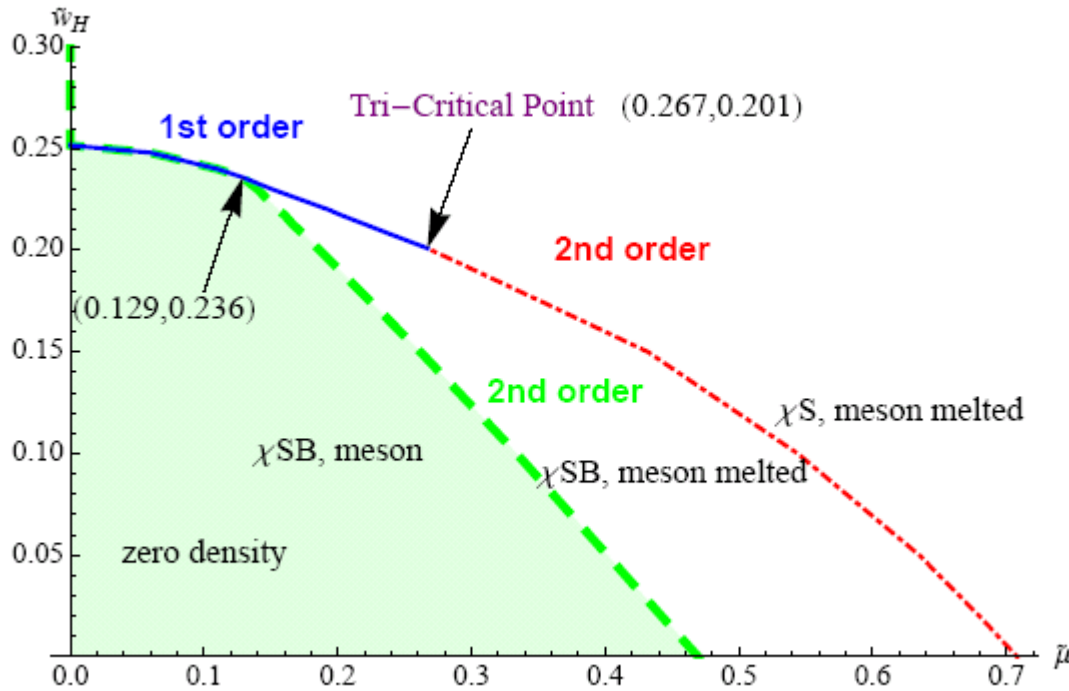
$$\bar{\psi}i(-iA^t\gamma_0)\psi \rightarrow \bar{\psi}\mu\gamma_0\psi$$

Second Order Chiral Symmetry Restoration – B + density



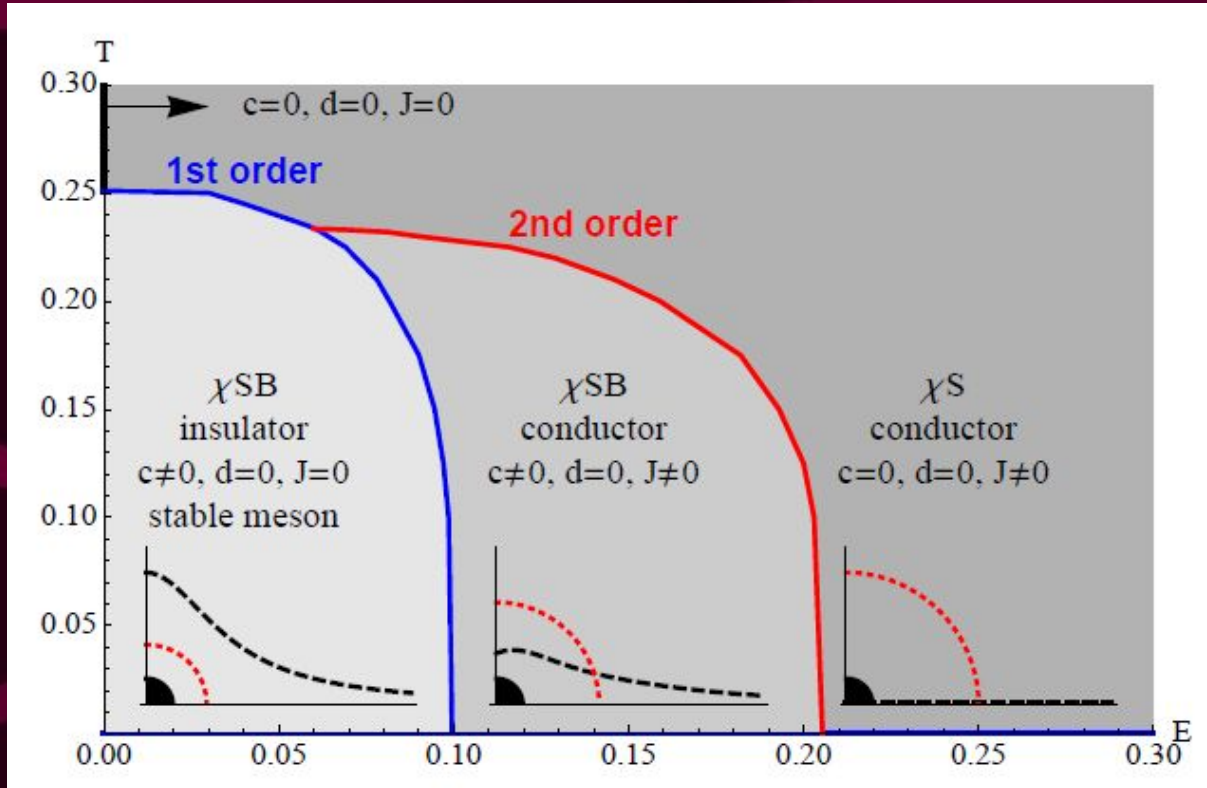
Phase Diagram for B Field Theory, $m=0$

JHEP
1003:132,2010.
 e-Print:
arXiv:1002.1885
 [hep-th]



QCD scenarios for comparison... Wilczek vs Philipsen (Lattice)

Phase Diagram for B=1,T,E Theory, m=0

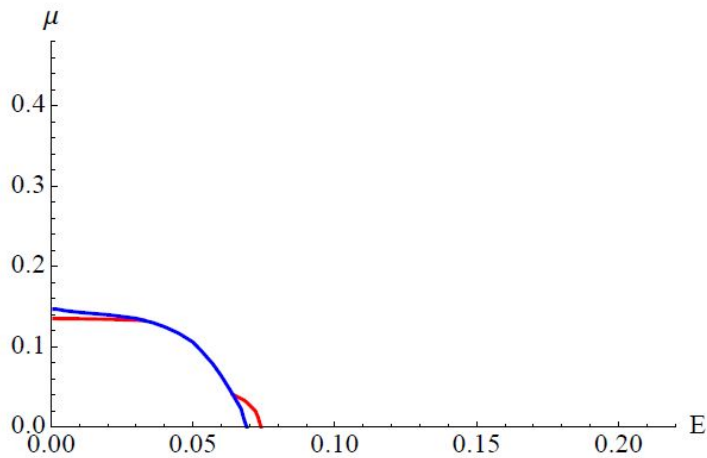


JHEP to appear
e-Print:
arXiv:1103.5627
[hep-th]

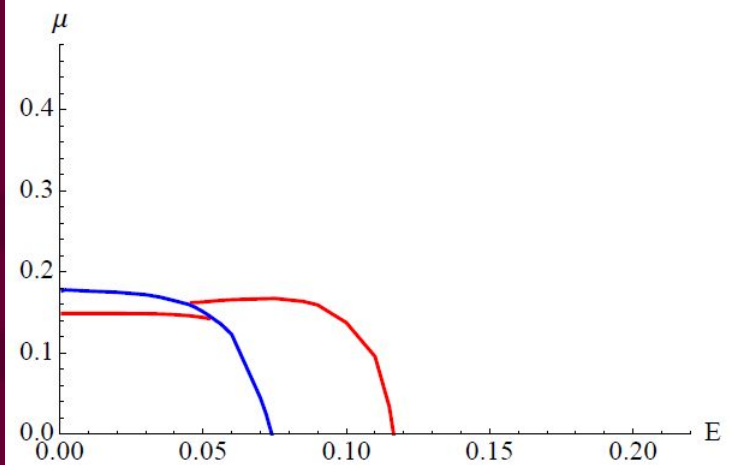
An E field tends to dissociate the mesons – when it does a current must be introduced (Karch, O'Bannon)

It opposes chiral symmetry breaking like μ .

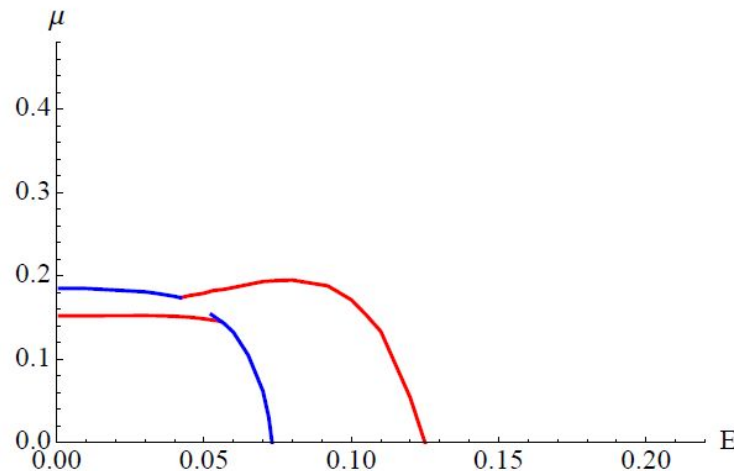
Phase Diagram for $B=1, T, E, \mu$ Theory, $m=0$



(a) $T = 0.232$



(b) $T = 0.224$



(c) $T = 0.222$

Thanks to
Astrid
Gebauer!

Second Order Mean Field Behaviour

A mean field second order transition is just an effective Landau
-Ginsberg (Higgs) Model

$$V_{\text{eff}} = V_0 + k(B - B_c)\phi^2 + \lambda\phi^4 + \dots$$

$$\phi \sim \sqrt{B - B_c}$$

A Phenomenological Operator

with Kristan Jensen

$$\Delta = 3$$

$$\tilde{S}_7 = -\mathcal{N}_7 \int d\rho \sqrt{1 + L'^2} \sqrt{\tilde{d}^2 + \rho^6 \left(1 + \frac{B^2}{w^4} + \frac{O^2}{w^{2\Delta}} \right)}.$$

holographic Berezinskii-Kosterlitz-Thouless transition

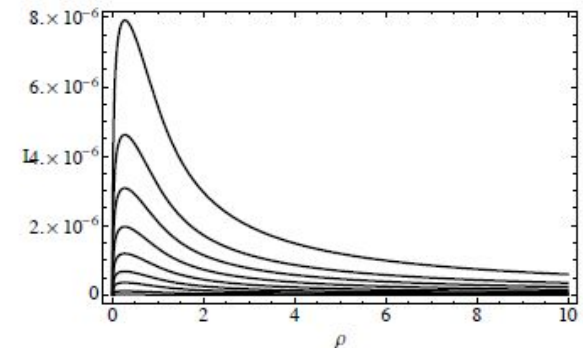
T=0 transition changes...

K. Jensen, A. Karch, D. T. Son, and E. G. Thompson, Phys. Rev. Lett. **105**, 041601 (2010), 1002.3159.

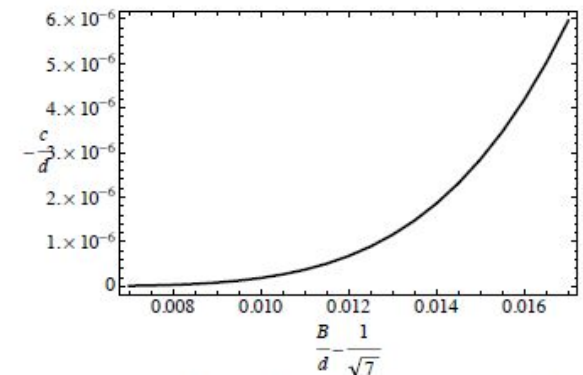
Exponential scaling of order parameter away from the transition...

D. B. Kaplan, J.-W. Lee, D. T. Son, and M. A. Stephanov, Phys. Rev. **D80**, 125005 (2009), 0905.4752.

First identified in D3/D5 system where d and B have same dimension... AdS/CM link



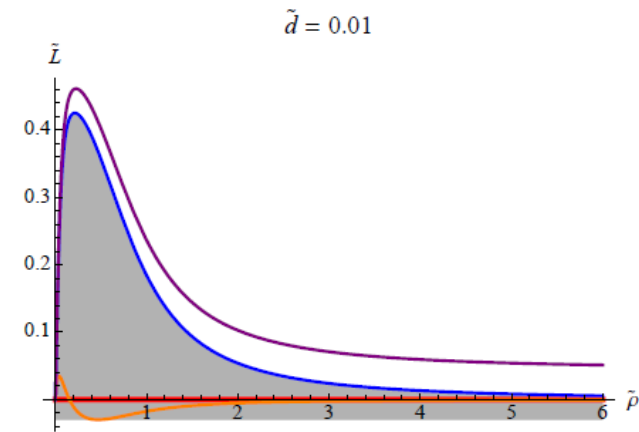
(a) The embedding L of a D5 brane in the D3 geometry for various B/\tilde{d} showing the BKT transition.



(b) A plot of the quark condensate c versus B across the D3/D5 BKT transition.

Instability of flat embedding

$$\mathcal{L}_7 = \sqrt{d^2 + O^2 + \rho^6 L'^2} + \frac{3O^2 L^2}{\rho^6 \sqrt{d^2 + O^2 + \rho^6}}$$



Small rho limit solutions:

$$\frac{L}{\rho} \sim \left(\frac{1}{\rho}\right)^\Delta$$

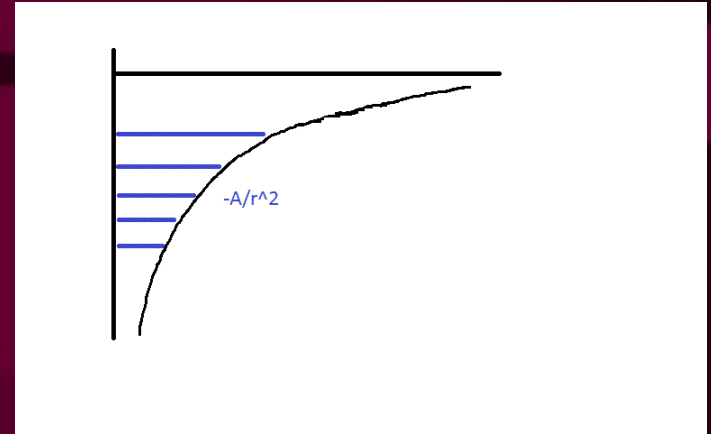
$$\Delta_{\pm} = \frac{p}{2} \pm \sqrt{\left(\frac{p}{2}\right)^2 + m^2}$$

$$m^2 = -3O^2 / (\tilde{d}^2 + O^2)$$

$$\Delta_{\text{IR}} = \frac{1 + \sqrt{\frac{\tilde{d}^2 - 11O^2}{\tilde{d}^2 + O^2}}}{2}$$

O and d enter on same footing because same dimension (not true in B case!) For fixed d raising O triggers complex [] an instability that correctly predicts the transition point...

The Schroedinger well becomes unstable ($A > 1/4$) with an infinite number of negative energy states growing from zero... leading to exponential behaviour...



Breitenlohner-Freedman (BF)

In our analysis we use the results for a scalar in AdS_{p+1} : The solution of the equation of motion is

$$\frac{L}{\rho} \sim \left(\frac{1}{\rho}\right)^\Delta \quad (11)$$

$$\Delta_{\pm} = \frac{p}{2} \pm \sqrt{\left(\frac{p}{2}\right)^2 + m^2} . \quad (12)$$

and the Breitenlohner-Freedman (BF) bound [65] is given by $-p^2/4$

$$AdS_2, m_{BF}^2 = -1/4$$

0+1d theory rules
IR?

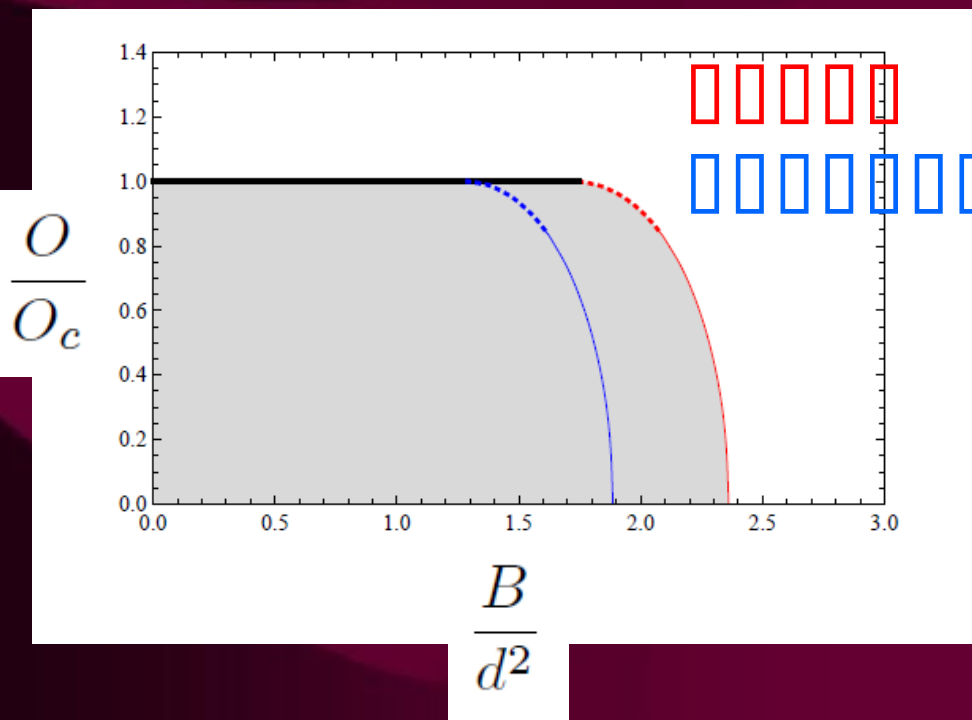
Small rho limit mass

$$m^2 = -3O^2 / (\tilde{d}^2 + O^2)$$

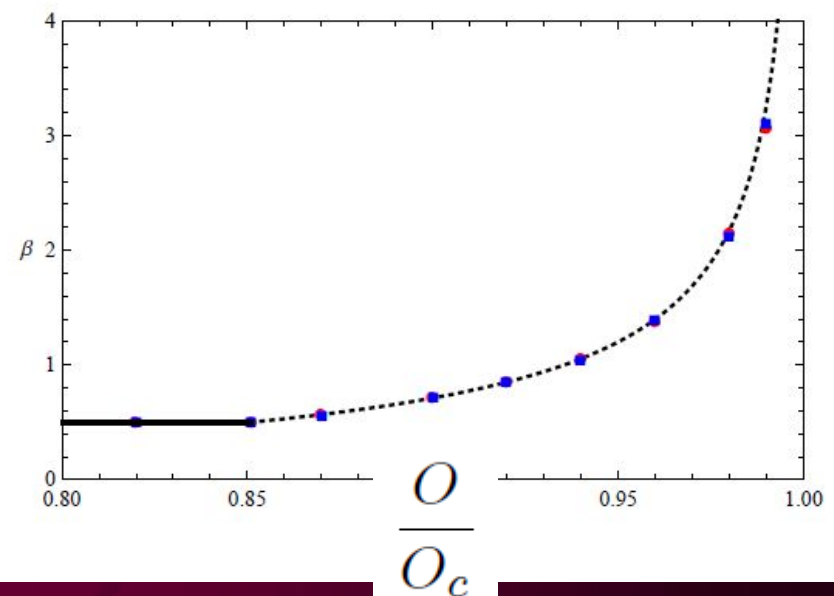
From Mean-Field 2nd Order to BKT

$$S_7 = \int d\rho \sqrt{1 + L'^2} \sqrt{d^2 + \rho^6 \left(1 + \frac{B^2}{w^4} + \frac{O^2}{w^6} \right)}$$

O+d triggers BKT.... B+d is second order mean-field... what about O+B+d:



$$c \sim (B - B_c)^\beta$$



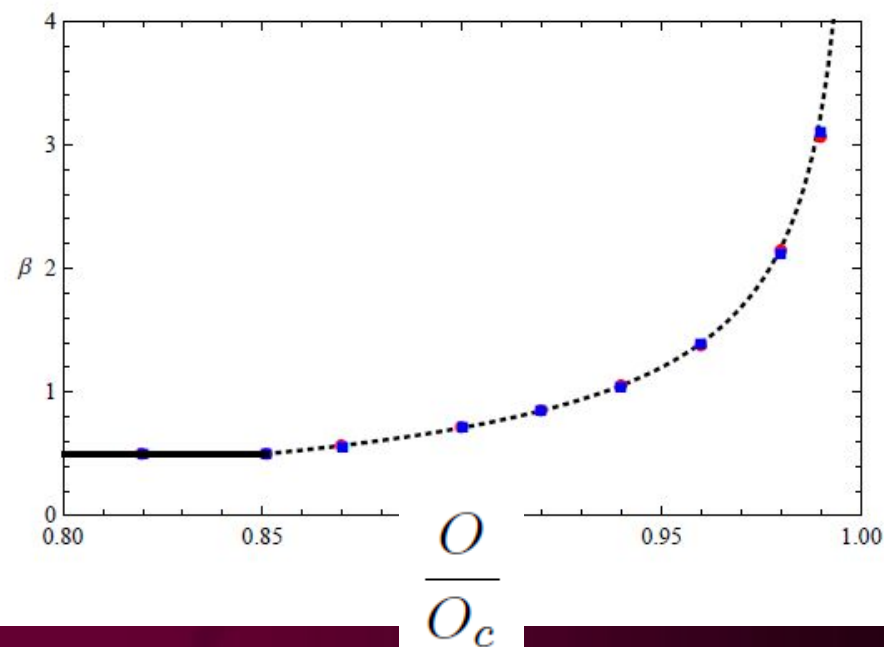
Effective Theory

$$V_{\text{eff}} = V_0 + k(B - B_c)\phi^2 + \lambda\phi^4 + \alpha_{\text{IR}}\phi^{p/(p-\Delta_{\text{IR}})}$$

$$\phi \sim (B - B_c)^\beta \equiv (B - B_c)^{\frac{p-\Delta_{\text{IR}}}{2\Delta_{\text{IR}}-2}}$$

$$\Delta_{\text{IR}} < \Delta_c = 3/4p$$

$$\beta = \frac{1}{2} \left(\sqrt{\frac{d^2 + O^2}{d^2 - 11O^2}} - 1 \right)$$



Comments

Lots of complex phase structure with $B, E, \mu, T, O!$

Non mean-field quantum critical points...

A full effective theory showing role of 0+1d IR... (pretty cool)

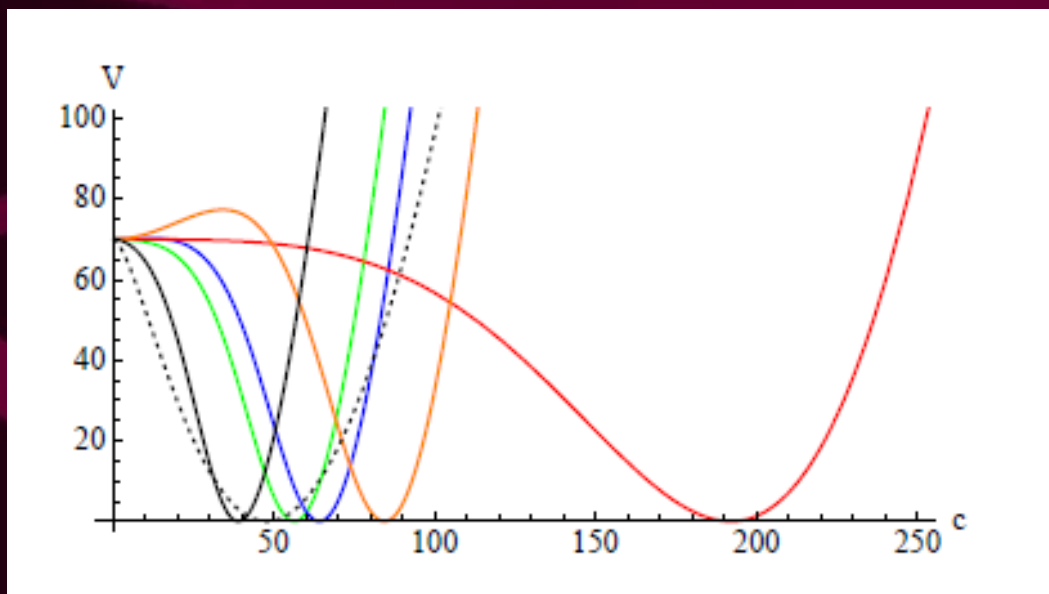
What's it good for?

Non-equilibrium Dynamics At The Transitions

AdS/CFT lets us solve full time dependence of non-equilibrium configurations... some sample ideas...

Seeking Inflating Gauge Theories

Nick Evans, James French, Keun-Young Kim, arXiv:1009.5678 [hep-th]



2nd order transition
leaves us at $c=0$ in
 $T=0$ V

Slow roll to $V=0$ with
+ve cosmological
constant \rightarrow inflation

Scalars are unnatural

An inflating strongly
coupled gauge theory?

Watching the roll with a B field

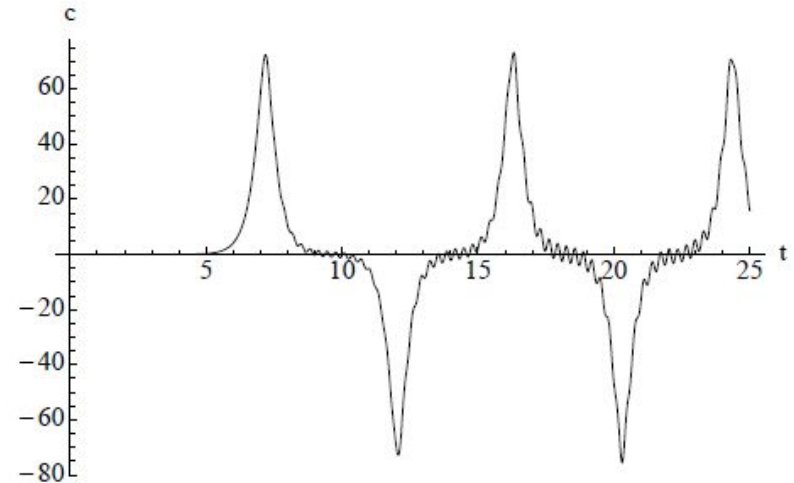
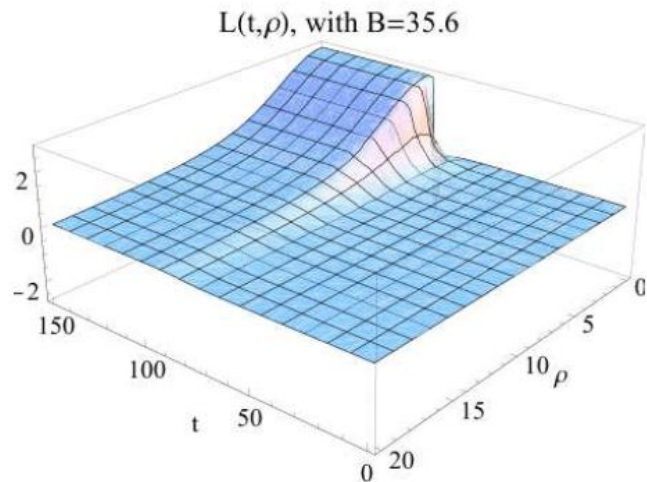
$$\mathcal{L} = -\tilde{T}_7 R^4 \int d\rho \beta \rho^3 \sqrt{\frac{g^{00} \dot{L}^2}{(\rho^2 + L^2)^2} + 1 + L'^2}$$

$$\beta = \sqrt{1 + \frac{B^2}{(\rho^2 + L^2)^2}}$$

$$ds^2 = -dt^2 + a(t)^2 \left(\frac{dr^2}{1 - kr^2} + r^2 (d\theta^2 + \sin^2 \theta d\phi^2) \right)$$

$$L(\rho, t = 0) = 0, \quad \dot{L}(\rho, t = 0) = v e^{-\rho^2}$$

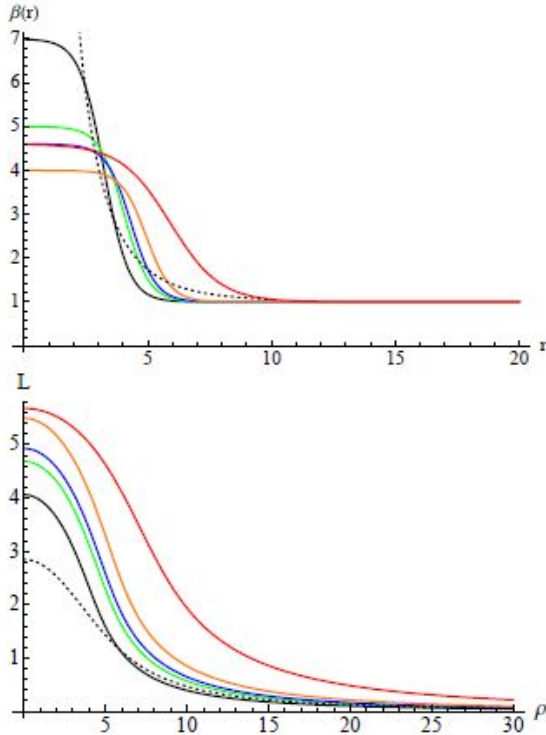
Probe limit – inflation
slow relative to N=4
glue dynamics.. Or
just a model...



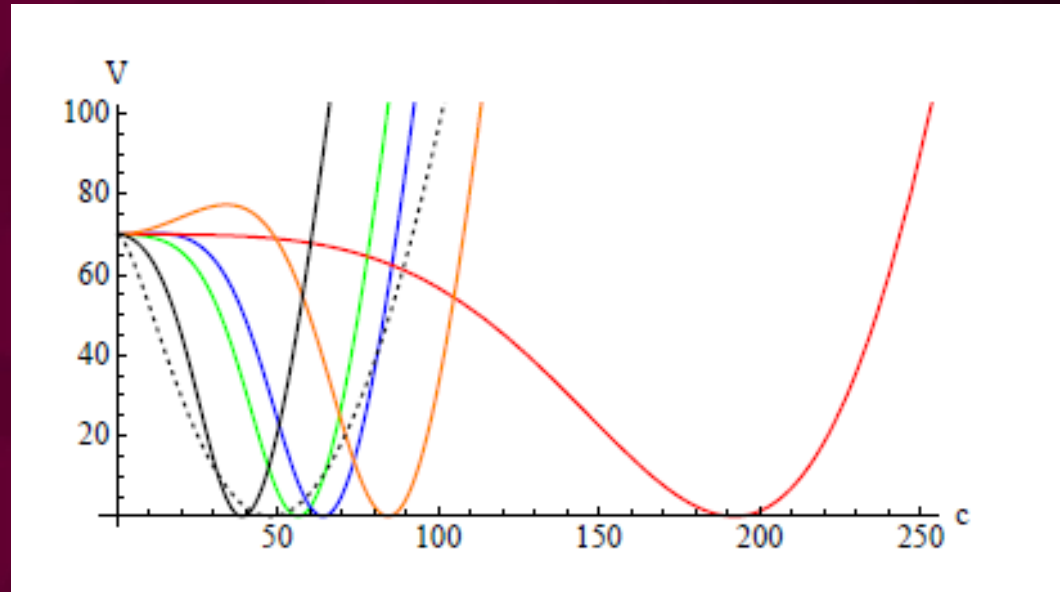
How should the coupling run to favour inflation?

$$e^{\Phi} = g_{\text{YM}}^2(r^2) = g_{\text{UV}}^2 \left(A + 1 - A \tanh [\Gamma(r - \lambda)] \right)$$

Cf AdS/QCD

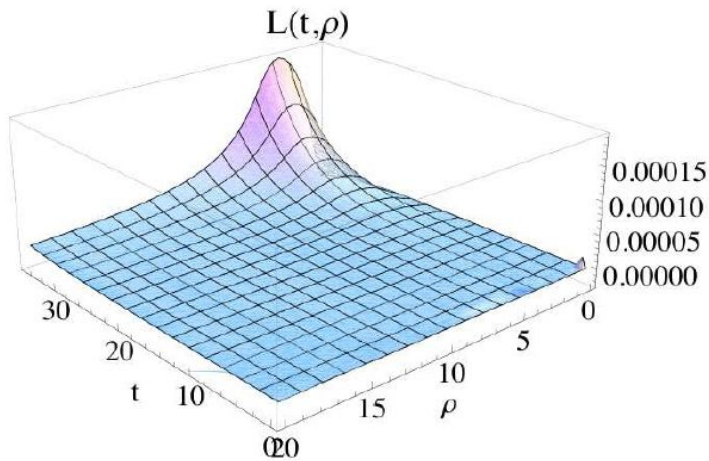


	Γ	A	λ
Black	1	3	3.240
Green	1	2	4.045
Blue	1	1.8	4.325
Orange	1	1.5	4.940
Red	0.5	1.8	5.882
Black(Dotted)	$B = 35.6$	-	-



Walking leads to a gap between \square and $\nabla \dots$

Early time rolling is around the dilaton kink.... Acceleration involves holographic behaviour...



$$\ddot{L}(\rho) \sim \left(\lambda^3 \frac{A\Gamma}{A+1} L(\lambda) + \lambda^4 L''(\lambda) \right) \delta(\rho - \lambda)$$

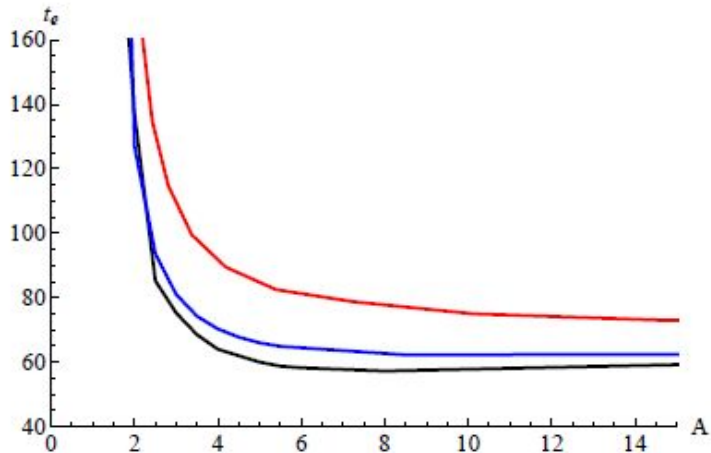


FIG. 11: t_e against A for a sequence of values of Γ s: 1.5(Black), 1(Blue), 0.5(Red). $v = 0.00001$. $H = \sqrt{70/3}$.

t_e is the transition time...

Walking dynamics may lead to inflationary dynamics...

Chiral Transition in Janik's Cooling Geometry

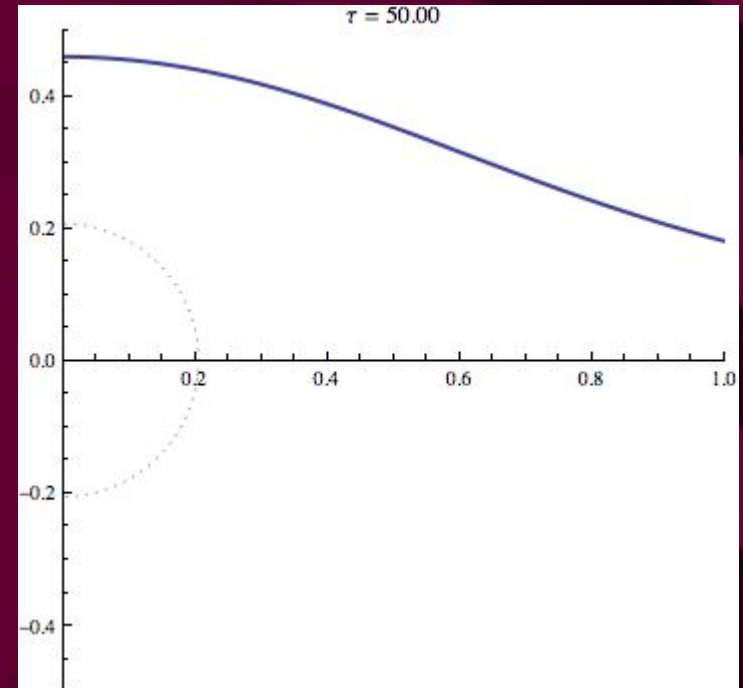
The black hole grows/shrinks changing the effective potential... With Ingo Kirsch, Tigran Kalaydzhyan (DESY)

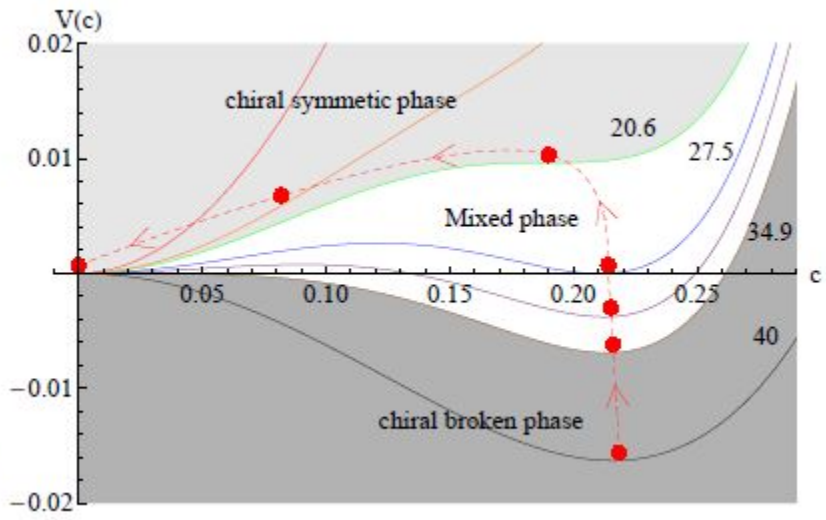
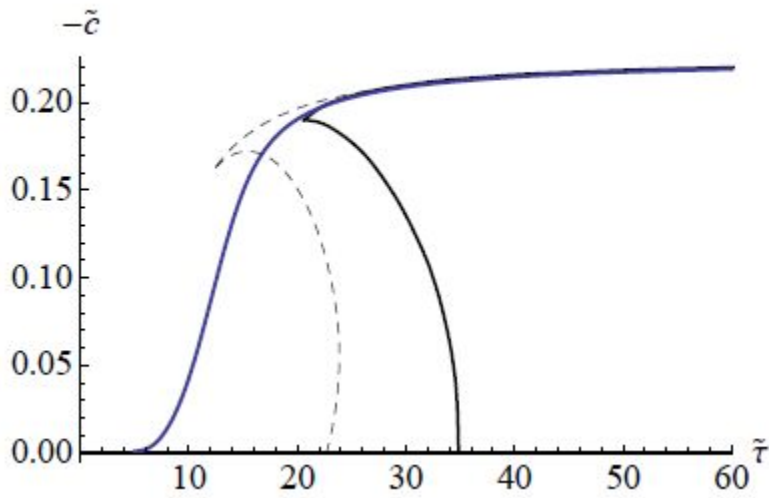
$$ds^2 = \frac{r^2}{R^2} (-e^{a(\tau,r)} d\tau^2 + e^{b(\tau,r)} \tau^2 dy^2 + e^{c(\tau,r)} dx_\perp^2) + \frac{R^2}{r^2} (d\rho^2 + \rho^2 d\Omega_3^2 + dL^2 + L^2 d\phi^2)$$

$$a(\tau, z) = \ln \left(\frac{(1 - v^4/3)^2}{1 + v^4/3} \right) + 2\eta_0 \frac{(9 + v^4)v^4}{9 - v^8} \left[\frac{1}{(\varepsilon_0^{3/8} \tau)^{2/3}} \right] + \mathcal{O} \left[\frac{1}{\tau^{4/3}} \right],$$

$$b(\tau, z) = \ln(1 + v^4/3) + \left(-2\eta_0 \frac{v^4}{3 + v^4} + 2\eta_0 \ln \frac{3 - v^4}{3 + v^4} \right) \left[\frac{1}{(\varepsilon_0^{3/8} \tau)^{2/3}} \right] + \mathcal{O} \left[\frac{1}{\tau^{4/3}} \right],$$

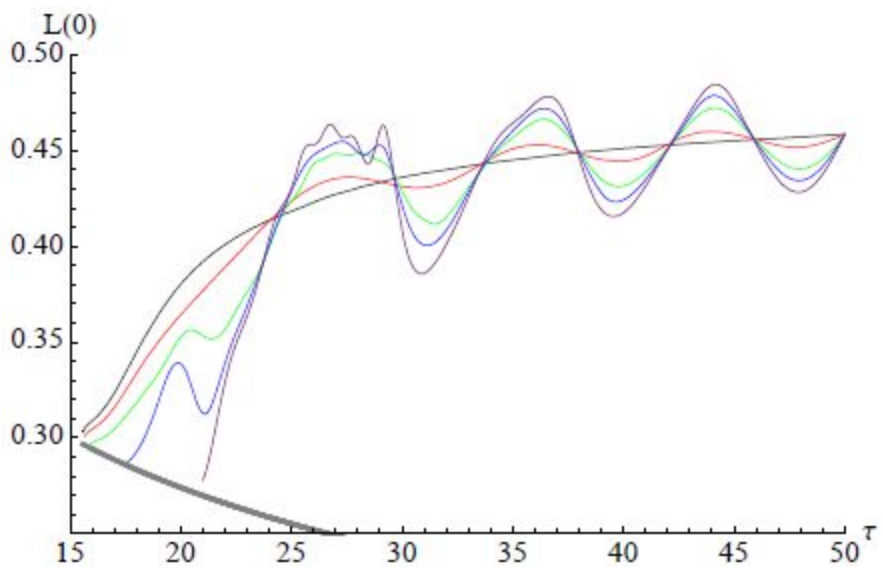
$$c(\tau, z) = \ln(1 + v^4/3) + \left(-2\eta_0 \frac{v^4}{3 + v^4} - \eta_0 \ln \frac{3 - v^4}{3 + v^4} \right) \left[\frac{1}{(\varepsilon_0^{3/8} \tau)^{2/3}} \right] + \mathcal{O} \left[\frac{1}{\tau^{4/3}} \right],$$





Equilibrium vs PDE solutions...

Bubble formation...



Conclusion

Strongly coupled phase diagrams with rich structure

arXiv:1002.1885 [hep-th], arXiv:1003.2694 [hep-th], arXiv:1103.5627 [hep-th]

Non mean-field quantum critical points... arXiv:1008.1889 [hep-th]

Out of equilibrium dynamics

Walking dynamics & inflation linked? arXiv:1009.5678 [hep-th]

Towards strongly coupled quenches, disordered chiral condensates etc... arXiv:1011.2519 [hep-th]

Wilsonian derivation of effective potential, inflaton phenomenology, horizon formation and thermalization...